

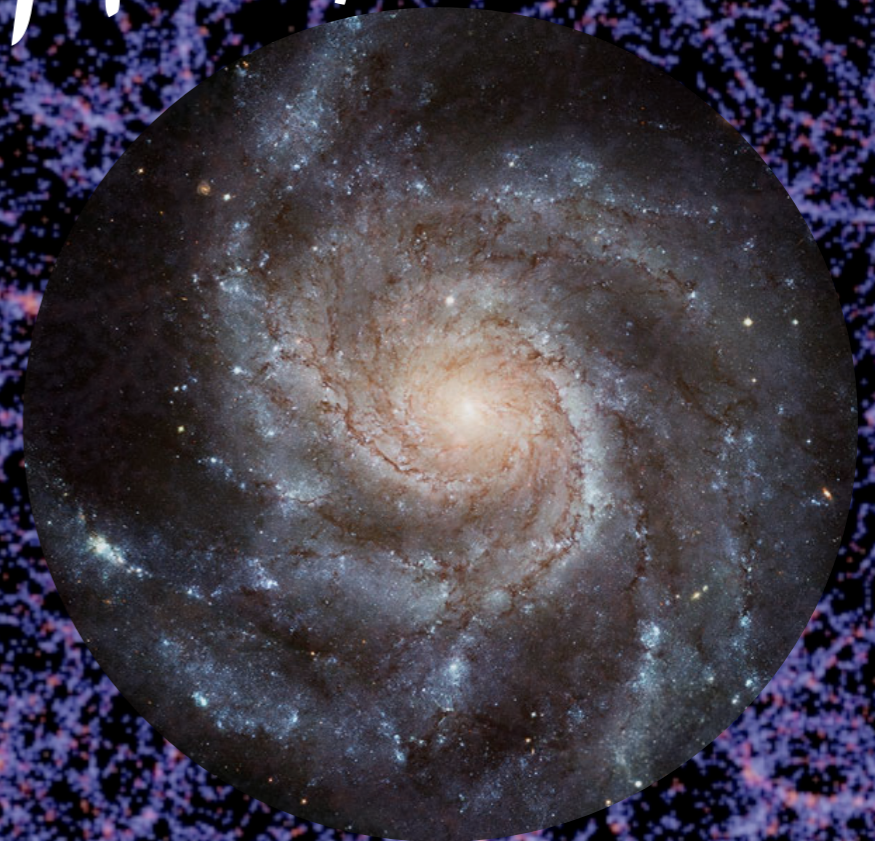
A visualization of the cosmic web, showing a complex network of blue and orange filaments and nodes against a dark background. The filaments represent the large-scale structure of the universe, with nodes indicating regions of high density and galaxy clusters.

Structure Formation

PROF. FRANK VAN DEN BOSCH
DEPT. OF PHYSICS & ASTRONOMY

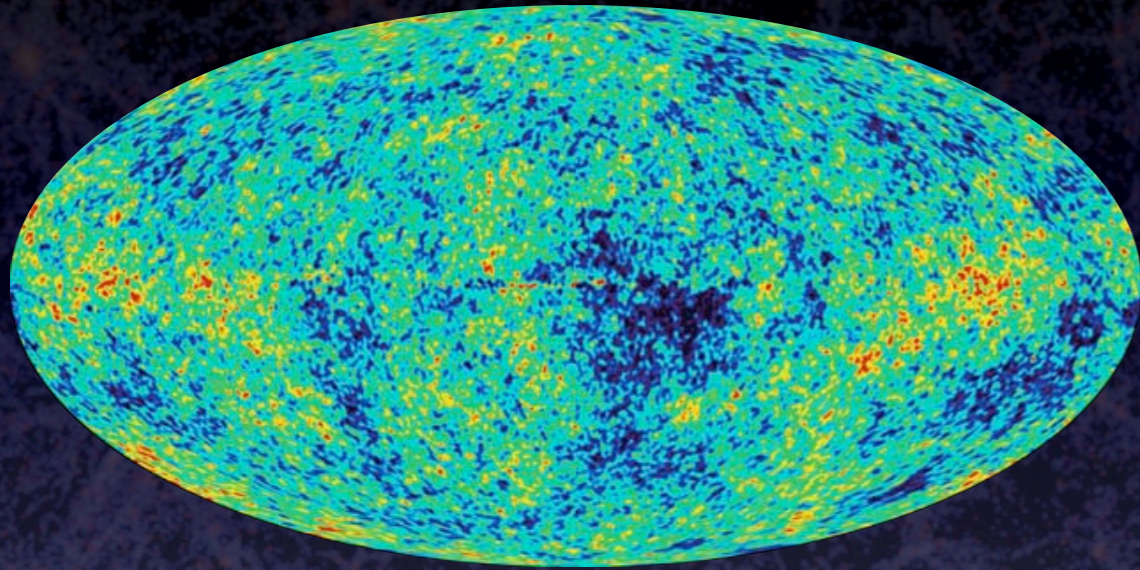
Structure Formation

From Inflation to Milky Way

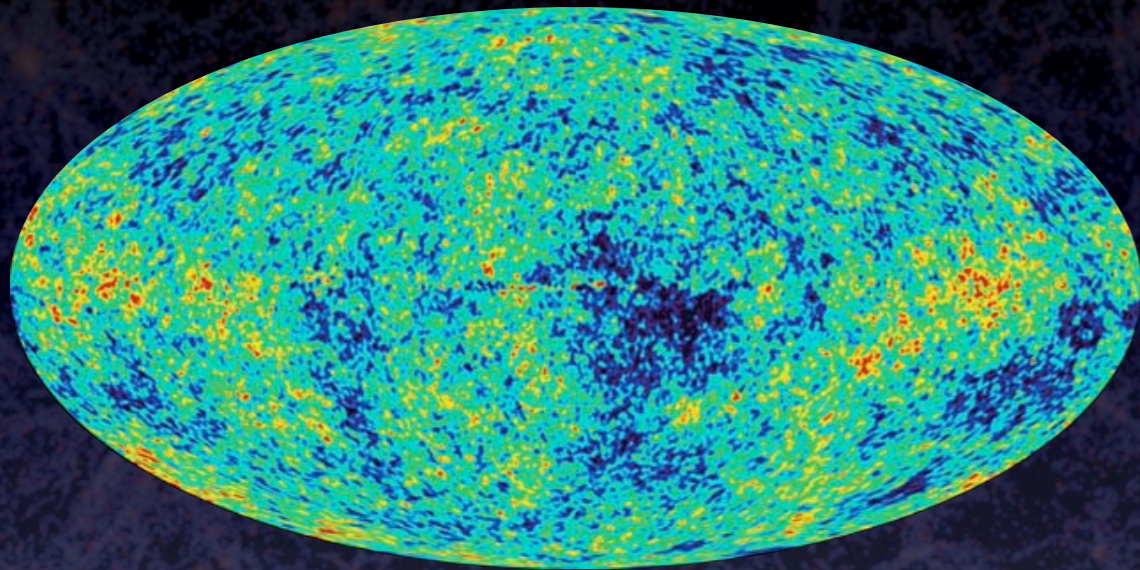


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Recap of last week....

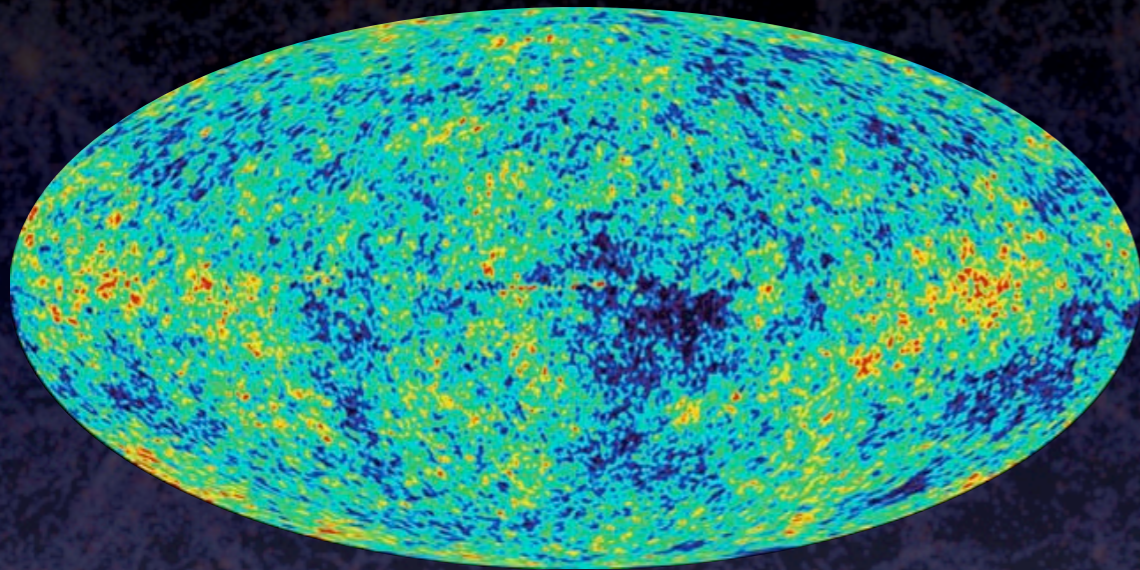


Recap of last week....



- ★ What is dark matter?
- ★ What is dark energy?
- ★ What is origin of perturbations?
- ★ Why is Universe flat?

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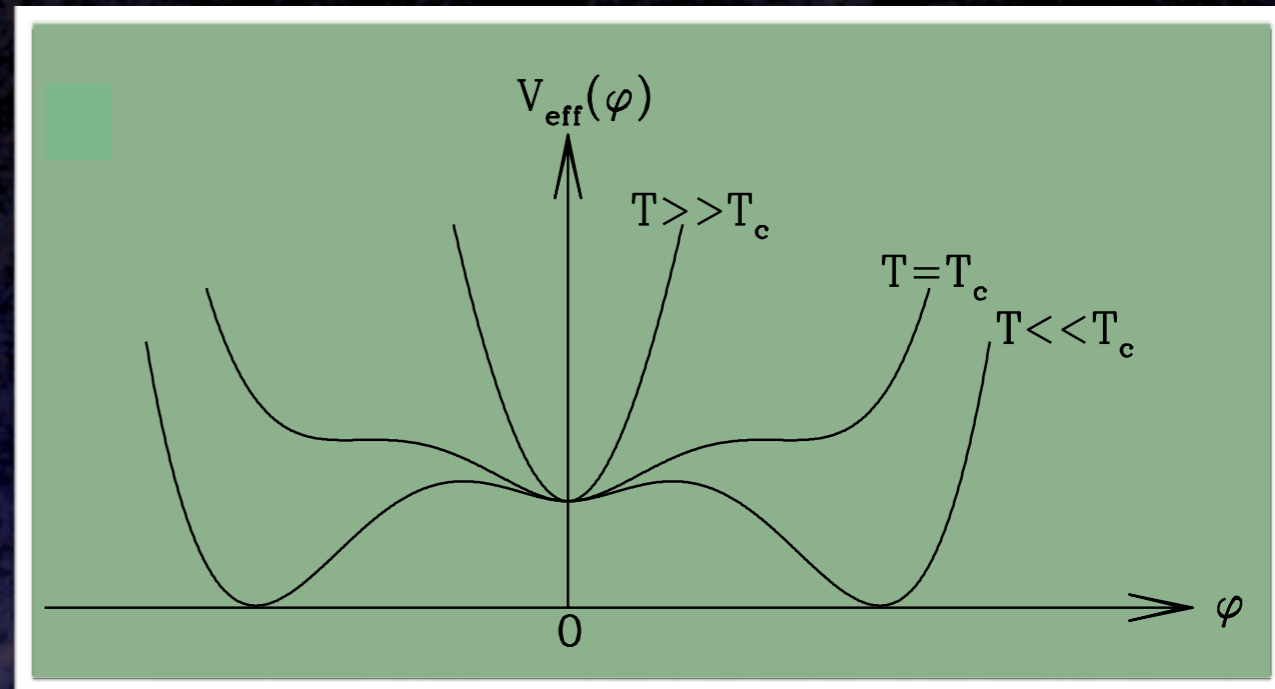


The Inflationary Universe

Inflation

Postulate

Shortly after Big Bang ($\sim 10^{-35}$ s), energy density of Universe is dominated by false vacuum state of a scalar field (the inflaton)



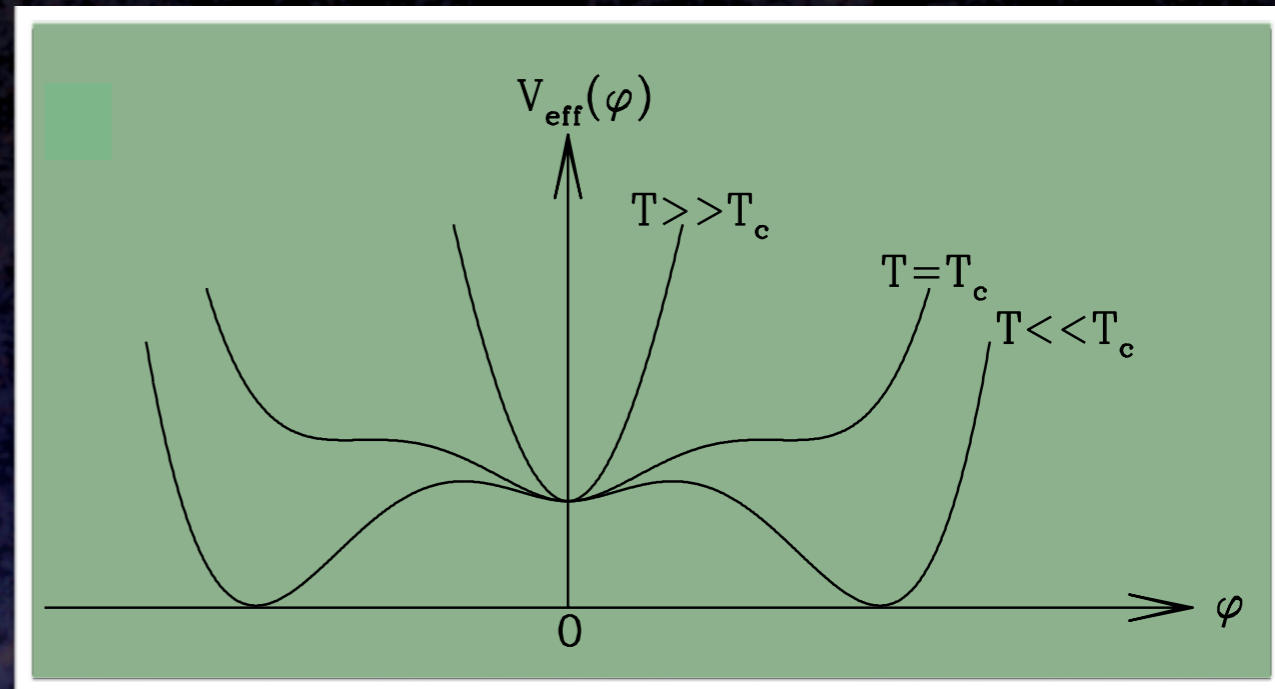
Inflaton acts like dark energy and causes exponential expansion
Inflation stops when inflaton `decays' to true vacuum state.
Energy of inflaton converted to particles & photons.

