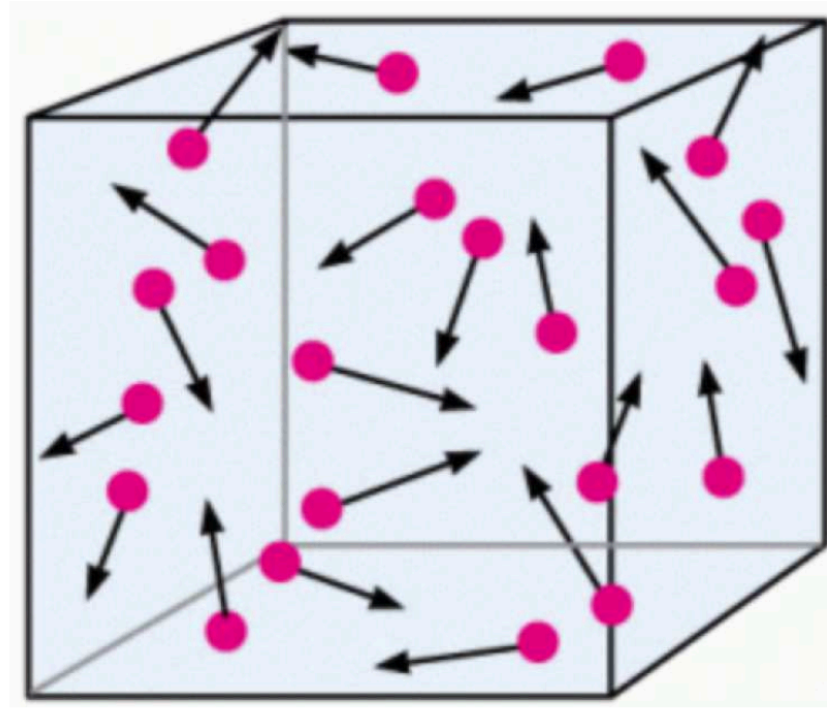


LECTURE 6

Part II: Kinetic Theory

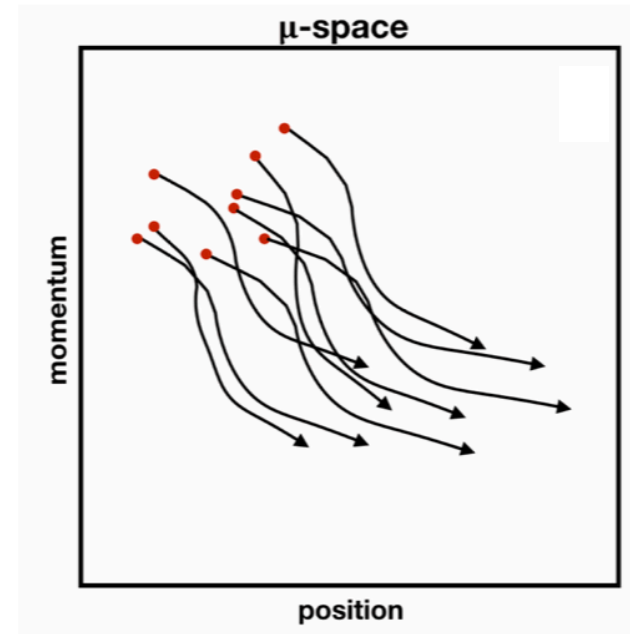
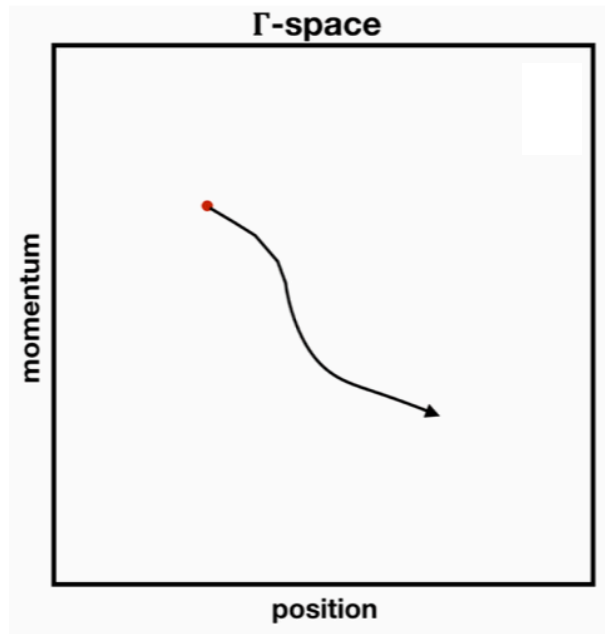


