The Role of Magnetic Fields in Population III Star Formation

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"What about magnetic fields?"

Increasingly well-explored space ...

Xu, et al (2008)

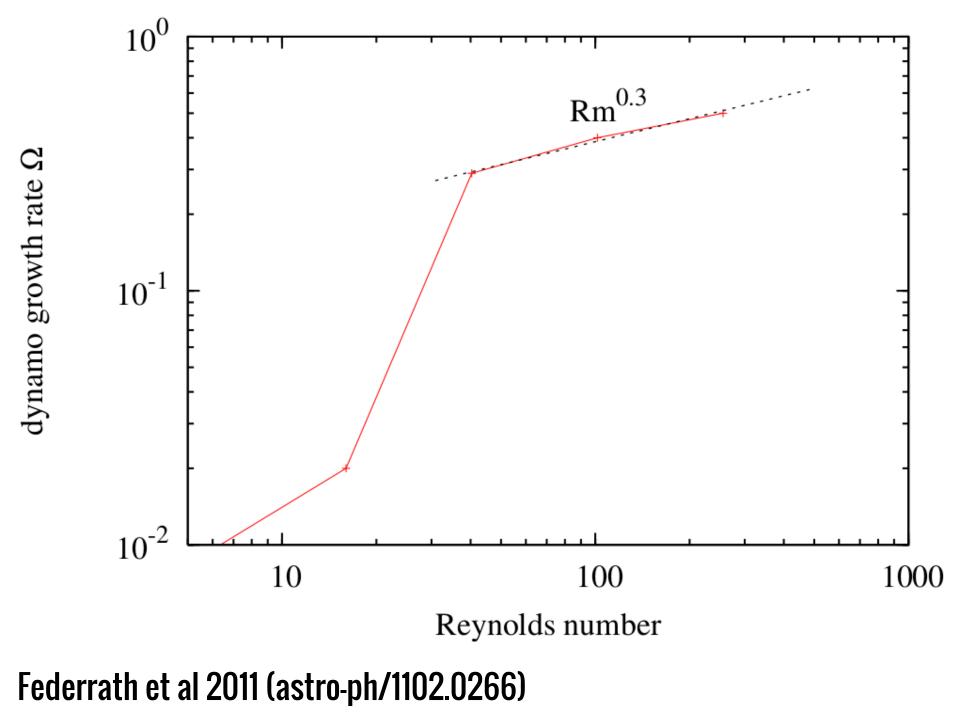
Biermann battery: <u>not</u> sufficient for dynamical importance!

(astro-ph/0807.2647)

Federrath, et al (2011)

Small-scale dynamo action on scales less than Jeans Length

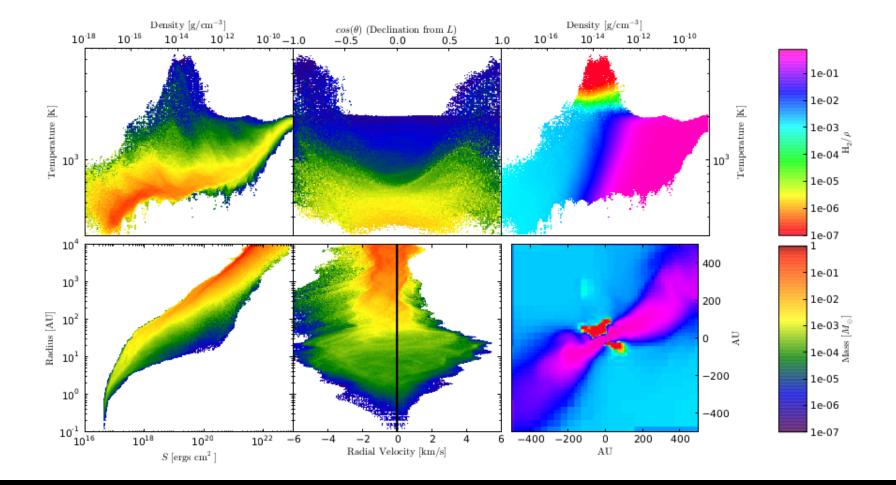
(astro-ph/1102.0266)