Inflation is envisioned to last for at least ~ 60 e-foldings, during which size of Universe increased by a factor $\sim 10^{26}$

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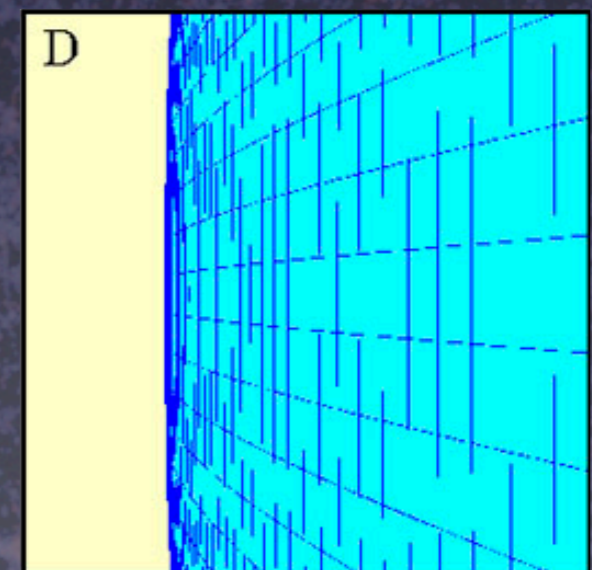
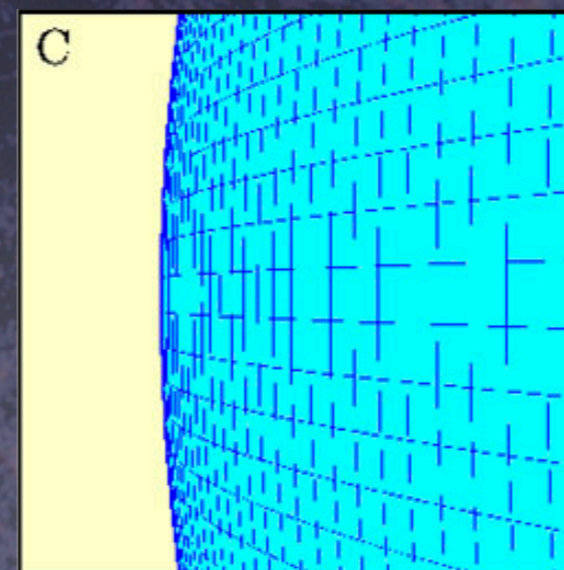
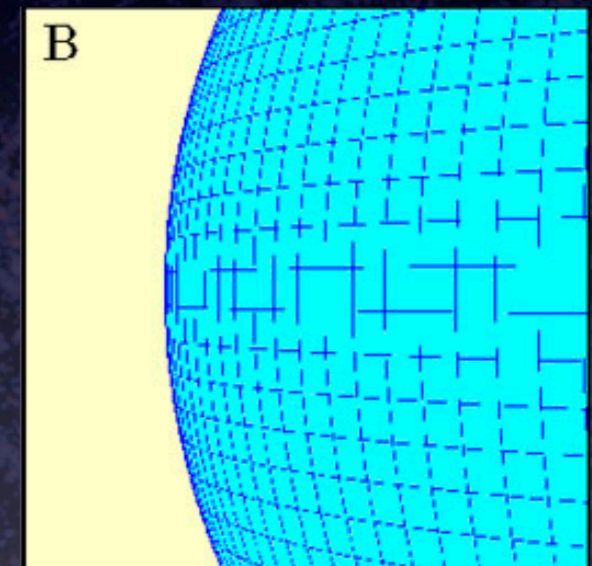
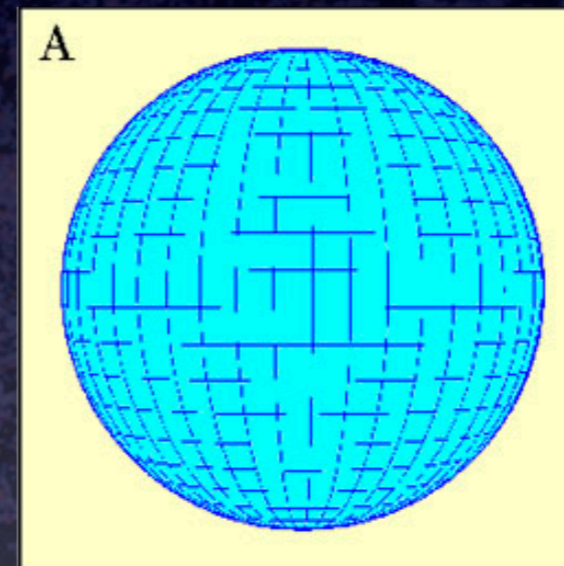
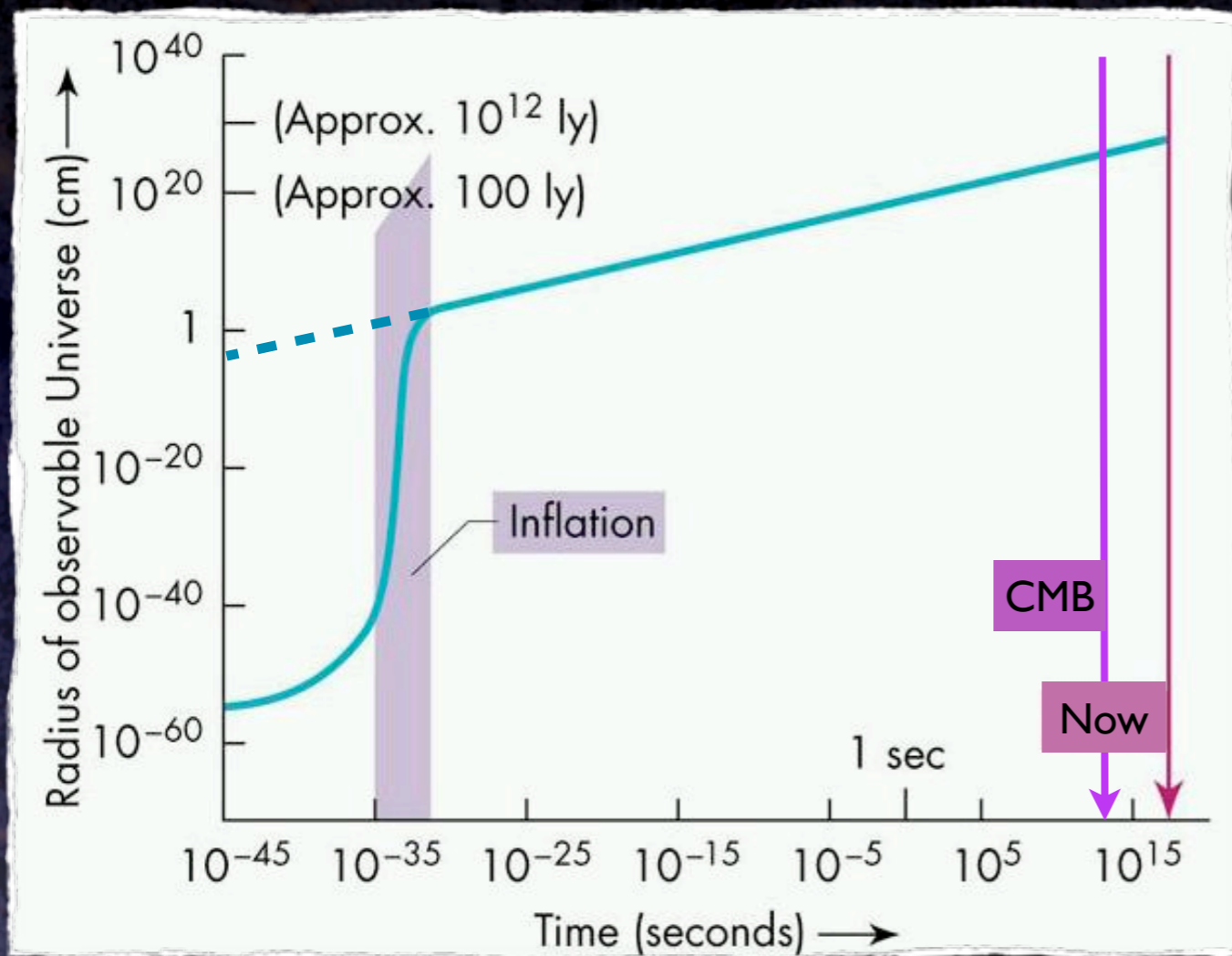


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For comparison, since CMB Universe has expanded by factor 10^3

Stretching Space



Because of huge expansion, any pre-existing curvature is inflated away --> post-inflationary Universe is flat

Quantum physics dictates that on very small scales, energy density associated with inflaton fluctuates...

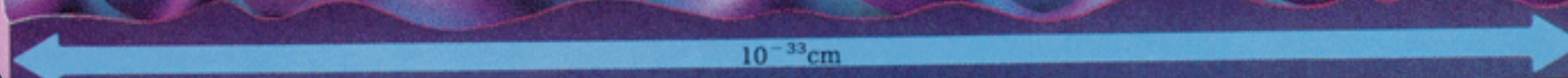
10^{-12} cm



10^{-30} cm



10^{-33} cm



Because of exponential expansion, these quantum fluctuations are inflated to fluctuations in energy density of Universe

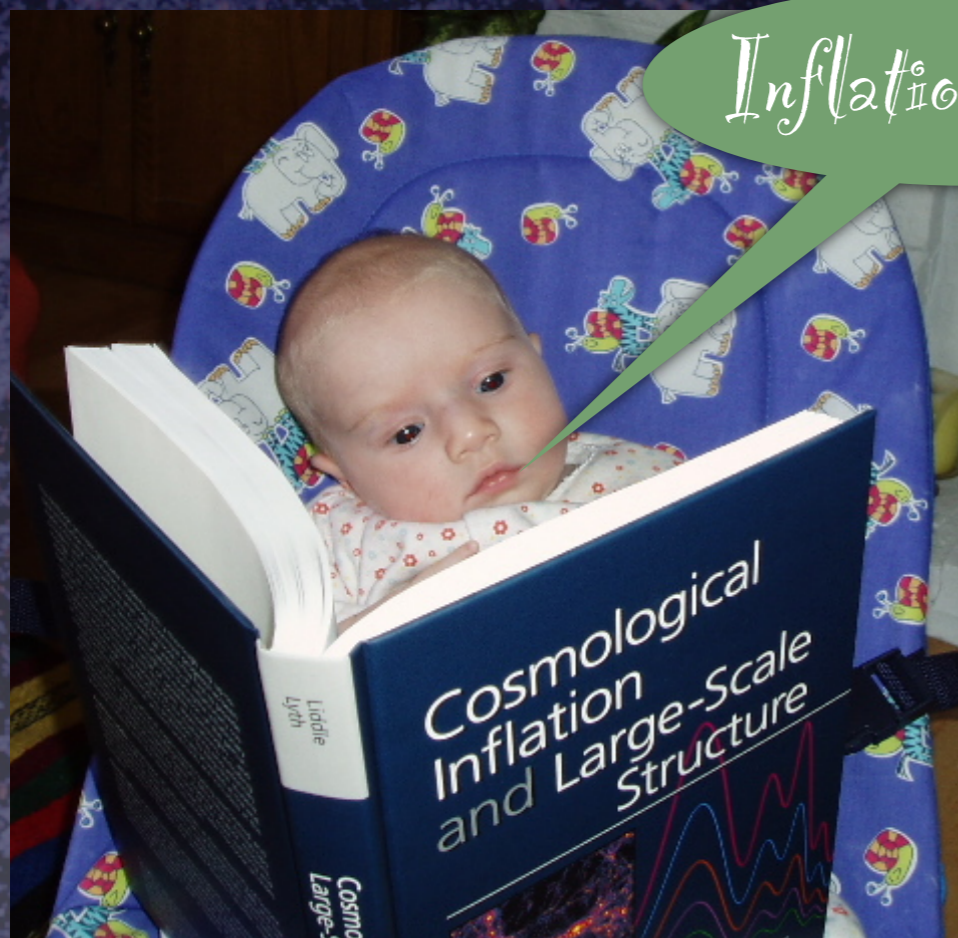
The Inflationary Universe

During inflation, a patch the size of a human hair (width) is inflated to patch larger than our Milky Way, in less than 10^{-33} s

Inflation solves the flatness problem in that it inflates away any pre-existing curvature

Inflation rocks

Because of quantum fluctuations, inflation also automatically predicts generation of density perturbations on wide range of scales



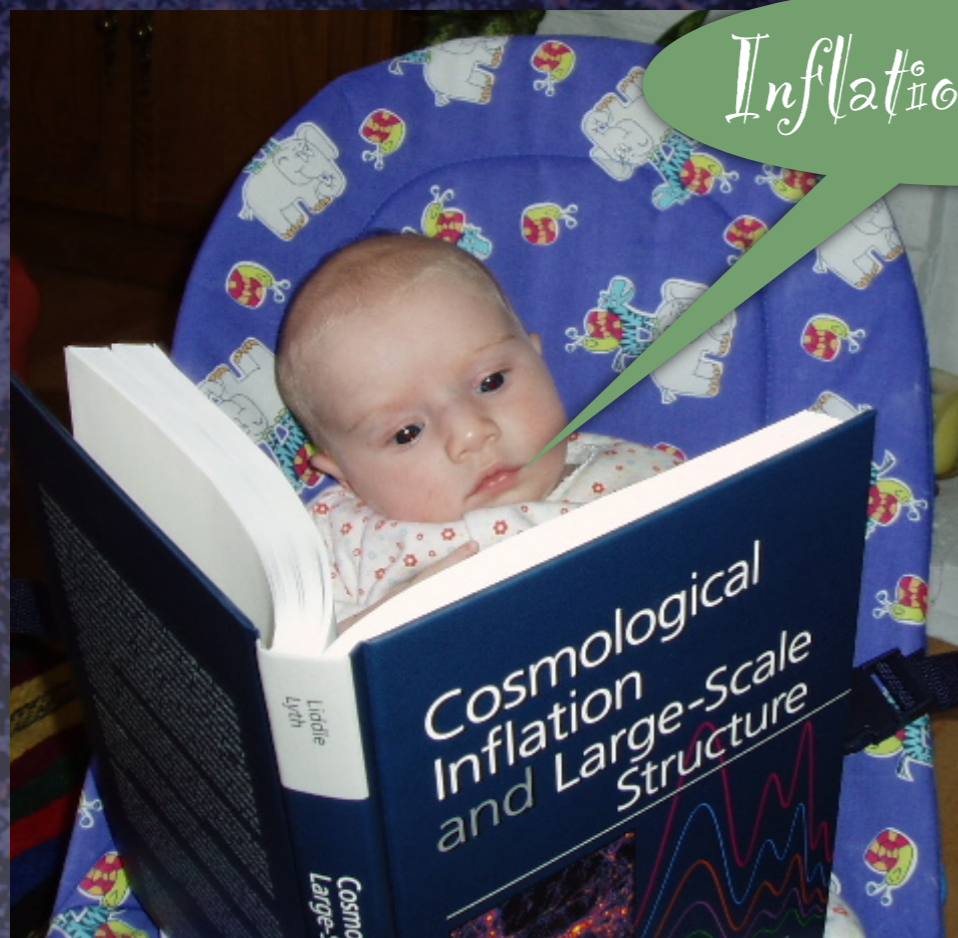
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The Structured Universe