Part II: Kinetic Theory

- 6: From Liouville to Boltzmann
- 7: From Boltzmann to Navier-Stokes
- 8: Stochasticity & the Langevin Equation
- 9: The Fokker-Planck Equation

N-particle Phase-Space (Γ -space):

the $6N$ -dimensional phase-space of a dynamical system is a space in which all possible states **of a system** are represented, which each possible space corresponding to one unique point in that phase-space.



1-particle Phase-Space (μ -space):

the 6-dimensional phase-space of a dynamical system is a space covering all possible phase-space coordinates **of individual particles**. Each particles corresponding to one point in that phase-space.

Note: unlike in Γ -space, in which two trajectories can *never* intersect one-another, in μ -space the trajectories (of individual particles) *can* cross one another.

The BBGKY Hierarchy

$$\frac{df^{(N)}}{dt} = \frac{\partial f^{(N)}}{\partial t} + \{f^{(N)}, \mathcal{H}\} = 0$$

Liouville Theorem

$$\frac{\partial f^{(k)}}{\partial t} = \{\mathcal{H}^{(k)}, f^{(k)}\} + \sum_{i=1}^k \int d^3\vec{q}_{k+1} d^3\vec{p}_{k+1} \frac{\partial U(|\vec{q}_i - \vec{q}_{k+1}|)}{\partial \vec{q}_i} \cdot \frac{\partial f^{(k+1)}}{\partial \vec{p}_i}$$

$$\frac{\partial f^{(1)}}{\partial t} = \{\mathcal{H}^{(1)}, f^{(1)}\} + \int d^3\vec{q}_2 d^3\vec{p}_2 \frac{\partial U(|\vec{q}_1 - \vec{q}_2|)}{\partial \vec{q}_1} \cdot \frac{\partial f^{(2)}}{\partial \vec{p}_1}$$

Hamiltonian

$$\mathcal{H}(\vec{q}_i, \vec{p}_i) = \sum_{i=1}^N \frac{\vec{p}_i^2}{2m} + \sum_{i=1}^N V(\vec{q}_i) + \frac{1}{2} \sum_{i=1}^N \sum_{\substack{j=1 \\ j \neq i}}^N U(|\vec{q}_i - \vec{q}_j|)$$

external
potential

2-body
interaction
potential

reduced k -particle DF

$$f^{(k)}(\vec{w}_1, \vec{w}_2, \dots, \vec{w}_k, t) \equiv \frac{N!}{(N-k)!} \int \prod_{i=k+1}^N d^6\vec{w}_i f^{(N)}(\vec{w}_1, \vec{w}_2, \dots, \vec{w}_N, t)$$

one-particle DF

$$f^{(1)}(\vec{w}_1, t) \equiv N \int \prod_{i=2}^N d^6\vec{w}_i f^{(N)}(\vec{w}_1, \vec{w}_2, \dots, \vec{w}_N, t) \quad \longrightarrow \quad f^{(1)}(\vec{q}, \vec{p}, t) = dN/d^3\vec{q}d^3\vec{p}$$

The Mayer Cluster Expansion

$$f^{(2)}(\vec{q}_1, \vec{q}_2, \vec{p}_1, \vec{p}_2) = f^{(1)}(\vec{q}_1, \vec{p}_1) f^{(1)}(\vec{q}_2, \vec{p}_2) + g(\vec{q}_1, \vec{q}_2, \vec{p}_1, \vec{p}_2)$$



short-hand notation

$$f^{(2)}(1, 2) = f^{(1)}(1) f^{(1)}(2) + g(1, 2)$$

2-point correlation function

$$f^{(3)}(1, 2, 3) = f(1) f(2) f(3) + f(1) g(2, 3) + f(2) g(1, 3) + f(3) g(1, 2) + h(1, 2, 3)$$

etc.