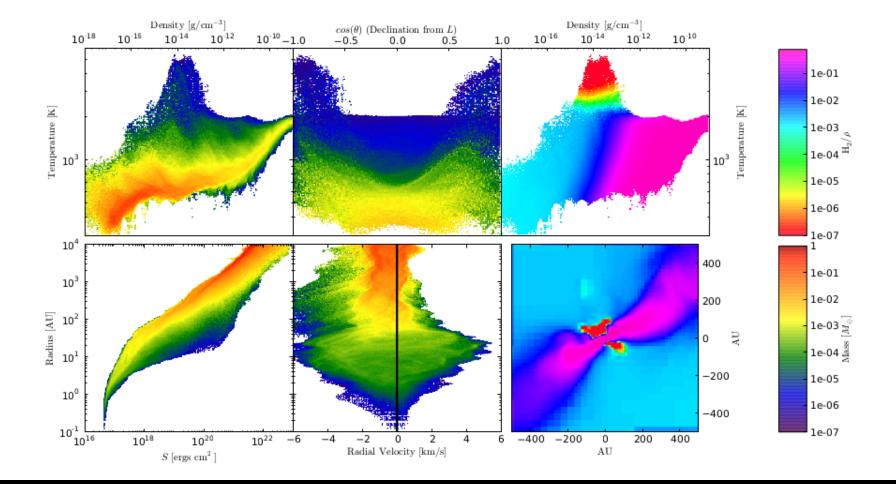
Schober, et al (2012)

Saturation of magnetic field energy from small-scale dynamo

(astro-ph/1204.0658)



Do magnetic fields alter disk structure?

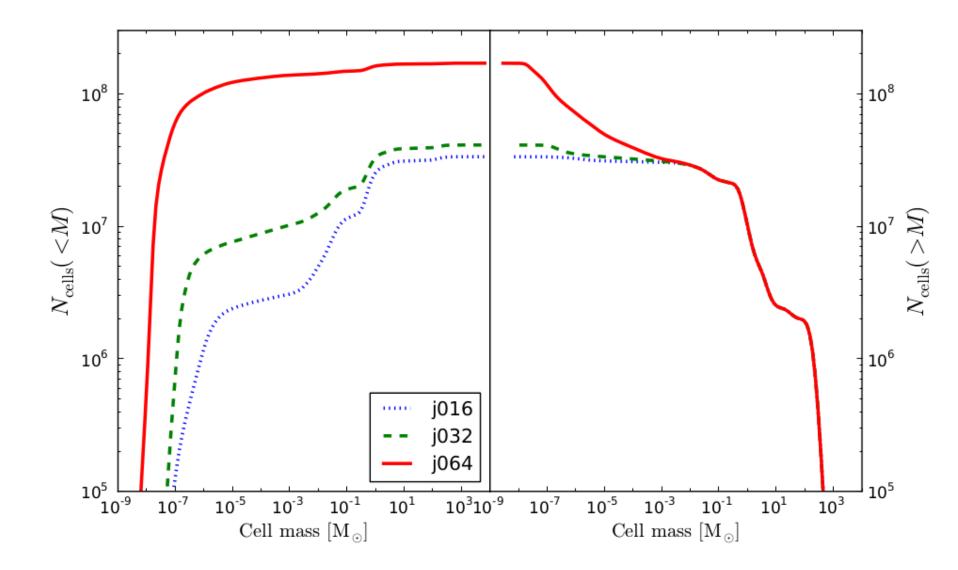


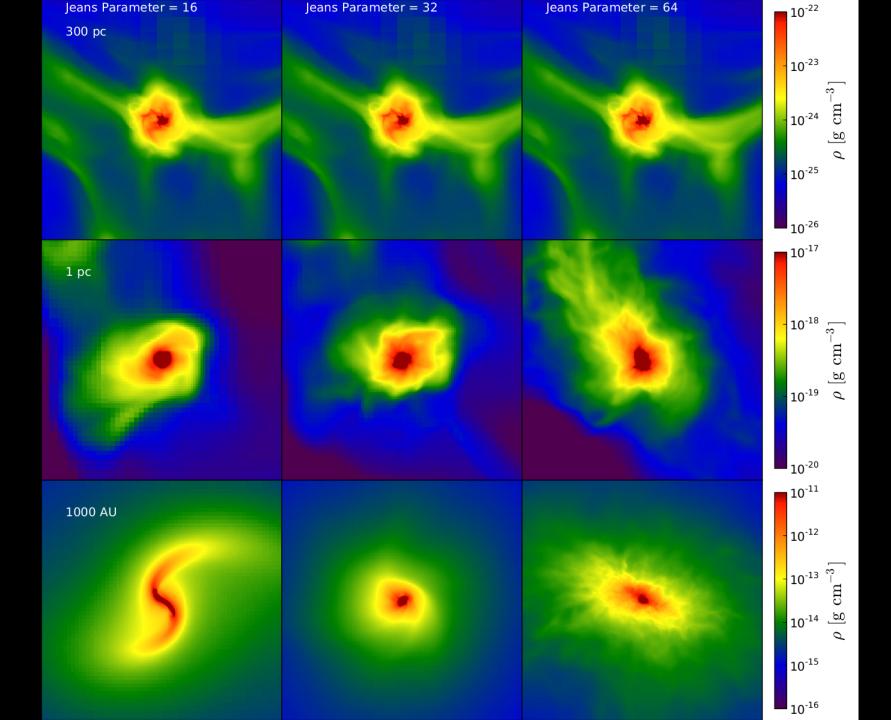
Do we see amplification in cosmology?