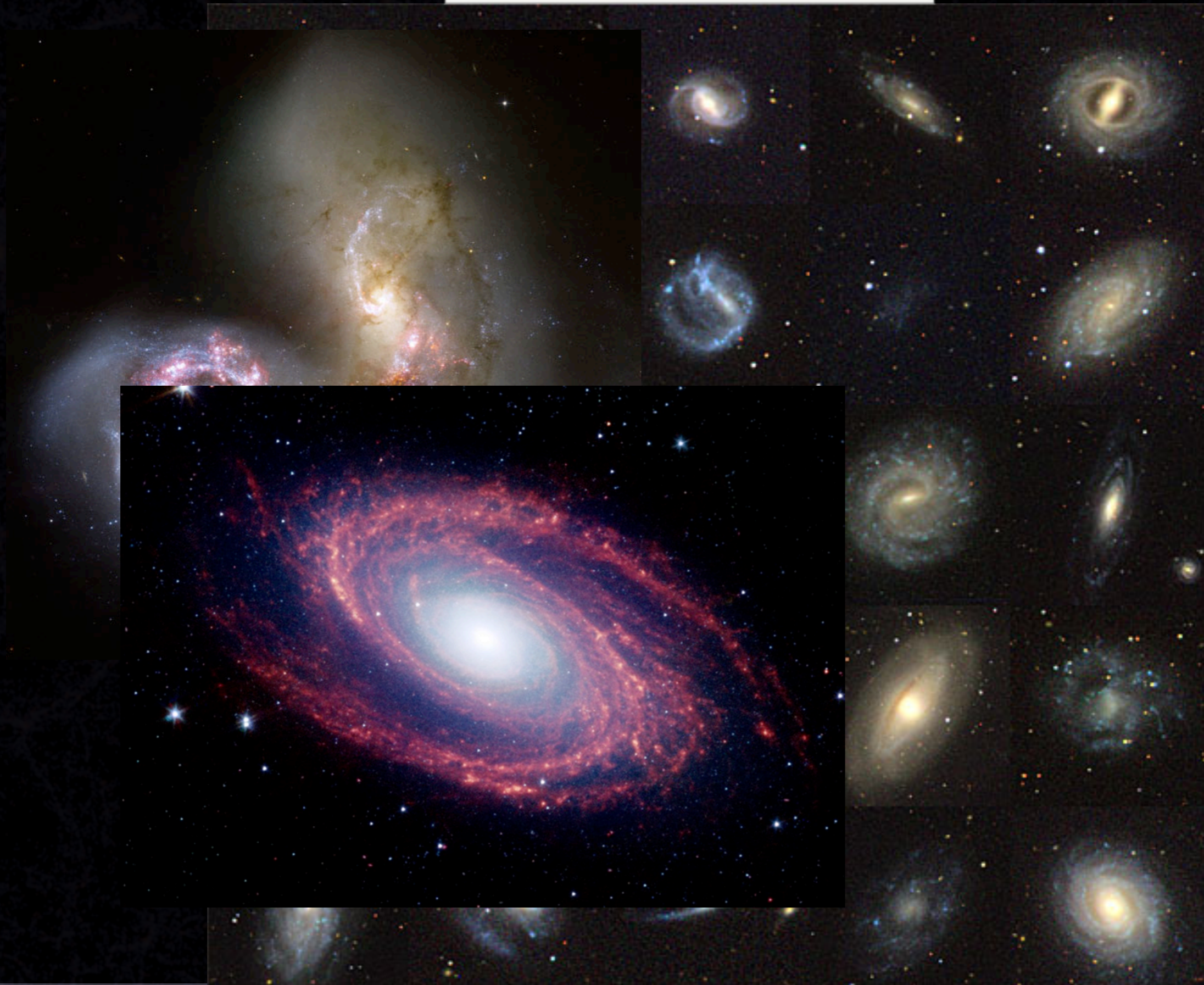
The Galaxy Zoo



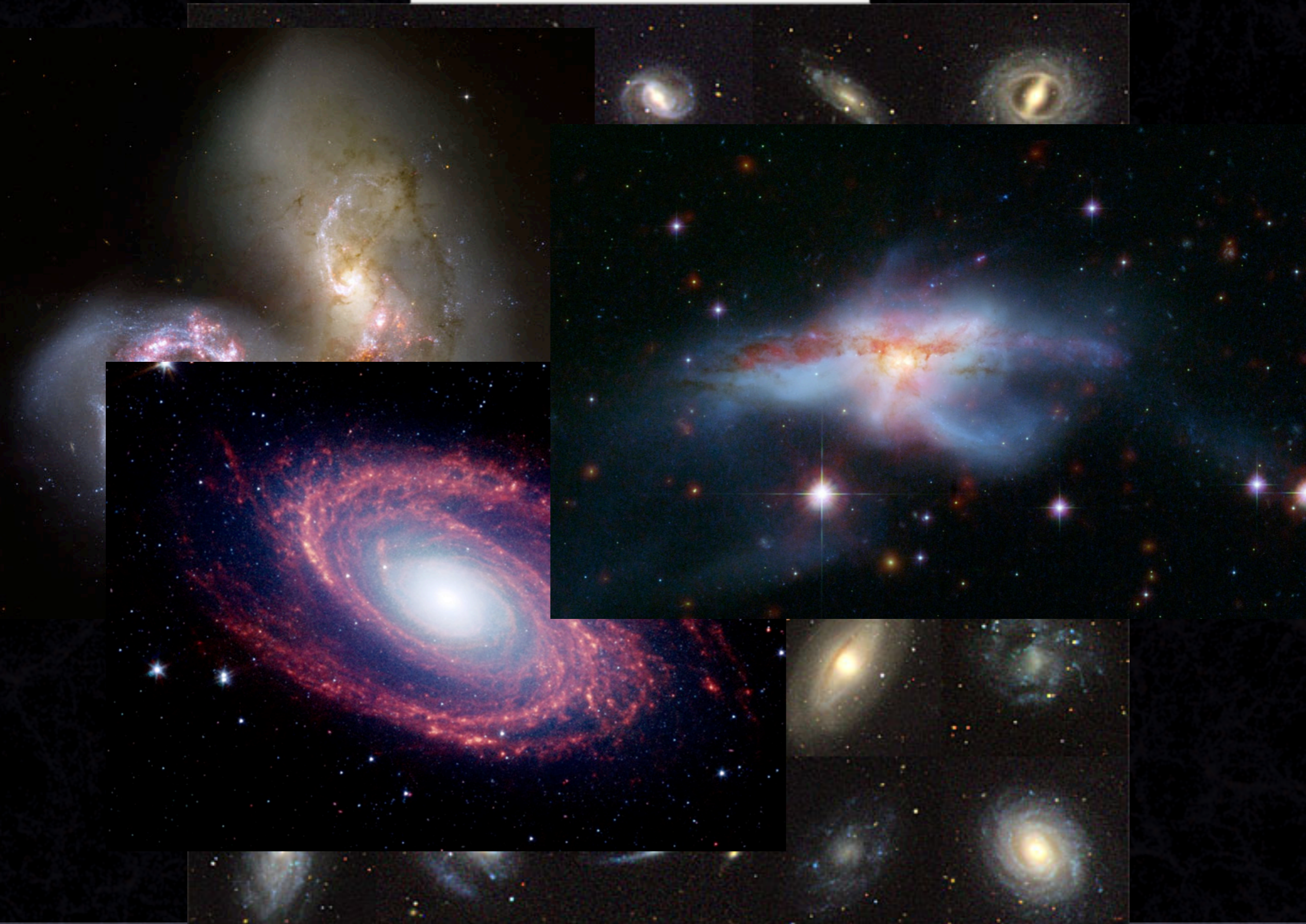
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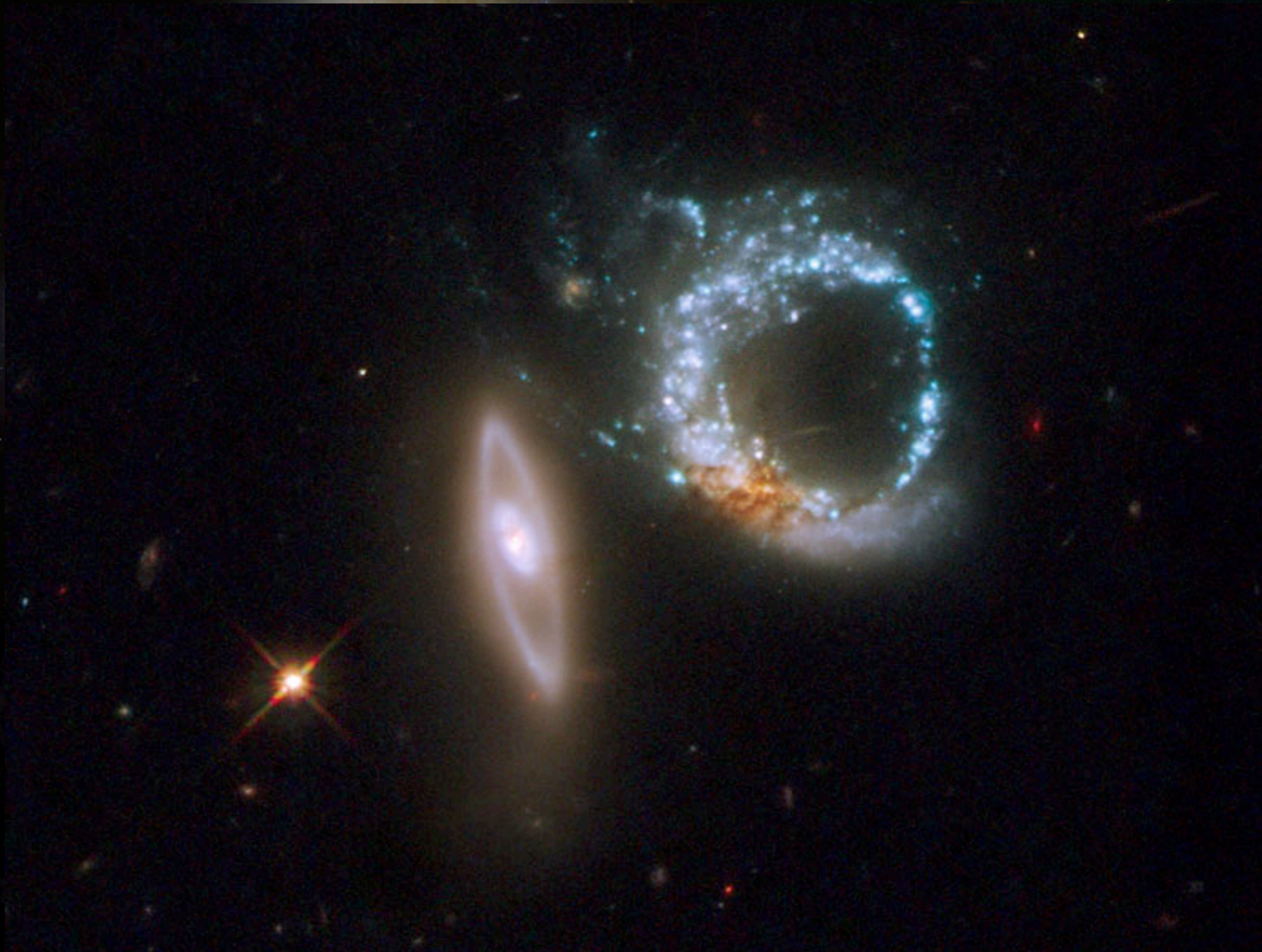
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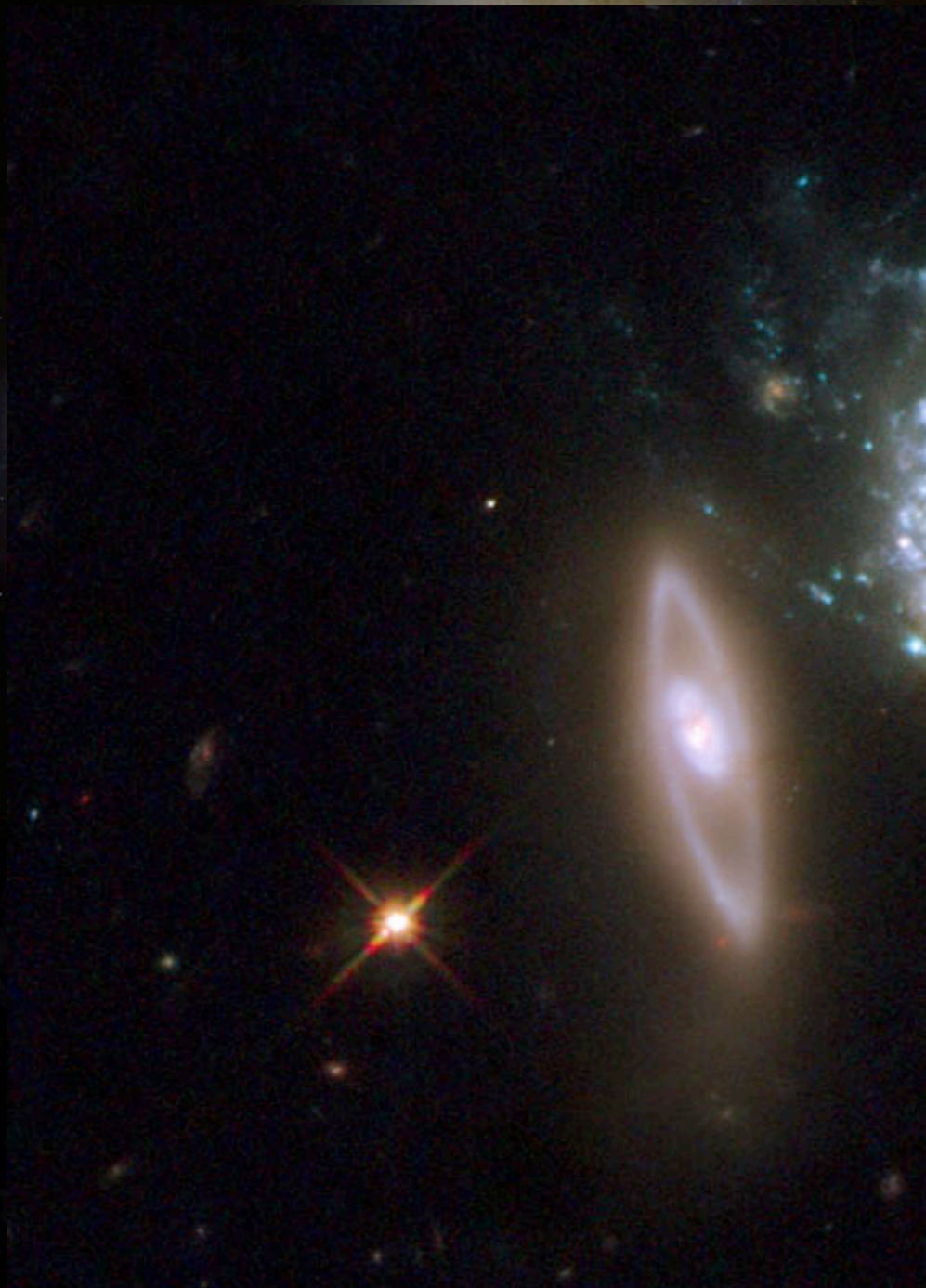
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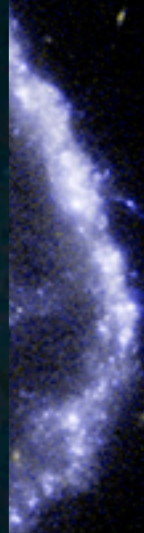
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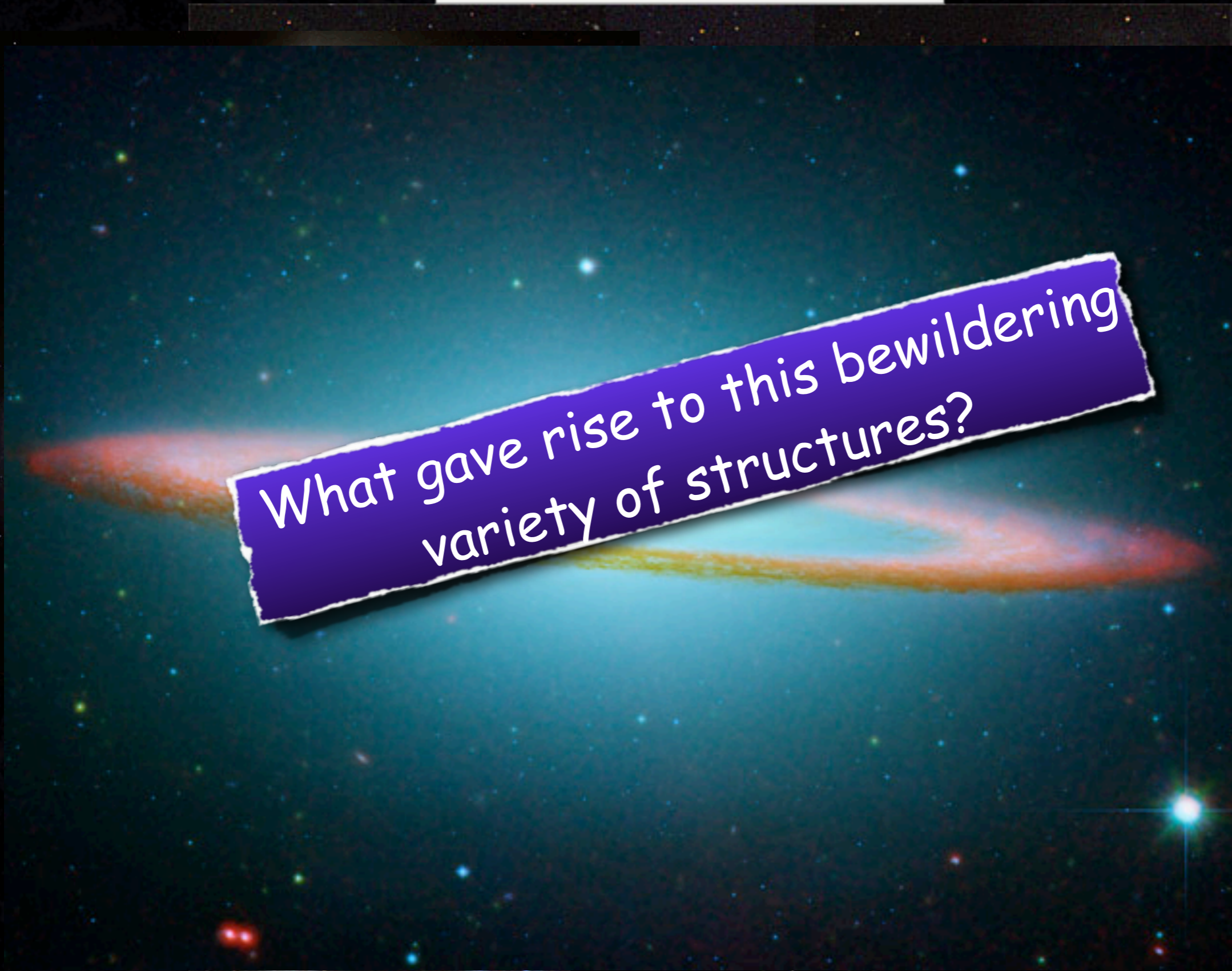


The Galaxy Zoo



The Galaxy Zoo

What gave rise to this bewildering variety of structures?