3-point correlation function

Correlations are induced by collisions (interactions) among the particles

[A] Collisionless System $\rightarrow g(1,2)=0 \rightarrow$ Collisionless Boltzmann Equation (CBE)

$$\frac{df^{(1)}}{dt} = \frac{\partial f^{(1)}}{\partial t} + \{f^{(1)}, \mathcal{H}^{(1)}\} = \frac{\partial f^{(1)}}{\partial t} + \vec{v} \cdot \frac{\partial f^{(1)}}{\partial \vec{x}} - \nabla \Phi \cdot \frac{\partial f^{(1)}}{\partial \vec{v}} = 0$$

[B] System with short-range collisions (neutral gas or liquid)

assumption of **molecular chaos**; $f^{(2)}(\vec{q}, \vec{q}, \vec{p}_1, \vec{p}_2) = f^{(1)}(\vec{q}, \vec{p}_1) f^{(1)}(\vec{q}, \vec{p}_2)$

\rightarrow Boltzmann Equation

$$\frac{df^{(1)}}{dt} = \frac{\partial f^{(1)}}{\partial t} + \{f^{(1)}, \mathcal{H}^{(1)}\} = \frac{\partial f^{(1)}}{\partial t} + \vec{v} \cdot \frac{\partial f^{(1)}}{\partial \vec{x}} - \nabla \Phi \cdot \frac{\partial f^{(1)}}{\partial \vec{v}} = I[f^{(1)}]$$

[C] System with long-range collisions (collisional Plasma or low-N gravitational system)

assumptions; $h(1,2,3)=0$ + fluid is homogeneous + $g(1,2)$ relaxes faster than $f^{(1)}$

\rightarrow Lenard-Balescu equation

$$\frac{\partial f(\vec{v}, t)}{\partial t} = -\frac{8\pi^4 n_e}{m_e^2} \frac{\partial}{\partial \vec{v}} \int d\vec{k} d\vec{v}' \vec{k} \vec{k} \cdot \frac{\phi^2(k)}{|\varepsilon(\vec{k}, \vec{k} \cdot \vec{v})|^2} \delta[\vec{k} \cdot (\vec{v} - \vec{v}')] \left[f(\vec{v}) \frac{\partial f}{\partial \vec{v}'} - f(\vec{v}') \frac{\partial f}{\partial \vec{v}} \right]$$

see chapter 27

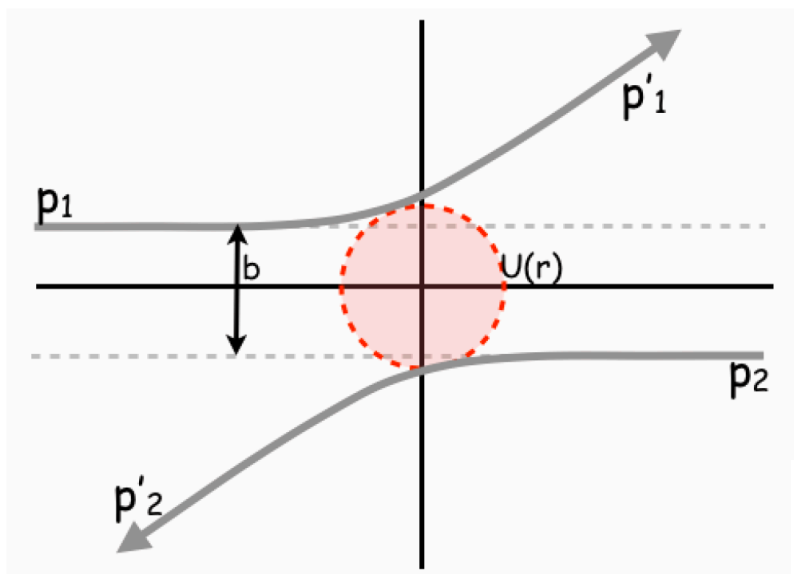
here $f = f^{(1)}$

From Boltzmann to Navier-Stokes

In what follows we focus on the 1-particle distribution function f , dropping the ⁽¹⁾-superscript