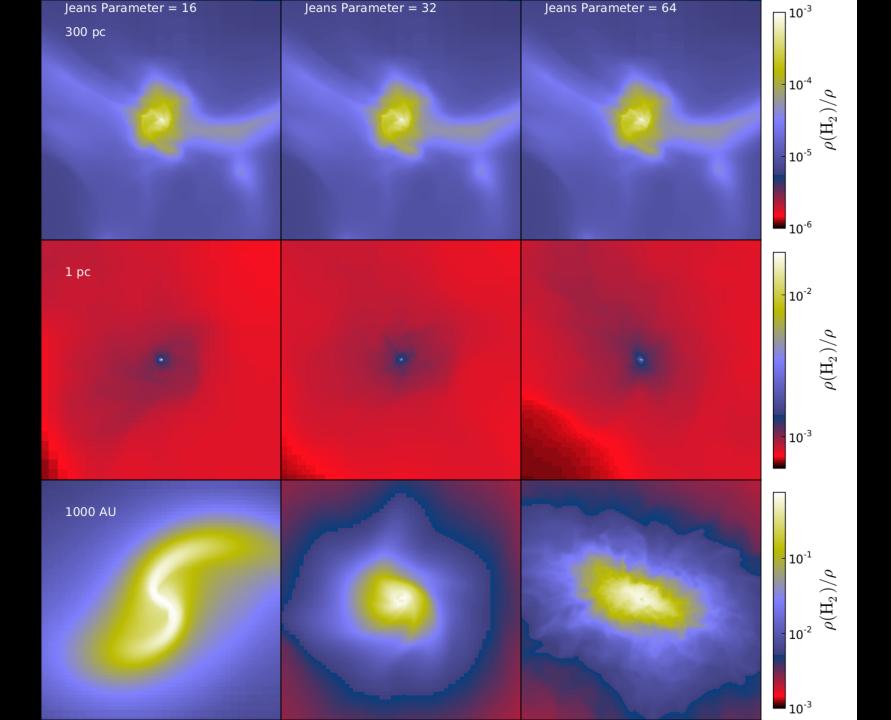
Seed field of 10^{-14} G at z = 99.

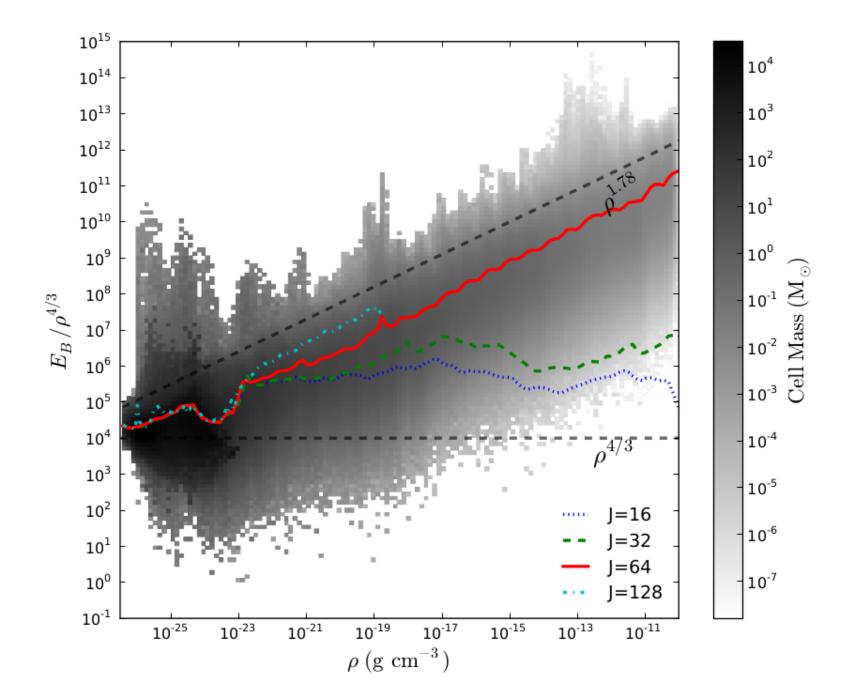
Resolve the 200 K Jeans Length by 16, 32, 64 cells.