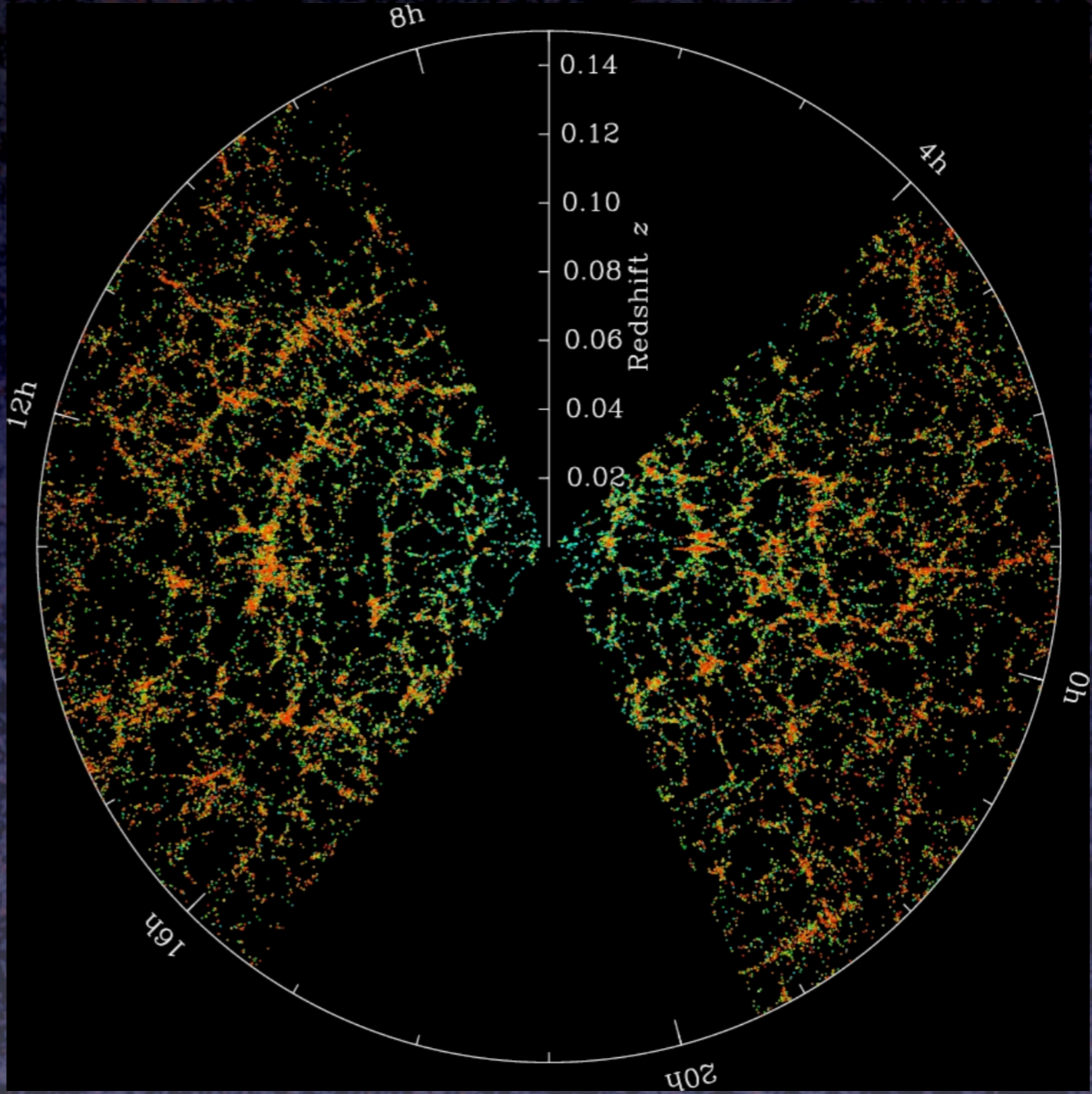


The Distribution of Galaxies

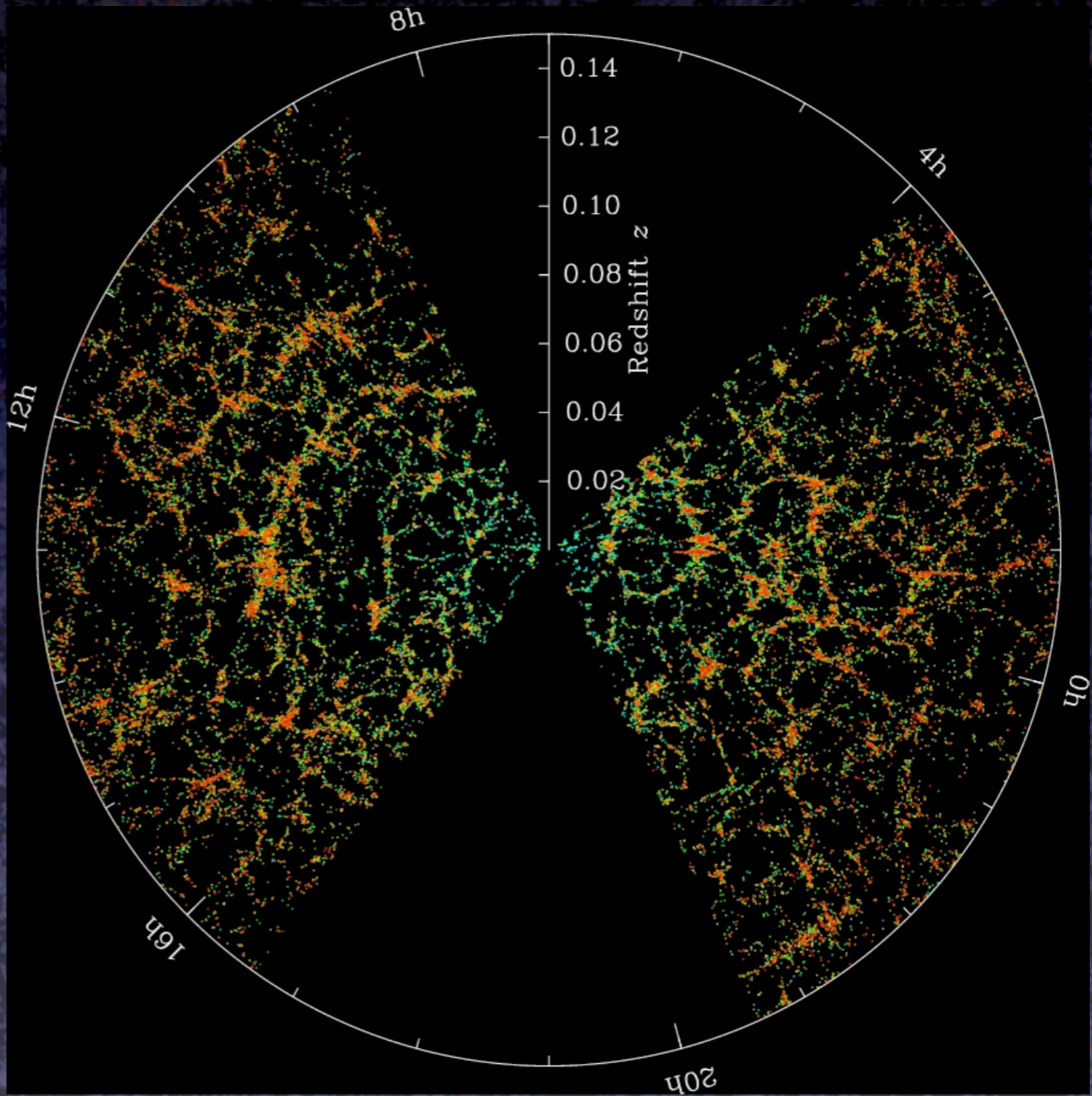


The Milky Way and more than 1.5 million galaxies from 2MASS, color coded according to redshift (blue = near, red = far)

Sloan Digital Sky Survey



Sloan Digital Sky Survey



Galaxy distribution is sponge-like; strong clustering



Massive clusters are the largest structures in the Universe,
and contain hundreds of (mainly elliptical) galaxies.

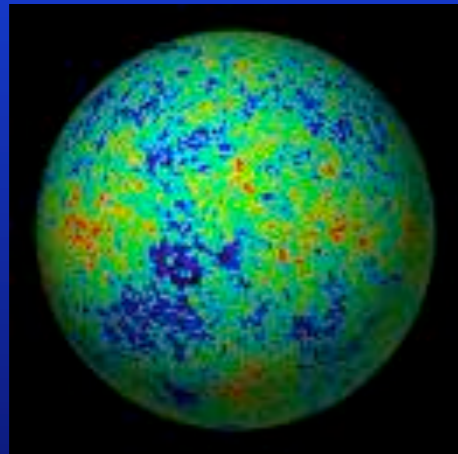
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A deep-field image of a galaxy cluster, showing hundreds of galaxies of various shapes and colors (yellow, white, blue, red) against a dark background. A large red circle is drawn around the central region of the cluster. A red arrow points from the center of the cluster to the edge of the red circle, with the text '~500,000 lightyears' written along the arrow.

~500,000 lightyears

The Puzzle of Structure Formation

At recombination:

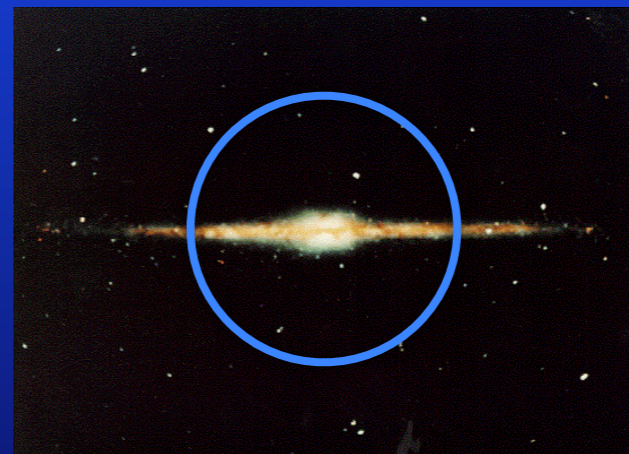


$$|\delta| < 10^{-5}$$

13 Gyr



Today, within Solar radius of MW



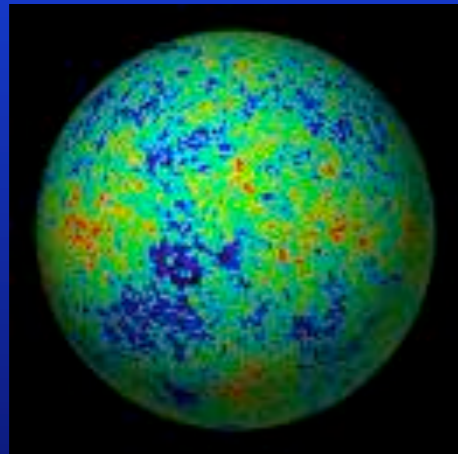
$$\delta \simeq 10^5$$

Dynamical time at Solar radius ~ 250 Myr

$$\delta(\mathbf{x}) = \frac{\rho(\mathbf{x}) - \bar{\rho}}{\bar{\rho}}$$

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At recombination:

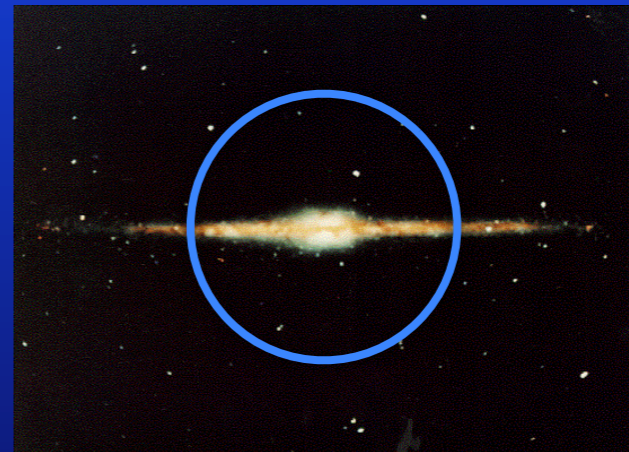


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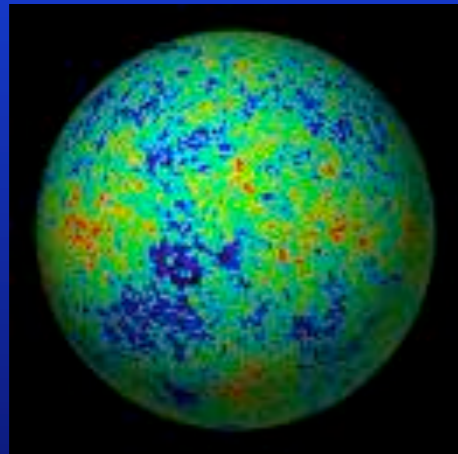
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Perturbations have grown
by more than 10 orders of magnitude
in less than 50 dynamical times

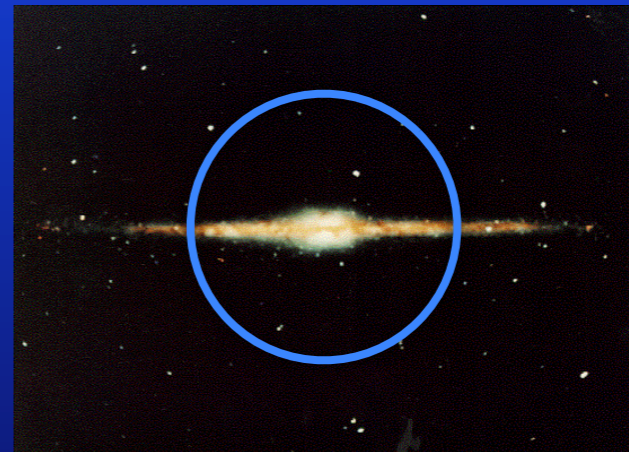
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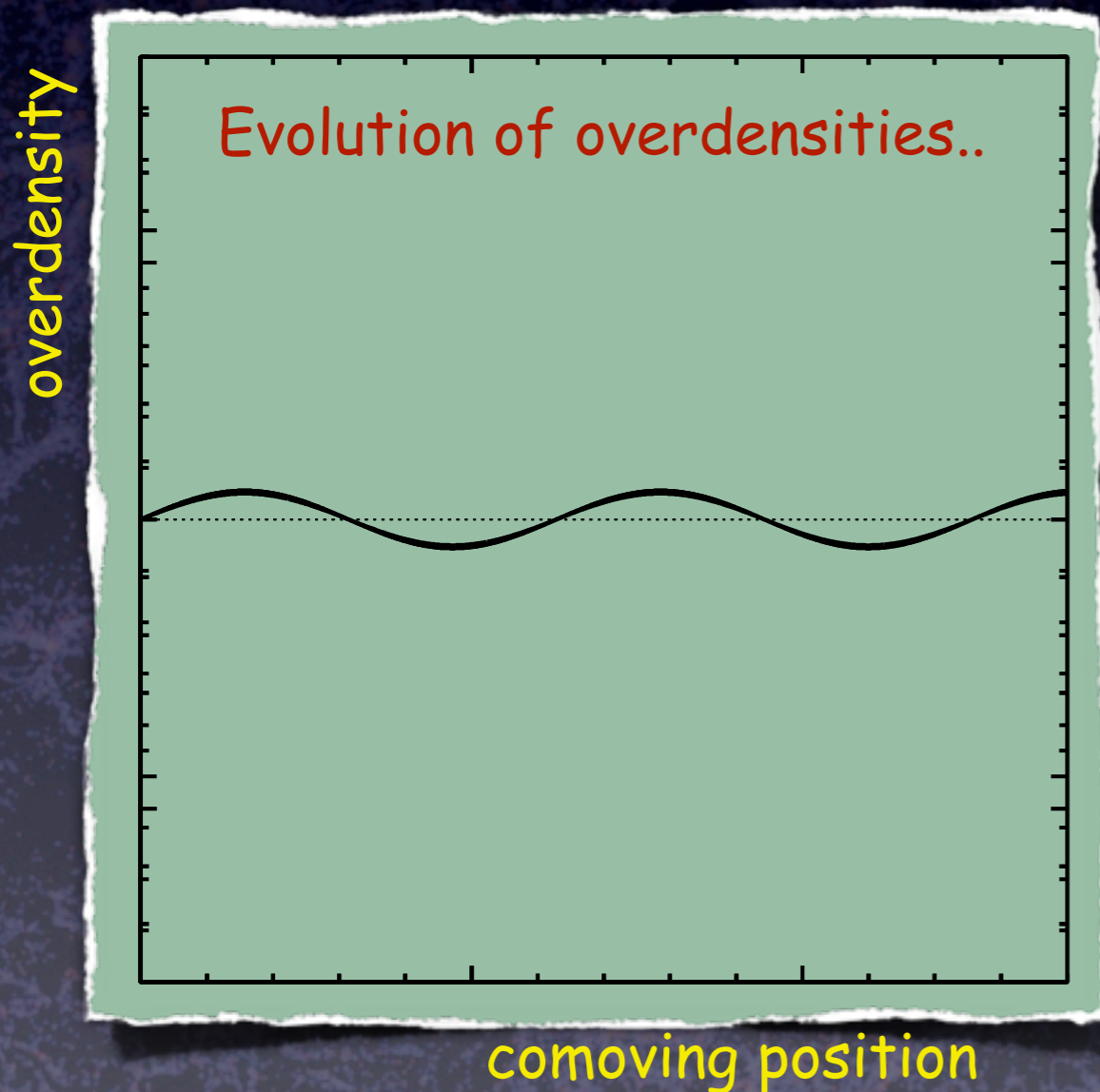
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Perturbations have grown
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HOW??