We consider **collision integral term** in the **Boltzmann equation**:

$$I[f] = (\partial f / \partial t)_{\text{coll}} = \int d^3 \vec{q}_2 d^3 \vec{p}_2 \frac{\partial U(|\vec{q}_1 - \vec{q}_2|)}{\partial \vec{q}_1} \cdot \frac{\partial f^{(2)}}{\partial \vec{p}_1}$$



consider the following **elastic** two-particle collision

$$\vec{p}_1 + \vec{p}_2 \rightarrow \vec{p}'_1 + \vec{p}'_2$$

these collisions obey:

momentum conservation: $\vec{p}_1 + \vec{p}_2 = \vec{p}'_1 + \vec{p}'_2$
energy conservation: $|\vec{p}_1|^2 + |\vec{p}_2|^2 = |\vec{p}'_1|^2 + |\vec{p}'_2|^2$

Write the rate at which particles of momentum \vec{p}_1 at \vec{x} experience collisions $\vec{p}_1 + \vec{p}_2 \rightarrow \vec{p}_1' + \vec{p}_2'$ as:

$$\mathcal{R} = \omega(\vec{p}_1, \vec{p}_2 | \vec{p}_1', \vec{p}_2') f^{(2)}(\vec{x}, \vec{x}, \vec{p}_1, \vec{p}_2) d^3\vec{p}_2 d^3\vec{p}_1' d^3\vec{p}_2'$$

The function $\omega(\vec{p}_1, \vec{p}_2 | \vec{p}_1', \vec{p}_2')$ depends on the **interaction potential** $U(\mathbf{r})$ and can be calculated (in principle) via **differential cross sections**

Momentum & energy conservation $\rightarrow \omega(\vec{p}_1, \vec{p}_2 | \vec{p}_1', \vec{p}_2') \propto \delta^3(\vec{P} - \vec{P}') \delta(E - E')$

with $\vec{P} = \vec{p}_1 + \vec{p}_2$ and $\vec{P}' = \vec{p}_1' + \vec{p}_2'$

Time reversibility $\rightarrow \omega(\vec{p}_1, \vec{p}_2 | \vec{p}_1', \vec{p}_2') = \omega(\vec{p}_1', \vec{p}_2' | \vec{p}_1, \vec{p}_2)$

Using principle of **molecular chaos** $\rightarrow f^{(2)}(\vec{x}, \vec{x}, \vec{p}_1, \vec{p}_2) = f^{(1)}(\vec{x}, \vec{p}_1) f^{(1)}(\vec{x}, \vec{p}_2)$



$$I[f] = \int d^3\vec{p}_2 d^3\vec{p}_1' d^3\vec{p}_2' \omega(\vec{p}_1', \vec{p}_2' | \vec{p}_1, \vec{p}_2) [f(\vec{p}_1') f(\vec{p}_2') - f(\vec{p}_1) f(\vec{p}_2)]$$

replenishing
collisions

depleting
collisions

for brevity we no longer write out the explicit x -dependence of the DF

What can we learn about the **equilibrium** distribution function, $f_{\text{eq}}(\mathbf{x}, \mathbf{p})$?

$$\left. \begin{array}{l} \text{Equilibrium} \rightarrow \partial f_{\text{eq}} / \partial t = 0 \\ \text{Ignore external potential \& spatial homogeneity} \rightarrow \{\mathcal{H}, f_{\text{eq}}\} = 0 \end{array} \right\} I[f] = 0$$

$$I[f] = \int d^3\vec{p}_2 d^3\vec{p}_1' d^3\vec{p}_2' \omega(\vec{p}_1', \vec{p}_2' | \vec{p}_1, \vec{p}_2) [f(\vec{p}_1') f(\vec{p}_2') - f(\vec{p}_1) f(\vec{p}_2)] = 0$$

$$\text{Detailed balance} \rightarrow f(\vec{x}, \vec{p}_1') f(\vec{x}, \vec{p}_2') - f(\vec{x}, \vec{p}_1) f(\vec{x}, \vec{p}_2) = 0$$

$$\rightarrow \log[f(\vec{p}_1)] + \log[f(\vec{p}_2)] = \log[f(\vec{p}_1')] + \log[f(\vec{p}_2')]$$

This has form of a **conservation law**, and suggests that $\log[f]$ must be equal to sum of conserved quantities, $A(\mathbf{p})$, that obey $A(\vec{p}_1) + A(\vec{p}_2) = A(\vec{p}_1') + A(\vec{p}_2')$

We have the following **collisional invariants**:

$A = 1$	particle number conservations
$A = \vec{p}$	momentum conservation
$A = \vec{p}^2 / (2m)$	energy conservation