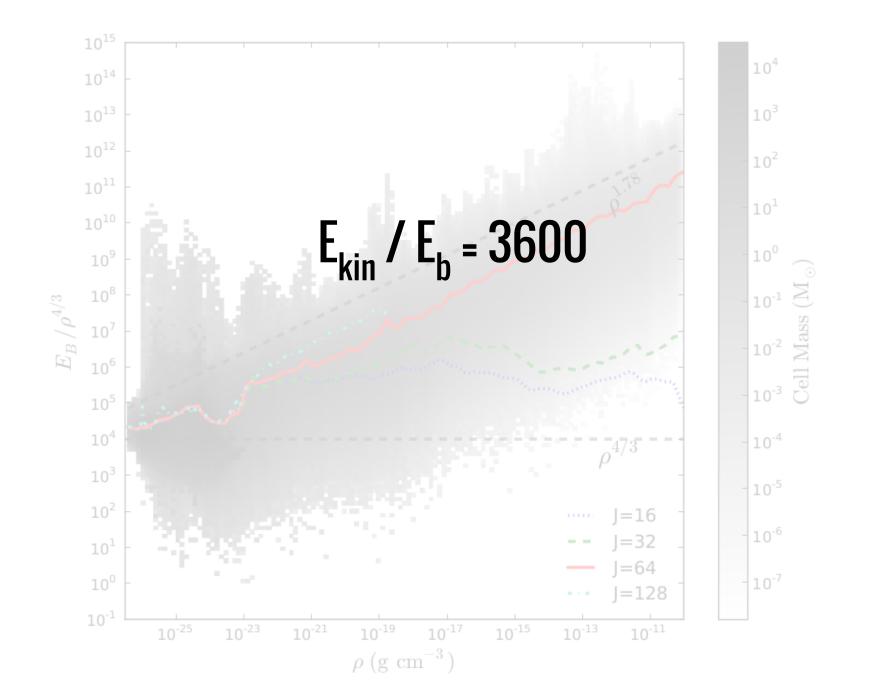
(astro-ph/1112.4479)