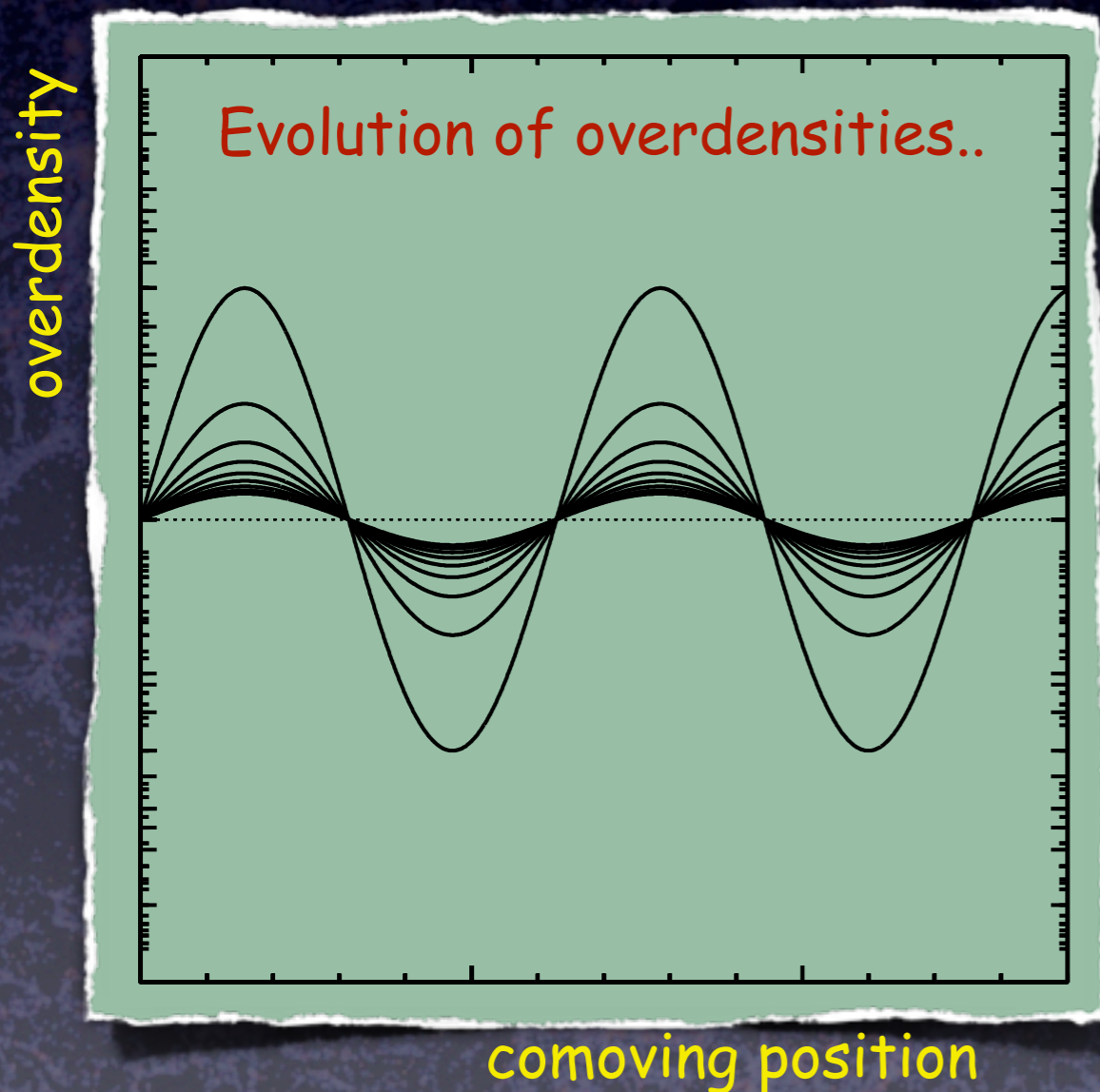
The Answer...

Gravitational Instability: slightly denser regions attract matter thus becoming even denser, etc.



The Answer...

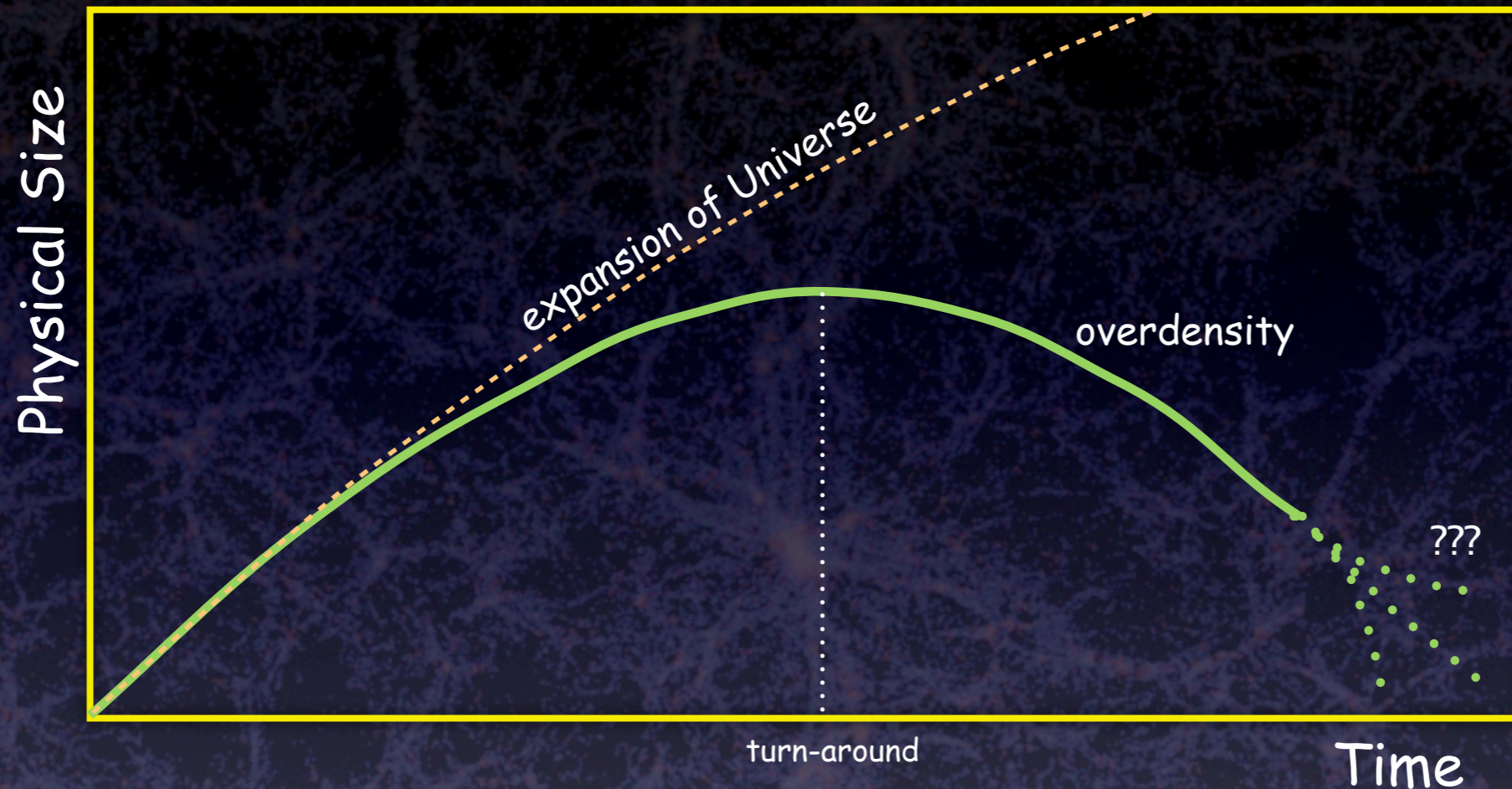
Gravitational Instability: slightly denser regions attract matter thus becoming even denser, etc.



This process continues until overdensities are of order unity.

At that point, overdensities 'turn around' (stop expanding) and start to collapse...

The Collapse of Perturbations



Evolution after turn-around depends on nature of matter

Dark Matter = collisionless → shell crossing

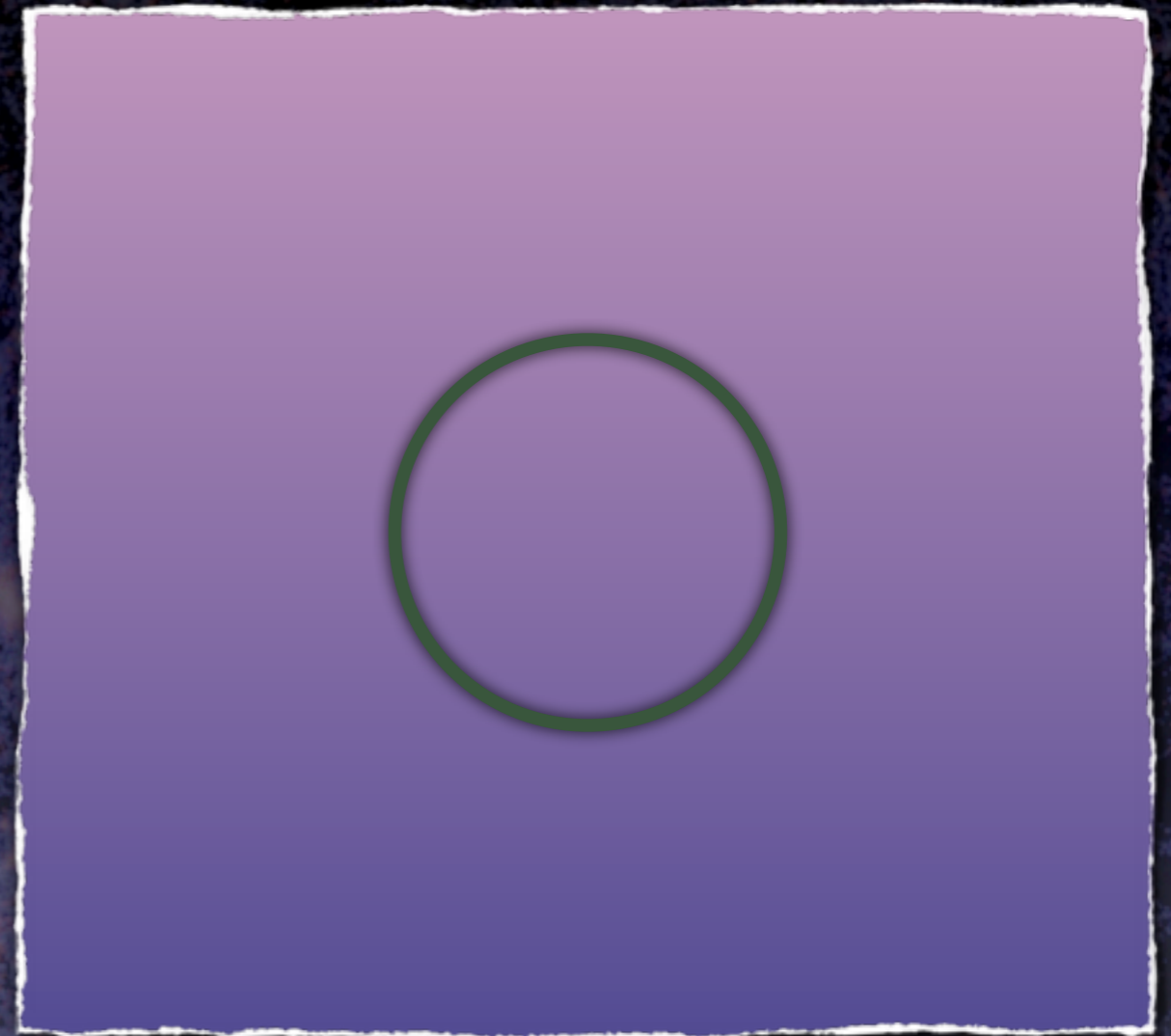
Baryonic Matter = collisional → shock heating

Evolution of shell of Cold Dark Matter



Onion Model

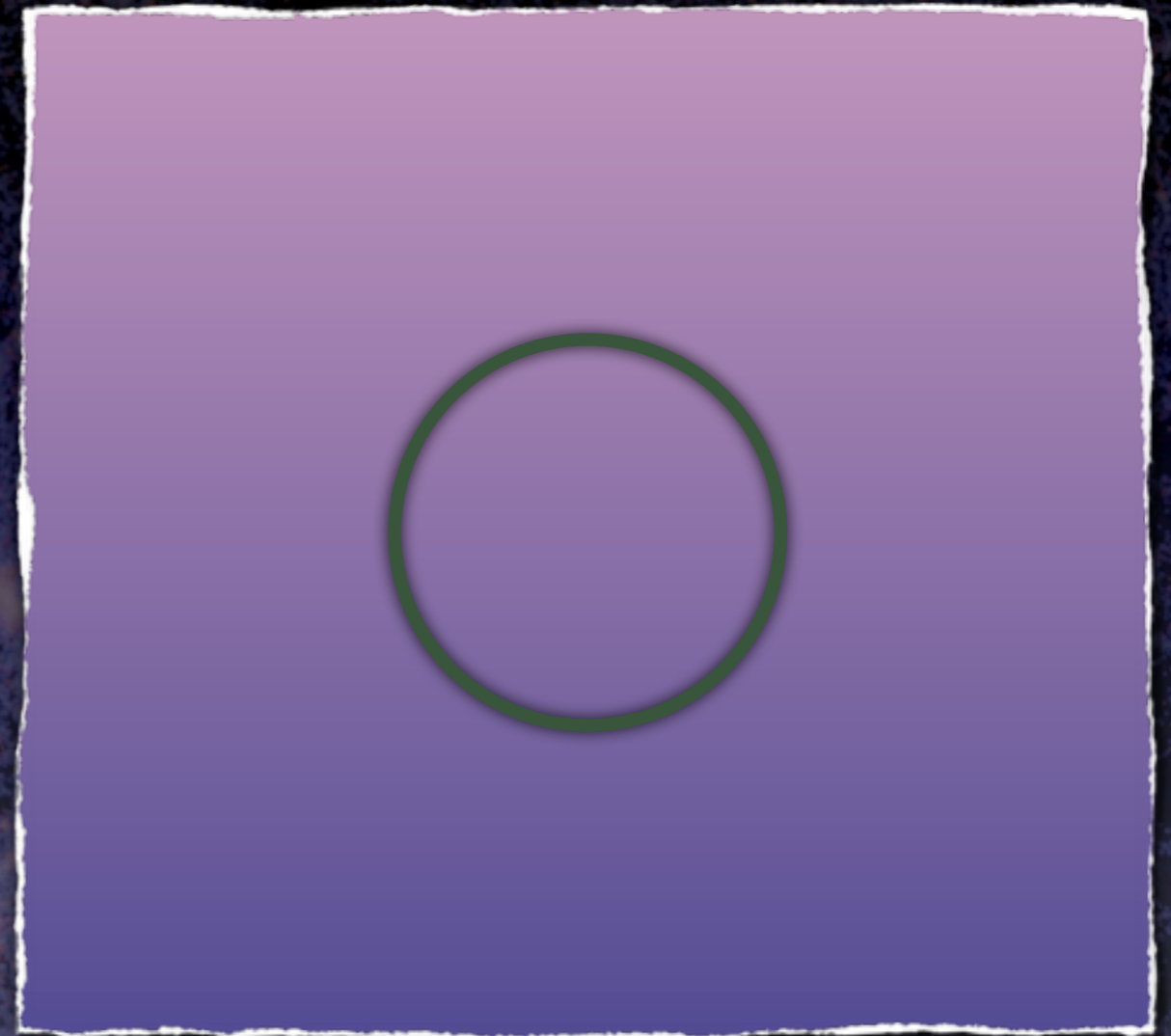
you can think of overdensity
as consisting of many
individual thin mass shells