This therefore suggests that $\log[f_{\text{eq}}(\vec{p})] \propto a_1 + a_2 \vec{p} + a_3 |\vec{p}|^2$

This therefore suggests that $\log[f_{\text{eq}}(\vec{p})] \propto a_1 + a_2 \vec{p} + a_3 |\vec{p}|^2$

It can be shown that this implies the **Maxwell-Boltzmann distribution**

$$f_{\text{eq}}(p) = \frac{n}{(2\pi m k_B T)^{3/2}} \exp \left[-\frac{p^2}{2m k_B T} \right]$$

In other words: **the MB-distribution is the equilibrium solution of the Boltzmann equation**

It can be shown that

$$\boxed{\int d^3\vec{p} A(\vec{p}) \left(\frac{\partial f}{\partial t} \right)_{\text{coll}} = 0}$$

Proof:

define
$$\mathcal{I}_1 = \int d^3\vec{p}_1 d^3\vec{p}_2 d^3\vec{p}_1' d^3\vec{p}_2' \omega(\vec{p}_1', \vec{p}_2' | \vec{p}_1, \vec{p}_2) A(\vec{p}_1) [f(\vec{p}_1') f(\vec{p}_2') - f(\vec{p}_1) f(\vec{p}_2)]$$

re-labelling 1 \leftrightarrow 2 and re-ordering yields

$$\mathcal{I}_2 = \int d^3\vec{p}_1 d^3\vec{p}_2 d^3\vec{p}_1' d^3\vec{p}_2' \omega(\vec{p}_1', \vec{p}_2' | \vec{p}_1, \vec{p}_2) A(\vec{p}_2) [f(\vec{p}_1') f(\vec{p}_2') - f(\vec{p}_1) f(\vec{p}_2)]$$

Starting from the first expression and swapping $p_1 \leftrightarrow p_1'$ yields

$$\mathcal{I}_3 = - \int d^3\vec{p}_1 d^3\vec{p}_2 d^3\vec{p}_1' d^3\vec{p}_2' \omega(\vec{p}_1, \vec{p}_2 | \vec{p}_1', \vec{p}_2') A(\vec{p}_1') [f(\vec{p}_1') f(\vec{p}_2') - f(\vec{p}_1) f(\vec{p}_2)]$$

re-labelling 1 \leftrightarrow 2 and re-ordering yields

$$\mathcal{I}_4 = - \int d^3\vec{p}_1 d^3\vec{p}_2 d^3\vec{p}_1' d^3\vec{p}_2' \omega(\vec{p}_1, \vec{p}_2 | \vec{p}_1', \vec{p}_2') A(\vec{p}_2') [f(\vec{p}_1') f(\vec{p}_2') - f(\vec{p}_1) f(\vec{p}_2)]$$

Time reversibility $\omega(\vec{p}_1', \vec{p}_2' | \vec{p}_1, \vec{p}_2) = \omega(\vec{p}_1, \vec{p}_2 | \vec{p}_1', \vec{p}_2')$ implies that $\mathcal{I}_4 = \mathcal{I}_3 = \mathcal{I}_2 = \mathcal{I}_1$

and thus $\mathcal{I}_1 = [\mathcal{I}_1 + \mathcal{I}_2 + \mathcal{I}_3 + \mathcal{I}_4]/4$

$$\mathcal{I}_1 = [\mathcal{I}_1 + \mathcal{I}_2 + \mathcal{I}_3 + \mathcal{I}_4]/4$$

substituting the expressions from the previous page:

$$\mathcal{I}_1 = \frac{1}{4} \int d^3\vec{p}_1 d^3\vec{p}_2 d^3\vec{p}_1' d^3\vec{p}_2' \omega(\vec{p}_1', \vec{p}_2' | \vec{p}_1, \vec{p}_2) \times \\ \{A(\vec{p}_1) + A(\vec{p}_2) - A(\vec{p}_1') - A(\vec{p}_2')\} [f(\vec{p}_1') f(\vec{p}_2') - f(\vec{p}_1) f(\vec{p}_2)]$$

$$A(\mathbf{p}) \text{ is a collisional invariant} \rightarrow A(\vec{p}_1) + A(\vec{p}_2) - A(\vec{p}_1') - A(\vec{p}_2') = 0 \rightarrow \mathcal{I}_1 = 0$$

Q.E.D.