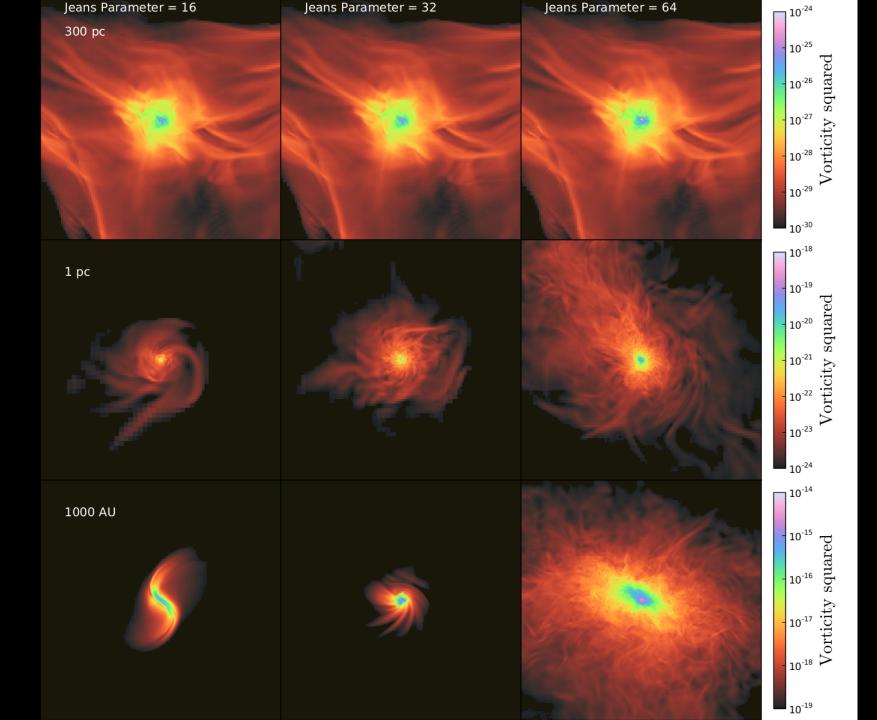


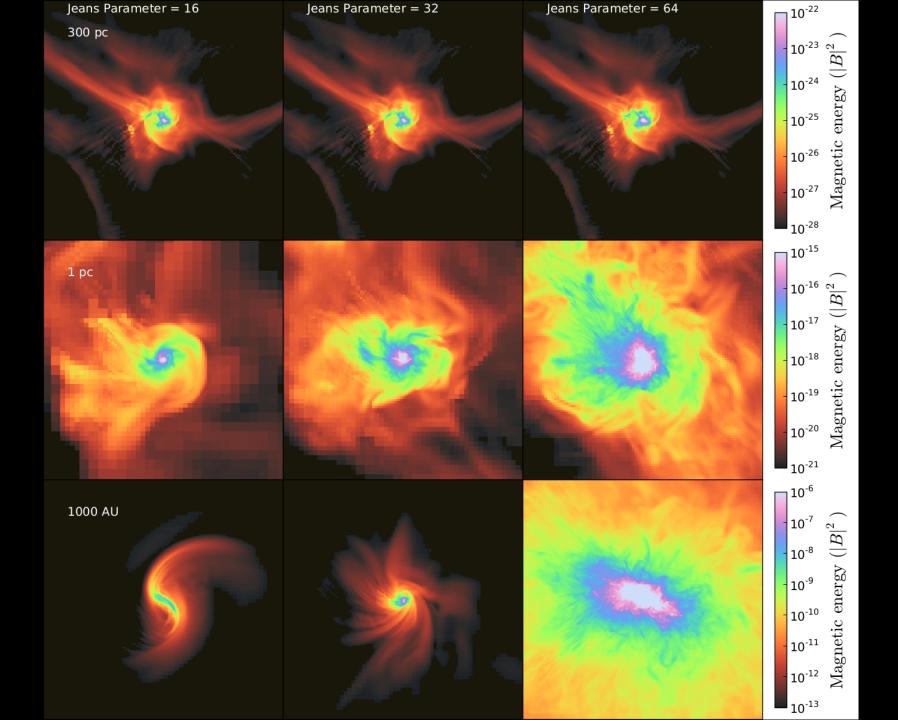


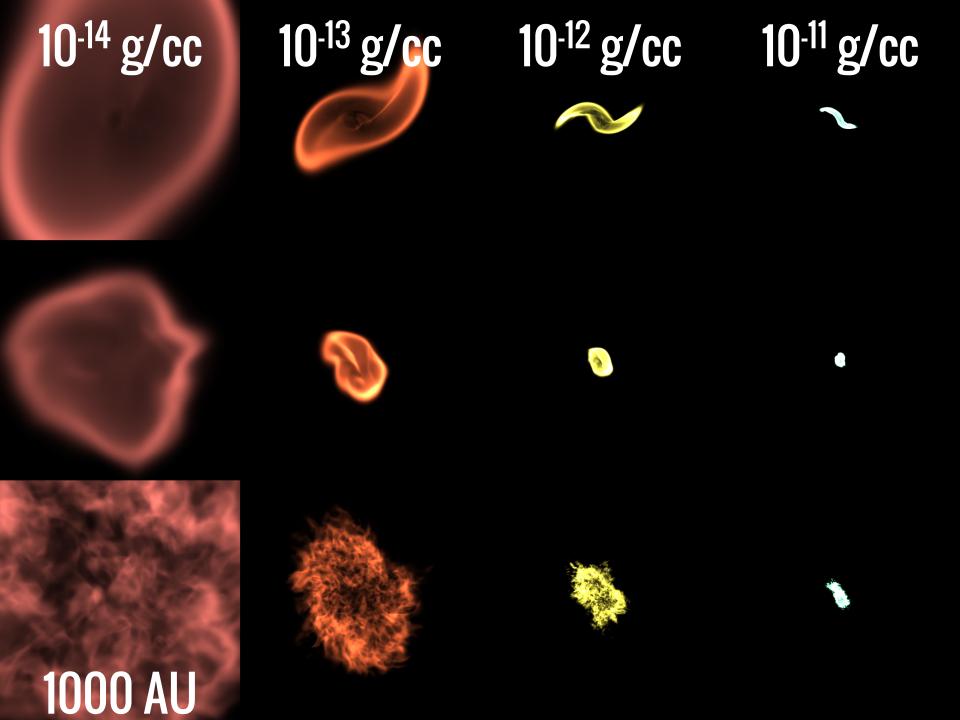
















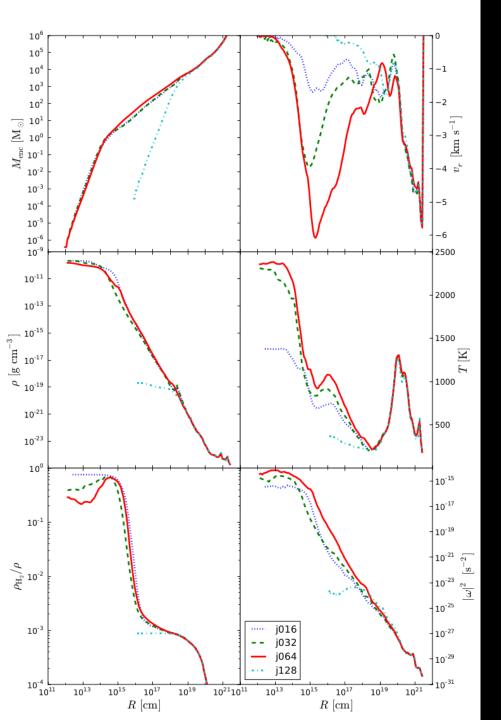