Evolution of shell of Cold Dark Matter



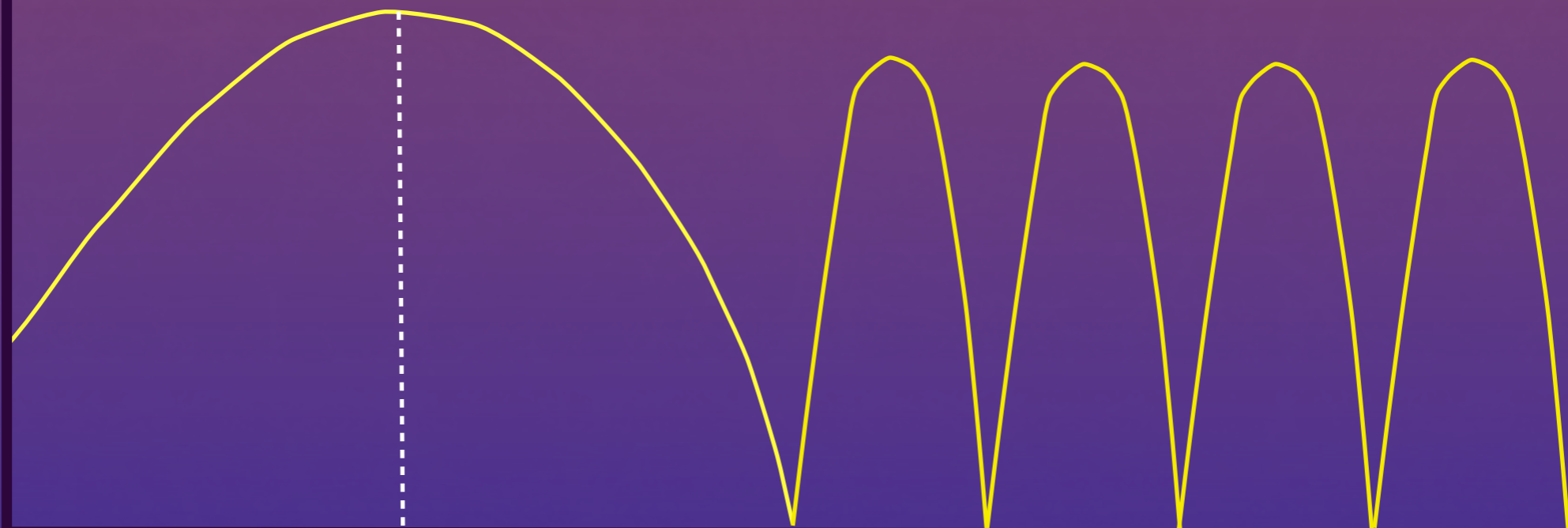
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Because dark matter has no pressure,
shell crosses itself and starts to oscillate

The Formation of a Dark Matter Halo

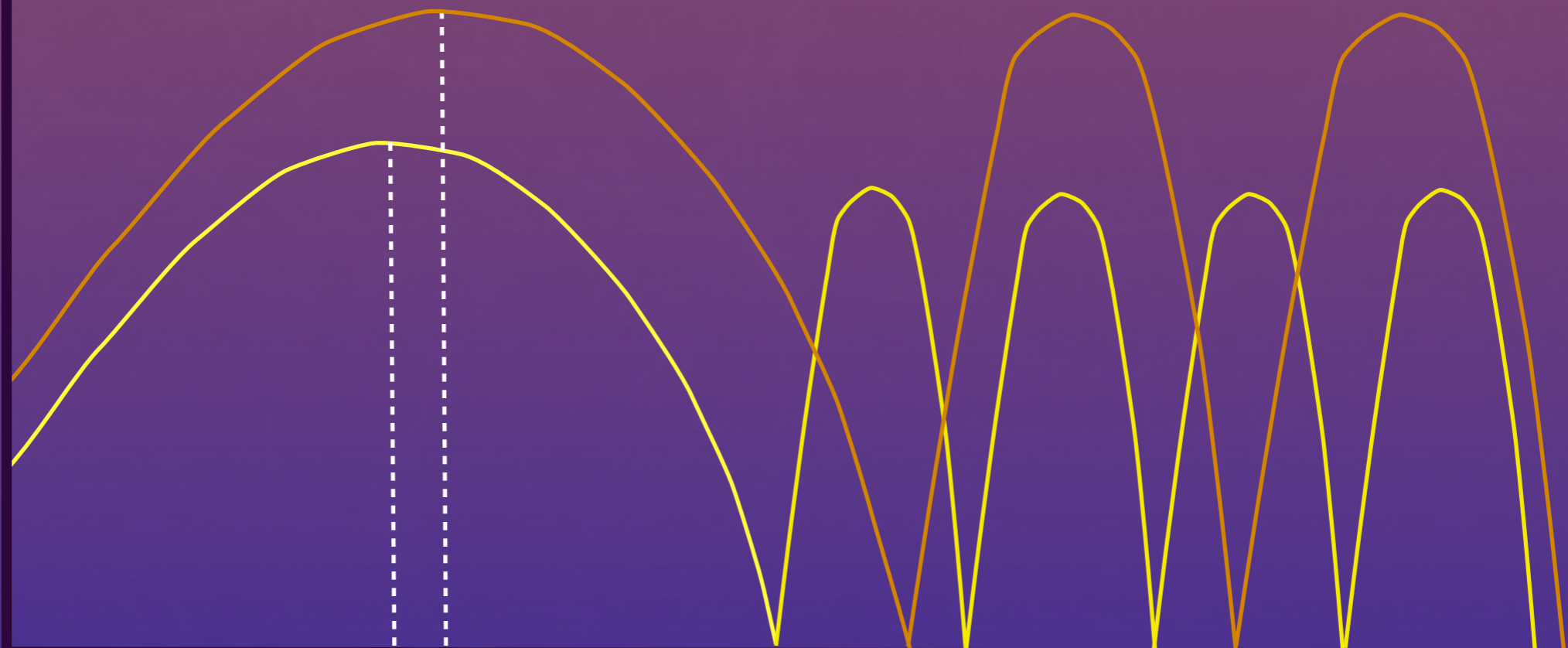
physical size



time

The Formation of a Dark Matter Halo

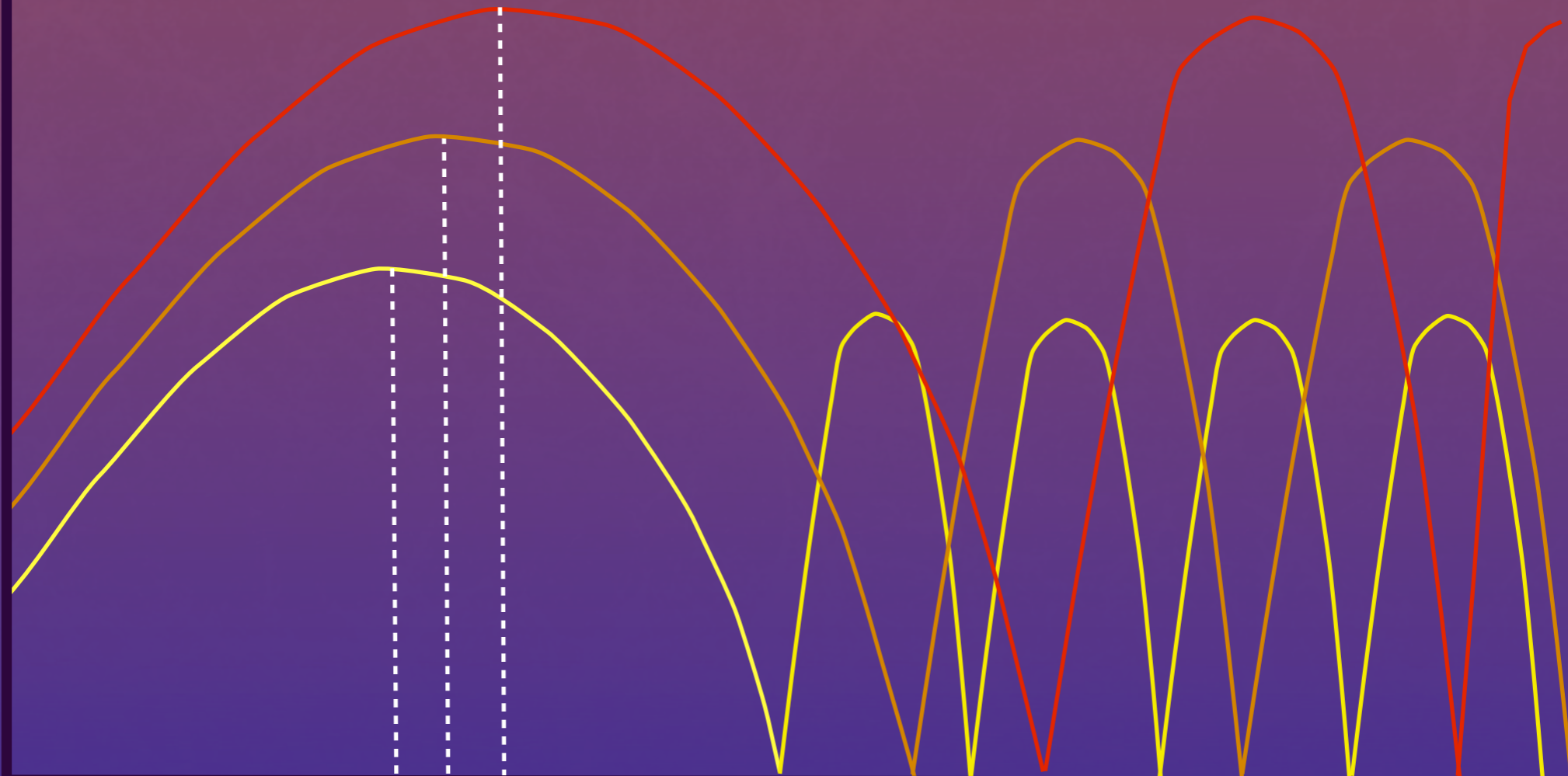
physical size



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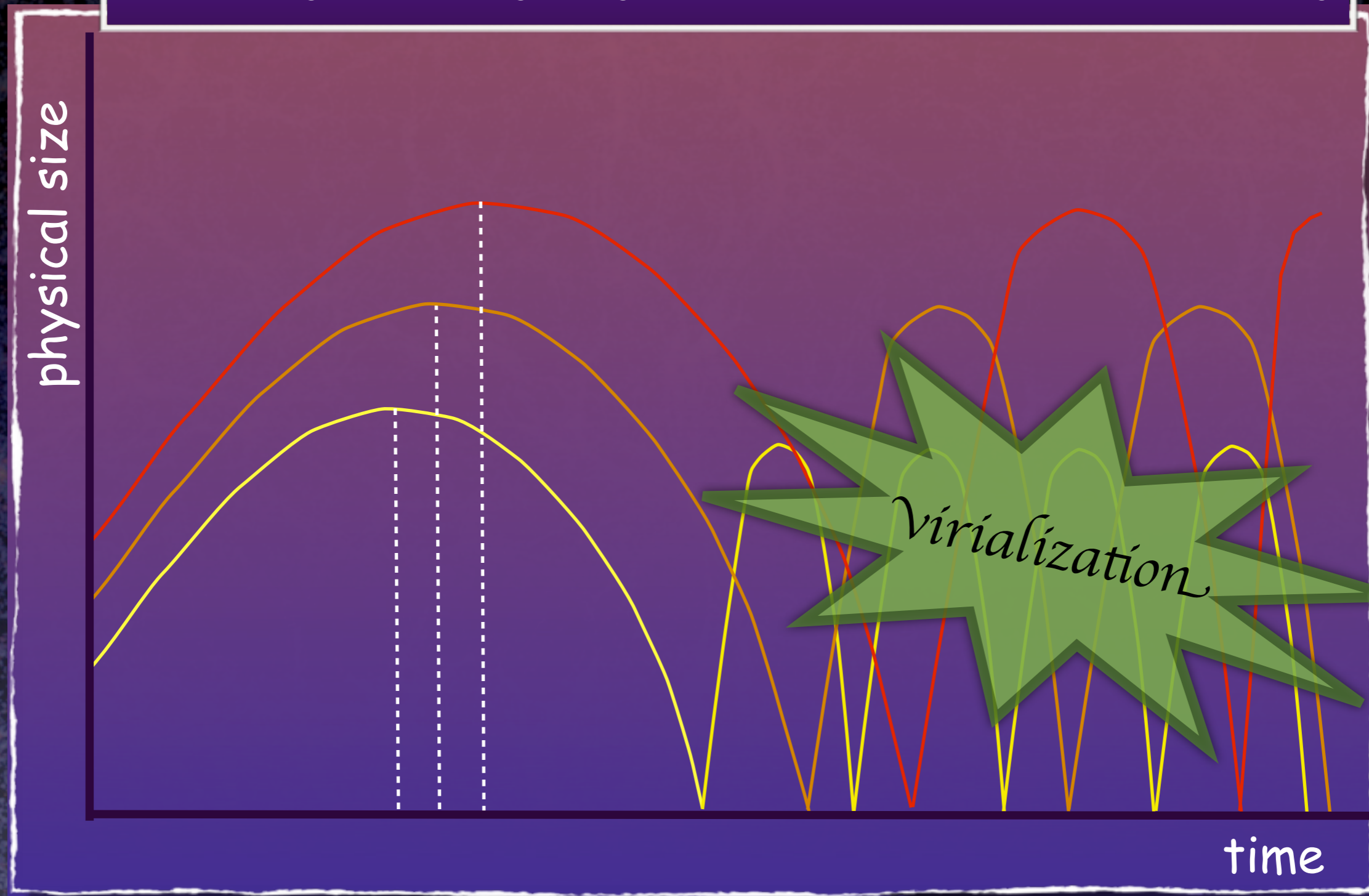
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The Formation of a Dark Matter Halo

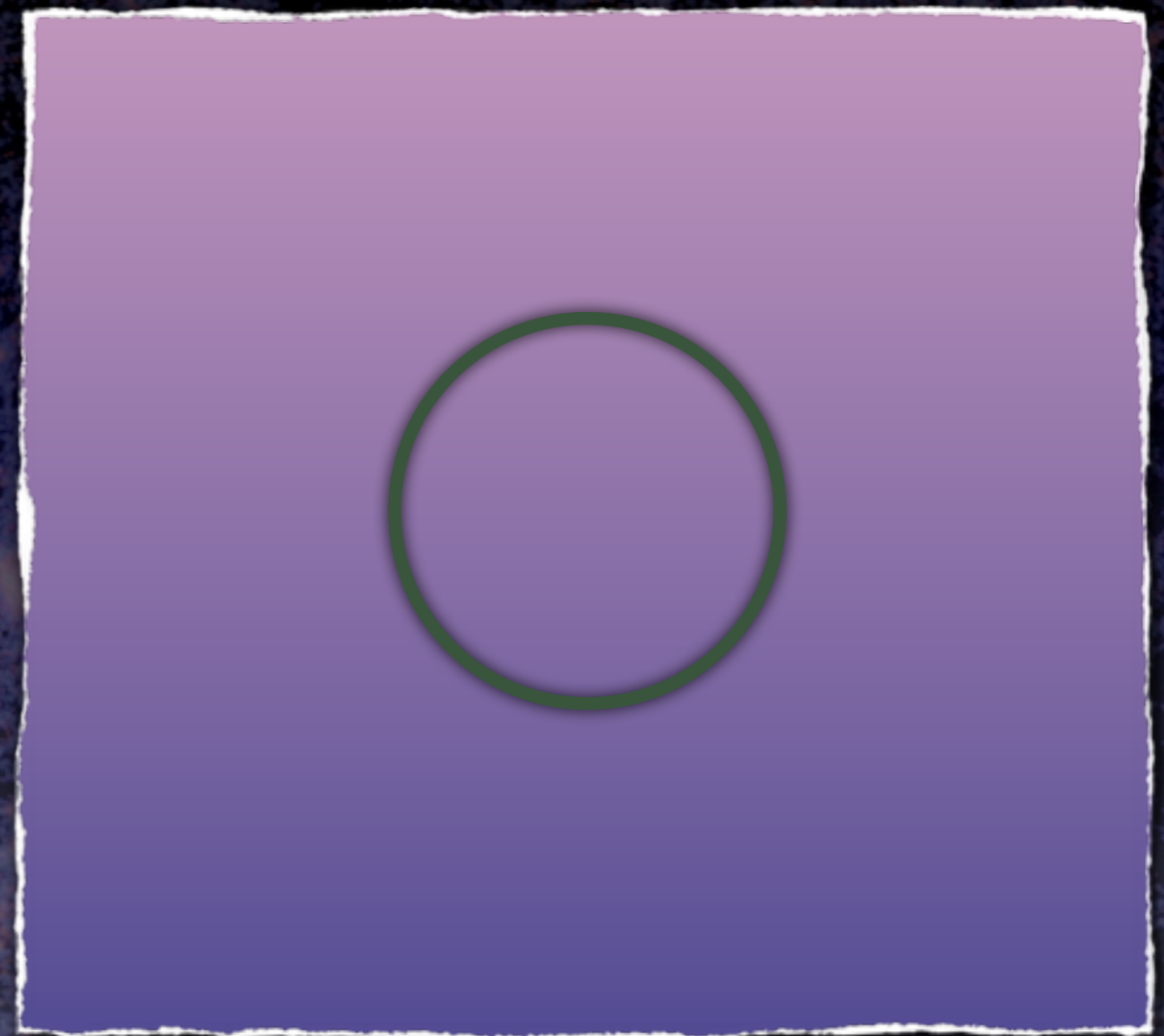


Individual oscillating shells interact gravitationally, exchanging energy (virializing), giving rise to a relaxed dark matter halo