As long as $A(\mathbf{p})$ is a collisional invariant we have that

$$\boxed{\int d^3\vec{p} A(\vec{p}) \left(\frac{\partial f}{\partial t} \right)_{\text{coll}} = 0}$$

we will use this shortly to obtain the Navier-Stokes equations from the Boltzmann equation

Solving the Boltzmann equation

$$\frac{df^{(1)}}{dt} = \frac{\partial f^{(1)}}{\partial t} + \{f^{(1)}, \mathcal{H}^{(1)}\} = \frac{\partial f^{(1)}}{\partial t} + \vec{v} \cdot \frac{\partial f^{(1)}}{\partial \vec{x}} - \nabla \Phi \cdot \frac{\partial f^{(1)}}{\partial \vec{v}} = I[f^{(1)}]$$

for the 7-dimensional DF $f(\mathbf{x}, \mathbf{v}, t)$ is a non-trivial task

Rather, we are going to solve **moment equations** of the Boltzmann equation

Consider a scalar function $Q(\mathbf{v})$. The expectation value for Q at location \mathbf{x} at time t is given by

$$\langle Q \rangle = \langle Q \rangle(\vec{x}, t) = \frac{\int Q(\vec{v}) f(\vec{x}, \vec{v}, t) d^3\vec{v}}{\int f(\vec{x}, \vec{v}, t) d^3\vec{v}}$$

using that

$$n = n(\vec{x}, t) = \int f(\vec{x}, \vec{v}, t) d^3\vec{v}$$

we see that

$$\int Q(\vec{v}) f(\vec{x}, \vec{v}, t) d^3\vec{v} = n \langle Q \rangle$$

Define $g(\vec{x}, t) = \int Q(\vec{v}) f(\vec{x}, \vec{v}, t) d^3\vec{v}$

$Q(\vec{v}) = 1$	\Rightarrow	$g(\vec{x}, t) = n(\vec{x}, t)$	number density
$Q(\vec{v}) = m$	\Rightarrow	$g(\vec{x}, t) = \rho(\vec{x}, t)$	mass density
$Q(\vec{v}) = m\vec{v}$	\Rightarrow	$g(\vec{x}, t) = \rho(\vec{x}, t) \vec{u}(\vec{x}, t)$	momentum flux density
$Q(\vec{v}) = \frac{1}{2}m(\vec{v} - \vec{u})^2$	\Rightarrow	$g(\vec{x}, t) = \rho(\vec{x}, t) \varepsilon(\vec{x}, t)$	specific energy density



these are our **macroscopic quantities** of interest...

NOTE: $\mathbf{u}(\mathbf{x}, t) = \langle \mathbf{v}(\mathbf{x}, t) \rangle$ is the mean velocity of all particles at location \mathbf{x} at time t

Rather than solving the Boltzmann equation $\frac{\partial f}{\partial t} + \vec{v} \cdot \nabla f - \nabla \Phi \cdot \frac{\partial f}{\partial \vec{v}} = \left(\frac{\partial f}{\partial t} \right)_{\text{coll}}$

we seek to solve the following moment equations:

$$\int Q(\vec{v}) \left[\frac{\partial f}{\partial t} + \vec{v} \cdot \nabla f - \nabla \Phi \cdot \frac{\partial f}{\partial \vec{v}} \right] d^3 \vec{v} = \int Q(\vec{v}) \left(\frac{\partial f}{\partial t} \right)_{\text{coll}} d^3 \vec{v}$$

In particular, we will focus on $Q(\mathbf{v})$ that are collisional invariants for which the rhs vanishes

$$\int Q(\vec{v}) \frac{\partial f}{\partial t} d^3 \vec{v} + \int Q(\vec{v}) \vec{v} \cdot \nabla f d^3 \vec{v} - \int Q(\vec{v}) \nabla \Phi \cdot \frac{\partial f}{\partial \vec{v}} d^3 \vec{v} = 0$$

NOTE: if we had started with the CBE, rather than the Boltzmann equation, we would have obtained the same moment equations.