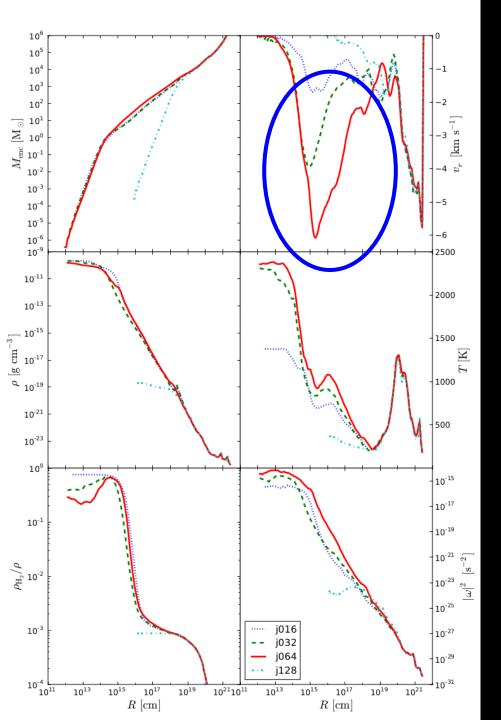




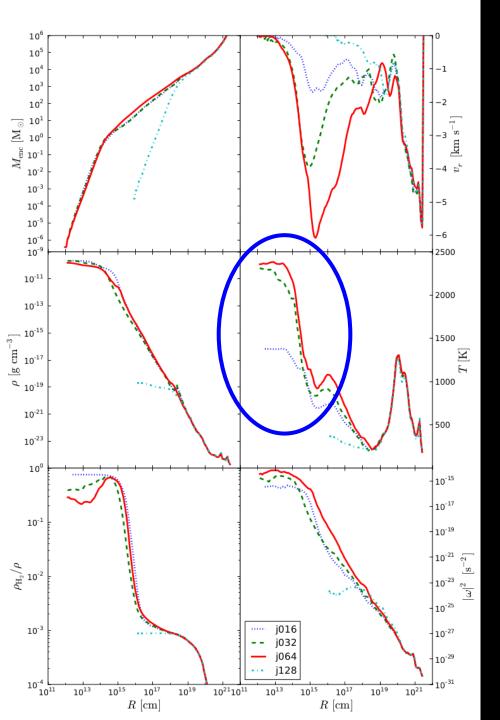
1000 AU



Three primary

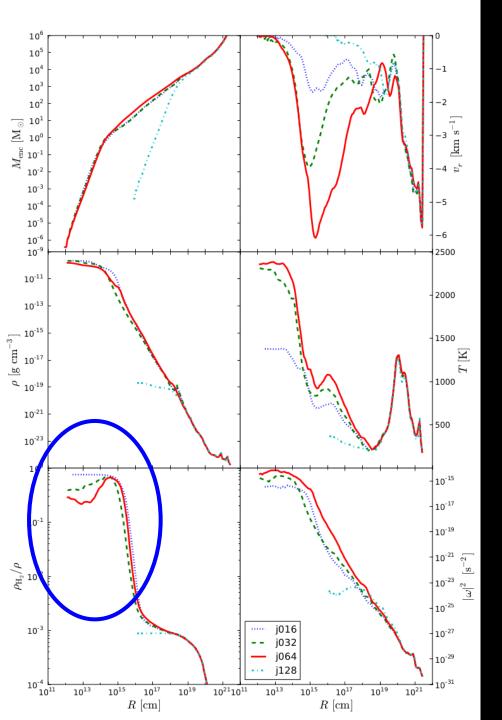


Increased inward velocity