Evolution of shell of Baryonic Matter



you can think of overdensity
as consisting of many
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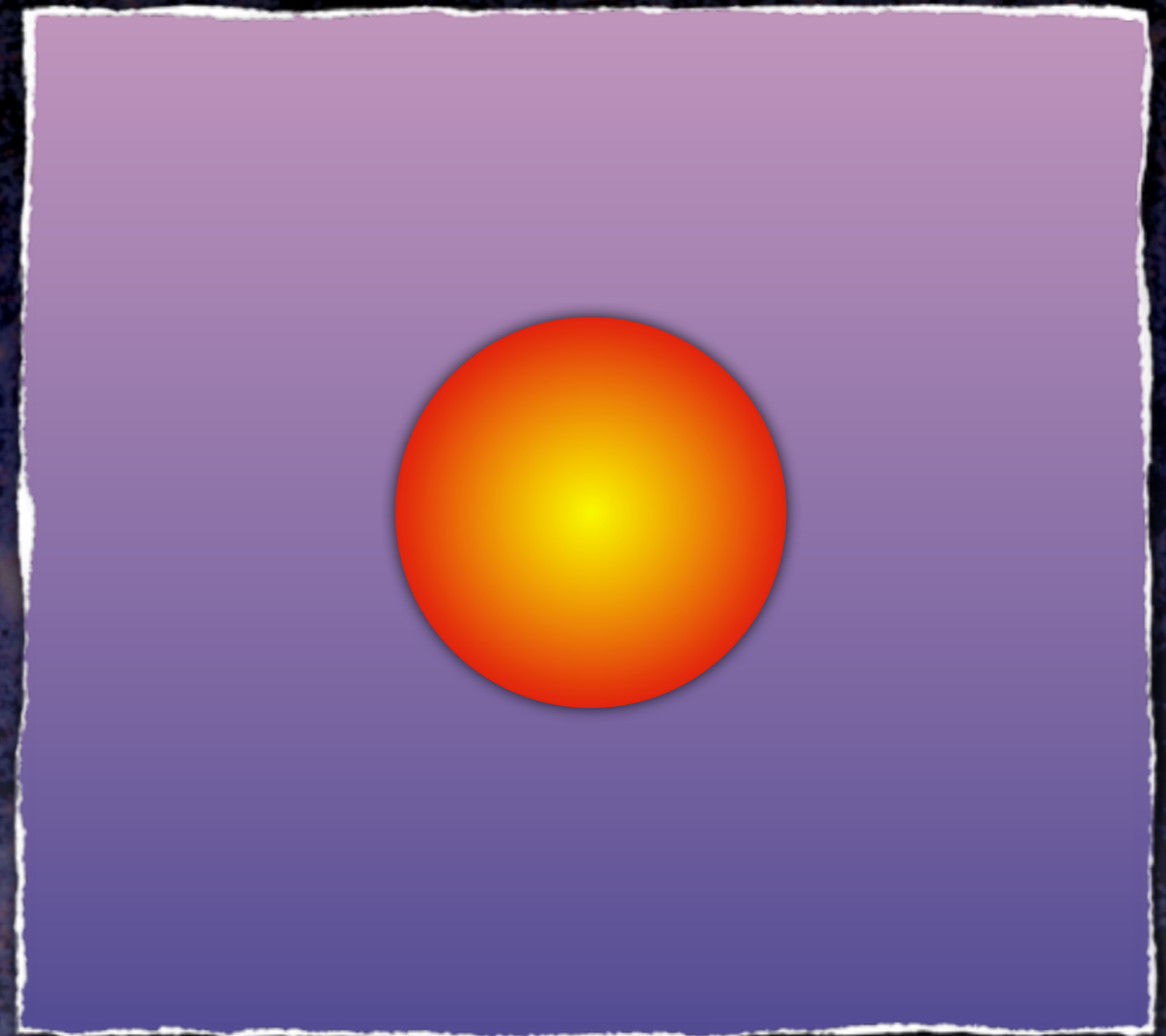


Evolution of shell of Baryonic Matter



Onion Model

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Because of pressure a shock develops,
which heats the gas and makes it expand

Evolution of shell of Baryonic Matter



Onion Model

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The End Result

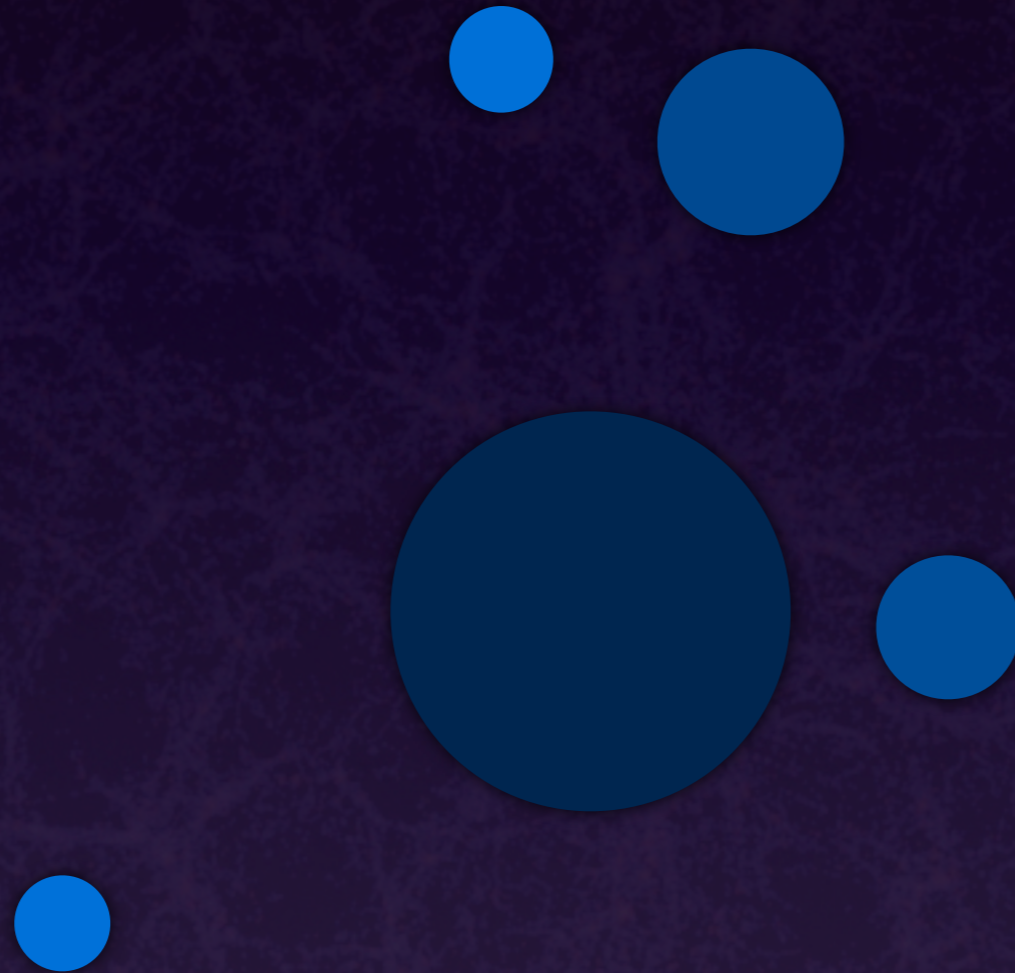


A dark matter halo
filled with hot gas

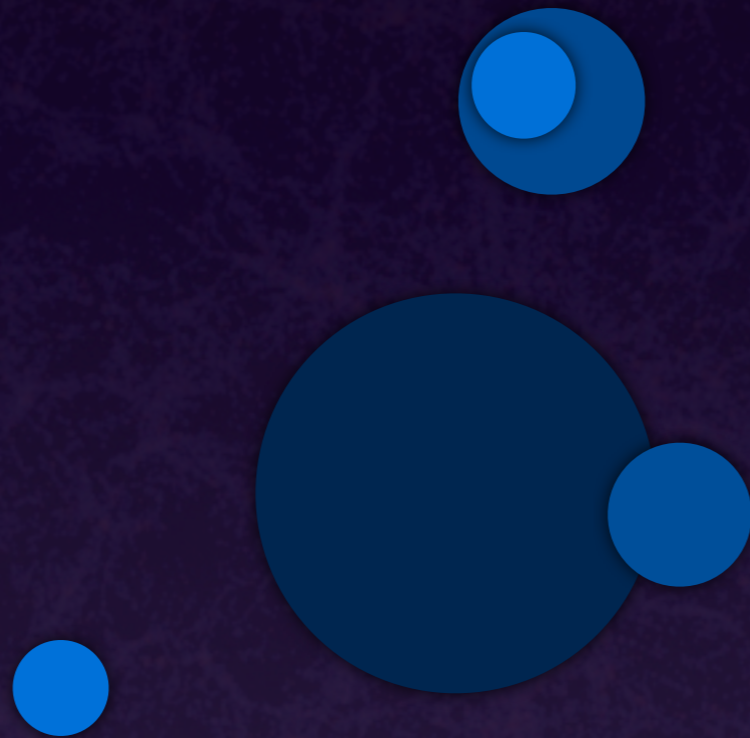
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The Hierarchical Growth of Dark Matter Haloes

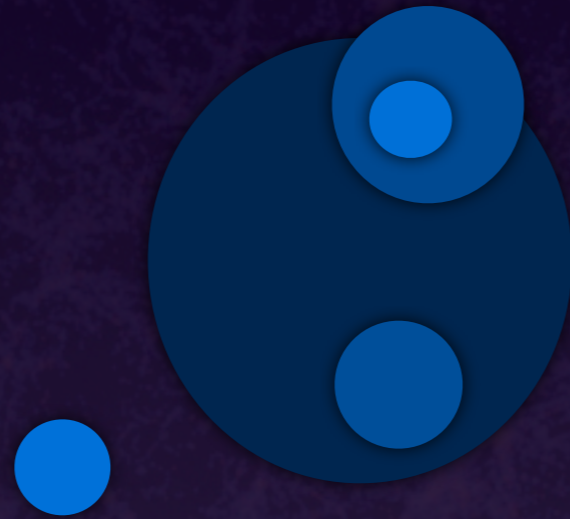
A region in space in which 5 dark matter haloes have formed



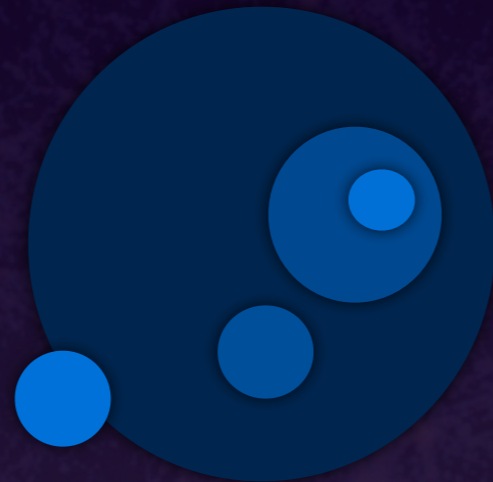
Dark matter haloes attract each other gravitationally...



consequently, they move towards each other....



and merge together, to form bigger haloes....



with substructure



Numerical Simulations

Start with box with many particles, whose spatial distribution reveals tiny fluctuations (as in CMB)

Let box `expand' (as Universe), and compute gravitational force between all particles



Propagate all particles according to the gravitational acceleration

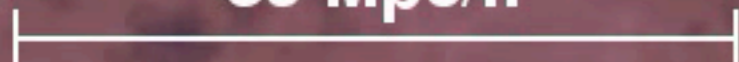
Repeat this procedure for as many time steps as needed

Distribution of dark matter with
tiny fluctuations in initial density



$z = 20.0$

50 Mpc/h



A visualization of the Millennium Simulation, showing a complex, interconnected network of particles. The particles are represented as small, glowing spheres in shades of purple, blue, and yellow, forming a dense, web-like structure. The background is a dark, deep purple. A horizontal scale bar is located at the top left, with the text "1 Gpc/h" above it. The text "Millennium Simulation" is written in yellow, and "10,077,696,000 particles" is written in white. The text "(z = 0)" is located in the bottom left corner.

1 Gpc/h

Millennium Simulation

10,077,696,000 particles

($z = 0$)