HENCE: what follows is valid for both collisional (short-range collisions) and collisionless systems

$$\int Q(\vec{v}) \frac{\partial f}{\partial t} d^3\vec{v} + \int Q(\vec{v}) \vec{v} \cdot \nabla f d^3\vec{v} - \int Q(\vec{v}) \nabla\Phi \cdot \frac{\partial f}{\partial \vec{v}} d^3\vec{v} = 0$$

$$\text{I} \quad + \quad \text{II} \quad - \quad \text{III} \quad = 0$$

$\begin{aligned} \text{I} & \int Q(\vec{v}) \frac{\partial f}{\partial t} d^3\vec{v} \\ \text{II} & \int Q(\vec{v}) v_i \frac{\partial f}{\partial x_i} d^3\vec{v} \\ \text{III} & \int Q(\vec{v}) \frac{\partial\Phi}{\partial x_i} \frac{\partial f}{\partial v_i} d^3\vec{v} \end{aligned}$
--

Integral I

$$\int Q(\vec{v}) \frac{\partial f}{\partial t} d^3\vec{v} = \int \frac{\partial Q f}{\partial t} d^3\vec{v} = \frac{\partial}{\partial t} \int Q f d^3\vec{v} = \frac{\partial}{\partial t} n \langle Q \rangle$$

Integral II

$$\int Q(\vec{v}) v_i \frac{\partial f}{\partial x_i} d^3\vec{v} = \int \frac{\partial Q v_i f}{\partial x_i} d^3\vec{v} = \frac{\partial}{\partial x_i} \int Q v_i f d^3\vec{v} = \frac{\partial}{\partial x_i} [n \langle Q v_i \rangle]$$

where we have used that

$$Q v_i \frac{\partial f}{\partial x_i} = \frac{\partial(Q v_i f)}{\partial x_i} - f \frac{\partial Q v_i}{\partial x_i} = \frac{\partial(Q v_i f)}{\partial x_i}$$

Integral III

$$\begin{aligned}\int Q \vec{F} \cdot \nabla_v f d^3\vec{v} &= \int \nabla_v \cdot (Q f \vec{F}) d^3\vec{v} - \int f \nabla_v \cdot (Q \vec{F}) d^3\vec{v} \\ &= \int Q f \vec{F} d^2S_v - \int f \frac{\partial Q F_i}{\partial v_i} d^3\vec{v} \\ &= - \int f Q \frac{\partial F_i}{\partial v_i} d^3\vec{v} - \int f F_i \frac{\partial Q}{\partial v_i} d^3\vec{v} \\ &= - \int f \frac{\partial \Phi}{\partial x_i} \frac{\partial Q}{\partial v_i} d^3\vec{v} = - \frac{\partial \Phi}{\partial x_i} n \left\langle \frac{\partial Q}{\partial v_i} \right\rangle\end{aligned}$$

$$\text{I} + \text{II} - \text{III} = 0 \quad \rightarrow$$

$$\frac{\partial}{\partial t} n \langle Q \rangle + \frac{\partial}{\partial x_i} [n \langle Q v_i \rangle] + \frac{\partial \Phi}{\partial x_i} n \left\langle \frac{\partial Q}{\partial v_i} \right\rangle = 0$$

Master Moment equation

This master moment equation holds for any **collisional invariant**, $Q(\mathbf{v})$, and for both **collisionless** and **collisional** (short-range forces only) systems

Let's consider $Q=m$ (mass conservation) and substitute this in the **Master Moment Equation**

using that $\langle m \rangle = m$, that $mn = \rho$, and that $\langle mv_i \rangle = m \langle v_i \rangle = mu_i$ we obtain that

$$\boxed{\frac{\partial \rho}{\partial t} + \frac{\partial \rho u_i}{\partial x_i} = 0}$$

which is the **continuity equation** (in index-form)

Let's consider $Q=mv_j$ (momentum conservation) and substitute this in the **Master Moment Equation**

using that $n\langle mv_j v_i \rangle = \rho\langle v_i v_j \rangle$, and that $\frac{\partial \Phi}{\partial x_i} n \left\langle \frac{\partial m v_j}{\partial v_i} \right\rangle = \frac{\partial \Phi}{\partial x_i} \rho \left\langle \frac{\partial v_j}{\partial v_i} \right\rangle = \frac{\partial \Phi}{\partial x_i} \rho \delta_{ij} = \rho \frac{\partial \Phi}{\partial x_j}$
 one finds that

$$\frac{\partial \rho u_j}{\partial t} + \frac{\partial \rho \langle v_i v_j \rangle}{\partial x_i} + \rho \frac{\partial \Phi}{\partial x_j} = 0$$