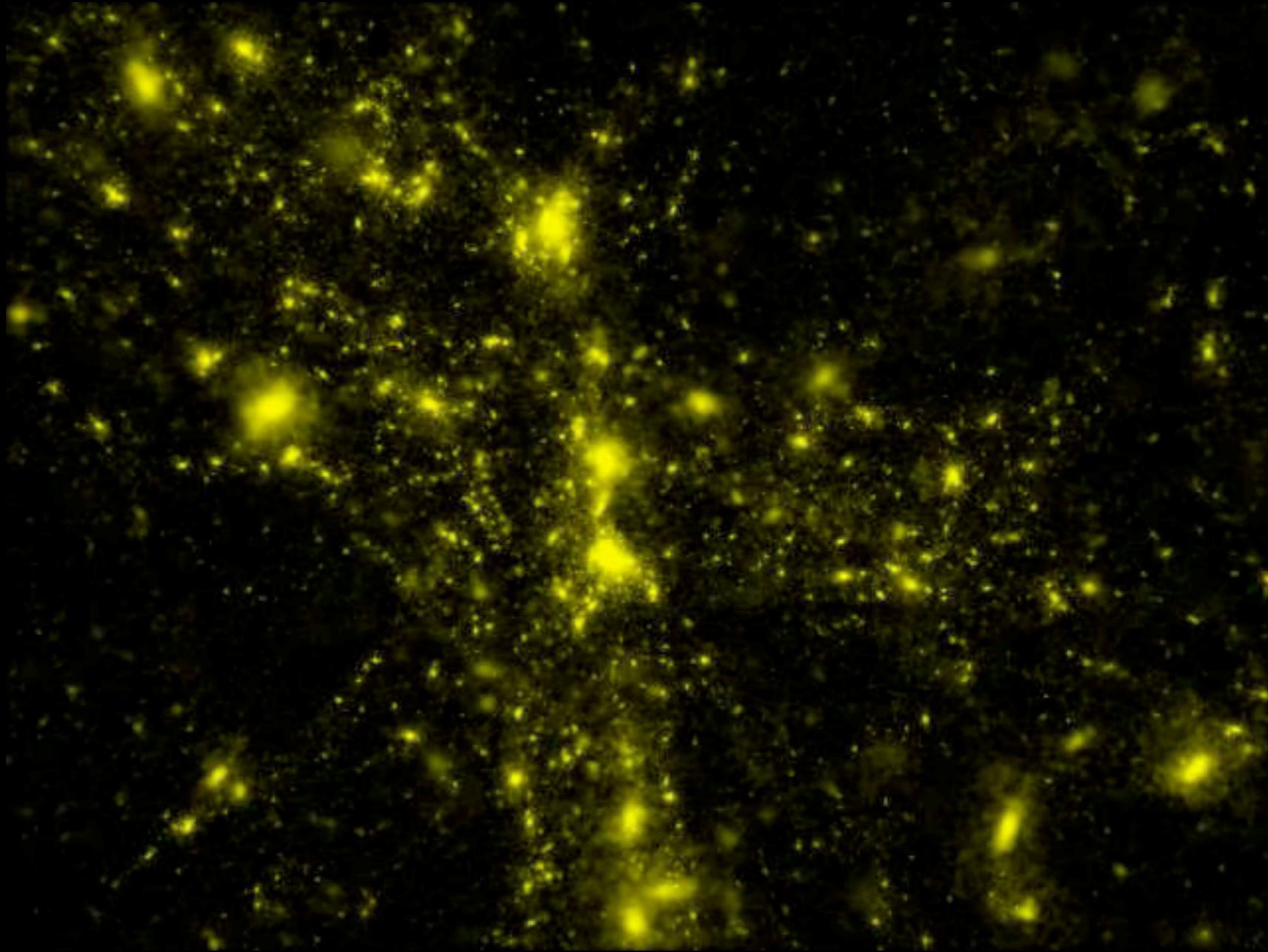
A Close-Up View of a Dark Matter Halo

A Close-Up View of a Dark Matter Halo



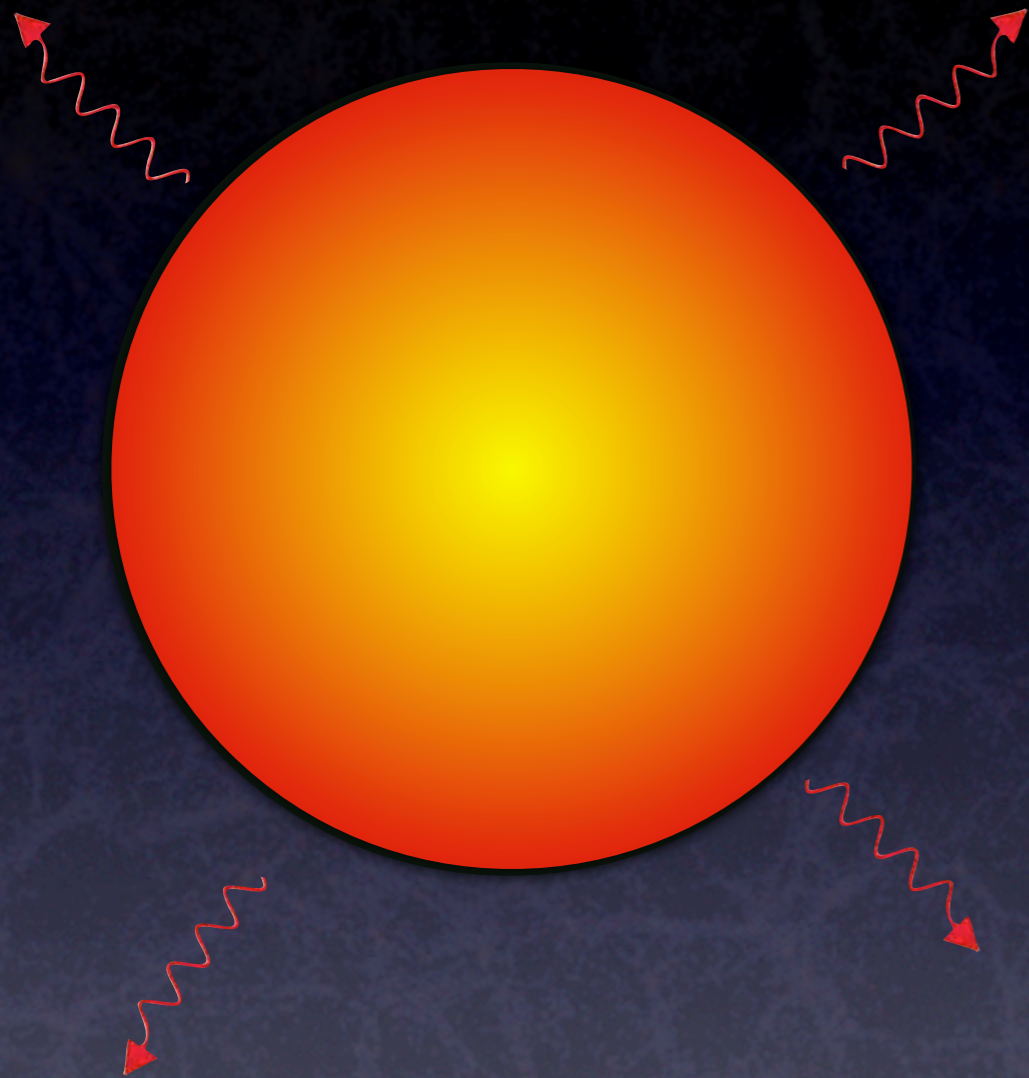
Flying Faster than the Speed of Light...

Flying Faster than the Speed of Light...



The Formation of Galaxies

Cooling & Disk Formation

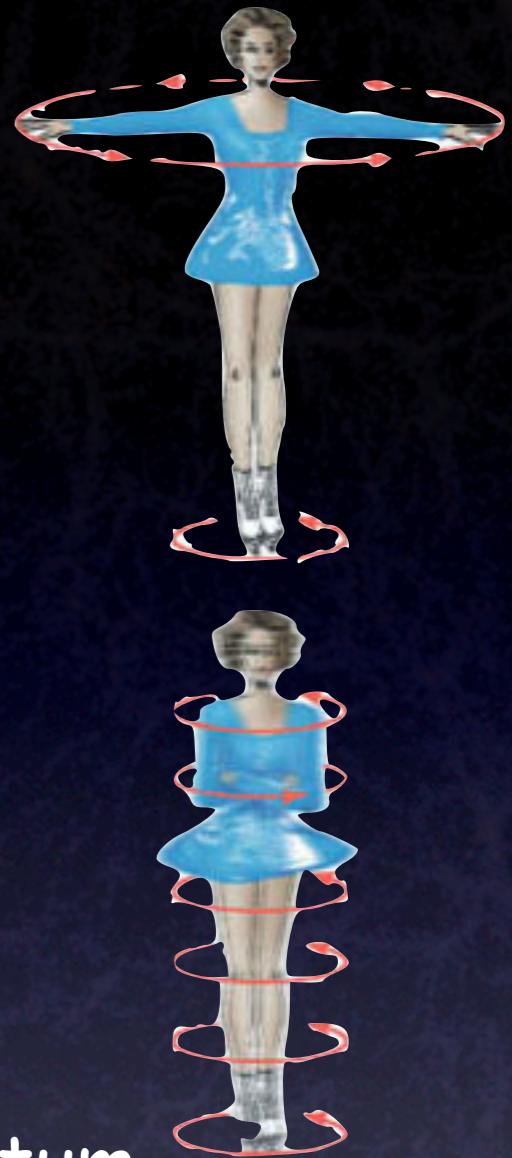


Hot gas radiates, emits photons which carry away energy: the gas cools

Due to pressure loss, gas starts to contract

Because of angular momentum conservation, the cooling baryons spin up and form a thin disk

Inside the disk the density gets very high, causing fragmentation and star formation: a disk galaxy is born...



Cooling & Disk Formation



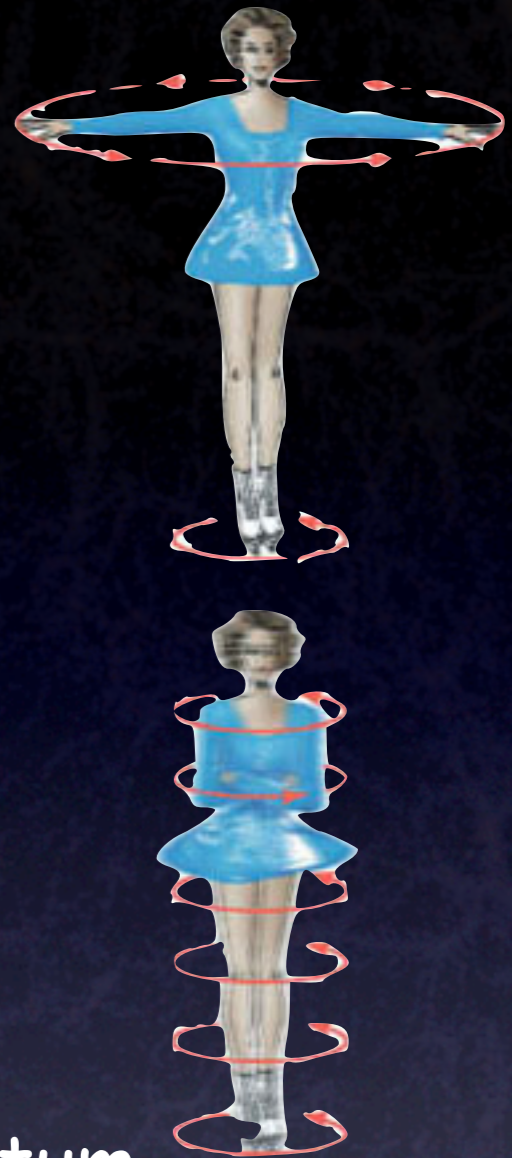
*Note that dark matter
does NOT cool*

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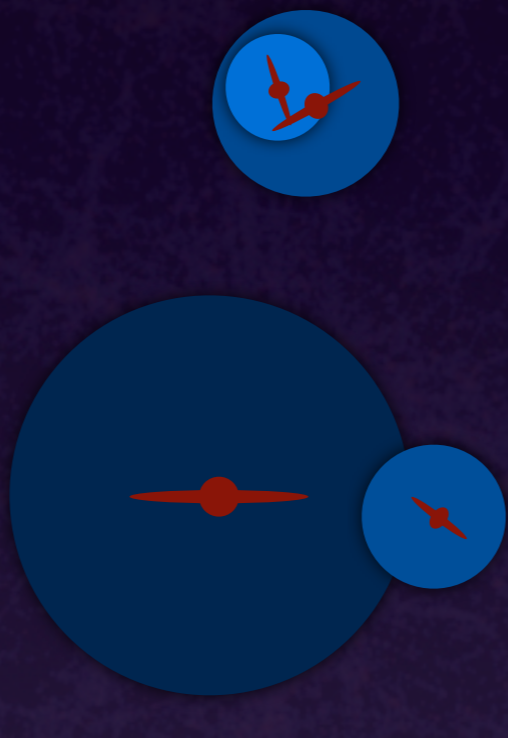
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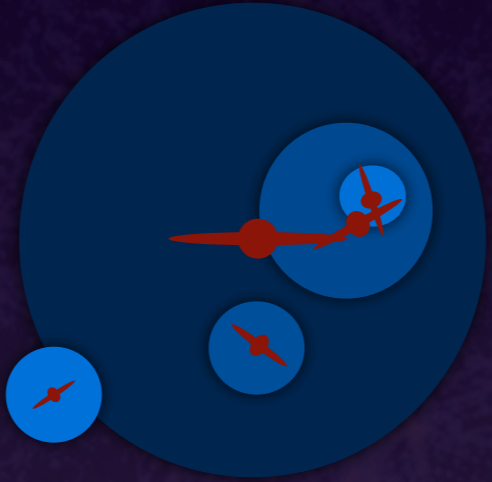
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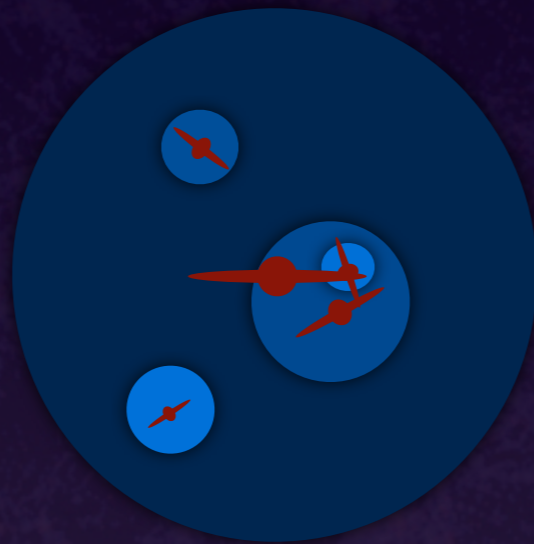




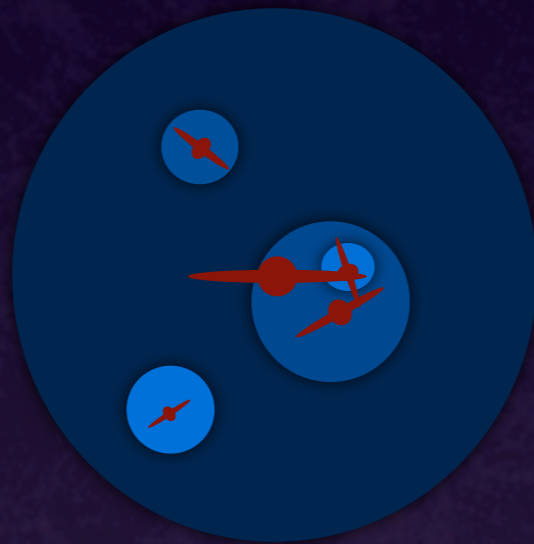




If dark matter haloes host galaxies,
clusters are a natural outcome of
hierarchical formation in CDM Universe



If dark matter haloes host galaxies,
clusters are a natural outcome of
hierarchical formation in CDM Universe



But what happens when two galaxies collide??

When two disks collide...

When two disks collide...



...an elliptical emerges

Galaxy Formation in a nutshell...

- ★ Small perturbations, due to quantum fluctuations, grow and collapse to form dark matter haloes
- ★ Baryonic gas is shock heated to high temperatures
- ★ Baryonic gas cools and settles in center of halo; angular momentum conservation --> disk galaxy
- ★ Disks merge giving rise to population of ellipticals especially in denser environments (clusters)



Outstanding Problems: Some Feedback Please...



Simple calculations of cooling rates and star formation efficiencies predict that virtually ALL baryons should have formed stars

Outstanding Problems: Some Feedback Please...



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But observations show that only ~10% of baryons have turned into stars

Outstanding Problems: Some Feedback Please...



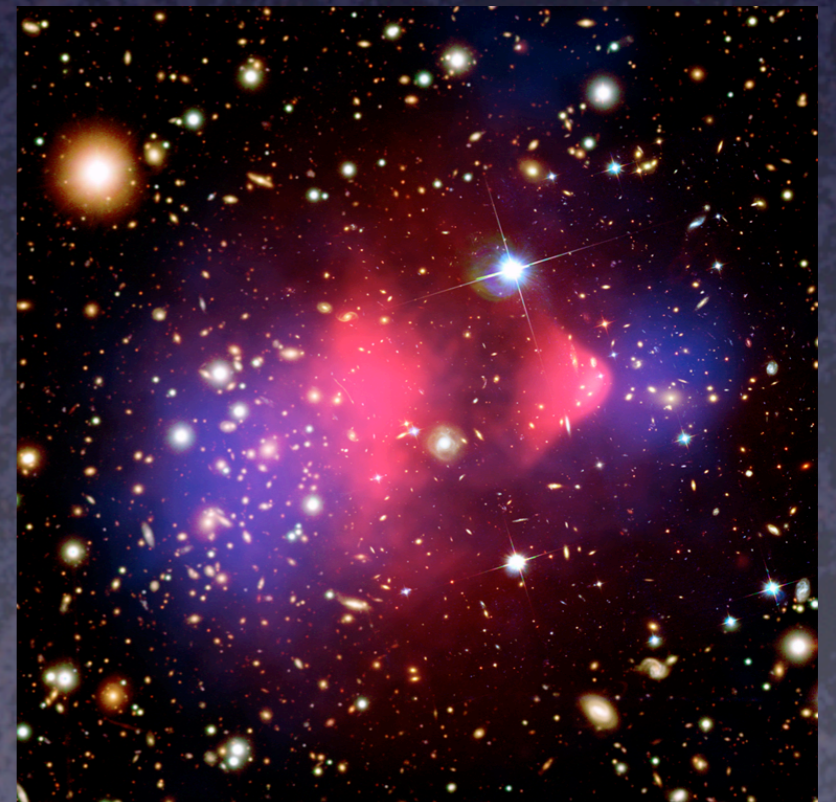
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Where are the other 90% ?

Hot gas in clusters (observed)

Warm-hot gas in filaments (elusive)



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But observations show that only ~10% of baryons have turned into stars

Where are the other 90% ?

Hot gas in clusters (observed)

Warm-hot gas in filaments (elusive)

Why so few stars ?

Feedback from supernovae & AGN?

We do NOT understand this process...





*Thank you
for believing...*



Remember me?