As detailed in the lecture notes, this can be manipulated to yield the **momentum equations**

$$\rho \frac{\partial u_j}{\partial t} + \rho u_i \frac{\partial u_j}{\partial x_i} + \frac{\partial [\rho \langle v_i v_j \rangle - \rho u_i u_j]}{\partial x_i} + \rho \frac{\partial \Phi}{\partial x_j} = 0$$

In order to make sense of the $\rho \langle v_i v_j \rangle - \rho u_i u_j$ term, it is useful to write the **microscopic particle velocity**, v , as the sum of a **streaming motion**, u , and a **random velocity**, w

$$\vec{v} = \vec{u} + \vec{w}$$

we have that $\langle \vec{w} \rangle = 0$ and $\langle \vec{v} \rangle = \vec{u}$

$\langle \cdot \rangle$ indicates an average over nearby particles (sometimes called a **fluid element**)

It is convenient to introduce the **stress tensor**

$$\sigma_{ij} \equiv -\rho \langle w_i w_j \rangle = -\rho \langle v_i v_j \rangle + \rho u_i u_j$$

Substituting this in the **momentum equations** we obtain the more common form:

$$\frac{\partial u_j}{\partial t} + u_i \frac{\partial u_j}{\partial x_i} = \frac{1}{\rho} \frac{\partial \sigma_{ij}}{\partial x_i} - \frac{\partial \Phi}{\partial x_j}$$

Momentum equations

Recall that these apply to both collisional (short-range force) and collisionless systems....

QUESTION: where has the impact of collisions (recall we started with a **collision integral**) gone??

ANSWER: it is 'hidden' in the detailed expression for the **stress tensor**

Momentum equations

$$\frac{\partial u_j}{\partial t} + u_i \frac{\partial u_j}{\partial x_i} = \frac{1}{\rho} \frac{\partial \sigma_{ij}}{\partial x_i} - \frac{\partial \Phi}{\partial x_j}$$

Collisionless Fluid:

No collisions, there is nothing special we can say about the stress tensor.

Momentum equations take on form above and are called the **Jeans equations**

Collisional Fluid:

As we will see, as long as the fluid is **Newtonian**, we have that $\sigma_{ij} = -P\delta_{ij} + \tau_{ij}$ with P the **hydrodynamic pressure**, and

$$\tau_{ij} = \mu \left[\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} - \frac{2}{3} \delta_{ij} \frac{\partial u_k}{\partial x_k} \right] + \eta \delta_{ij} \frac{\partial u_k}{\partial x_k}$$

the **deviatoric stress tensor** with μ and η the **shear viscosity** and **bulk viscosity**, respectively.

Substituting this in the momentum equations yields the **Navier-Stokes equations**

Setting $\mu = \eta = 0$ (assuming an **ideal fluid**), these become the **Euler equations**

Let's consider $Q=mv^2/2$ (energy conservation) and substitute this in the **Master Moment Equation**

As detailed in **Appendix J** of the lecture notes, working out the various terms ultimately yields

$$\frac{\partial}{\partial t} \left[\rho \left(\frac{u^2}{2} + \varepsilon \right) \right] = - \frac{\partial}{\partial x_k} \left[\rho \left(\frac{u^2}{2} + \varepsilon \right) u_k - \sigma_{jk} u_j + \rho \langle w_k \frac{1}{2} w^2 \rangle \right] - \rho u_k \frac{\partial \Phi}{\partial x_k}$$

which is known as the **energy equation** (in Lagrangian index-form).

Here $\varepsilon = \frac{1}{2} \langle w^2 \rangle$ is the **specific internal energy**.

As we will see in Part III of the lecture notes, the energy equation (for collisional fluids) can be recast as

$$\rho \frac{d\varepsilon}{dt} = -P \frac{\partial u_k}{\partial x_k} + \mathcal{V} - \frac{\partial F_{\text{cond},k}}{\partial x_k}$$

with \mathcal{V} the **rate of viscous dissipation**, and \mathbf{F}_{cond} the **conductive heat flux**.