Increased inward velocity

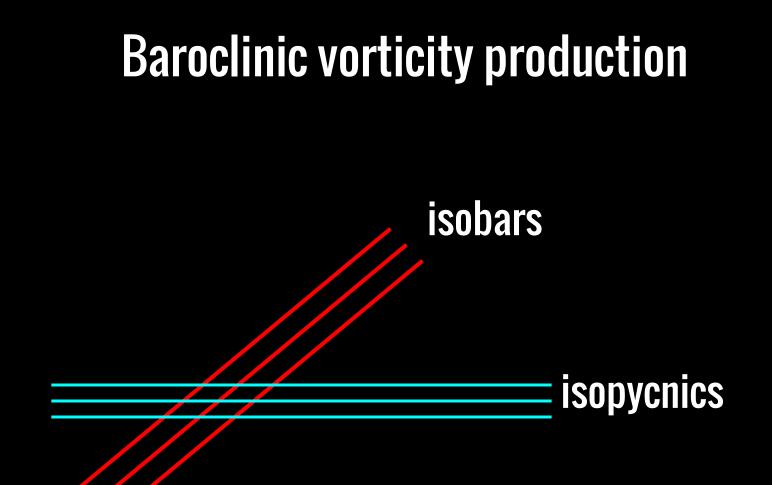
Increased core temperature



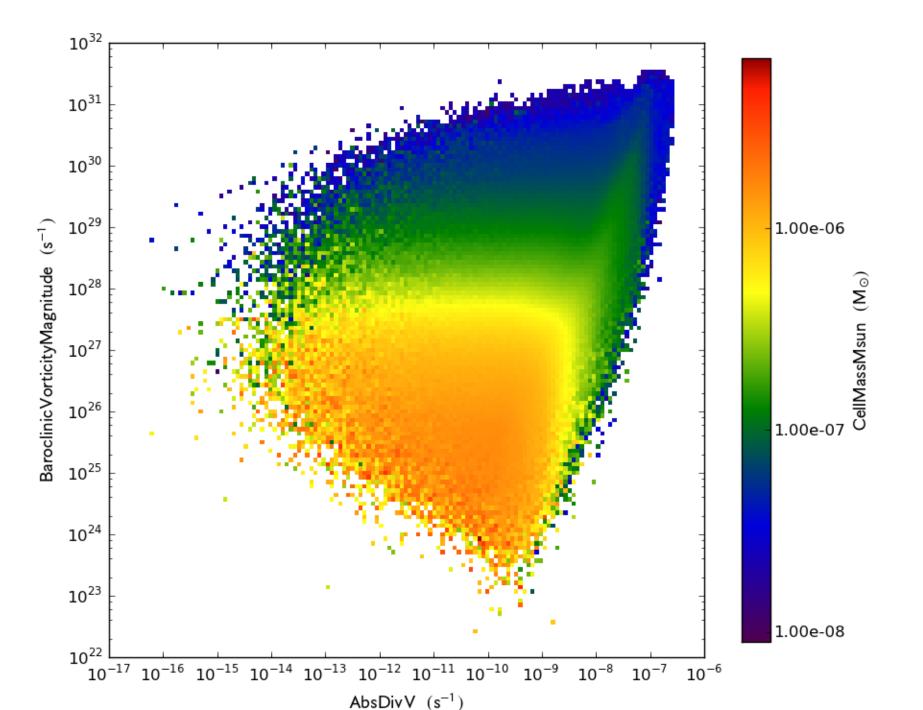
Increased inward velocity

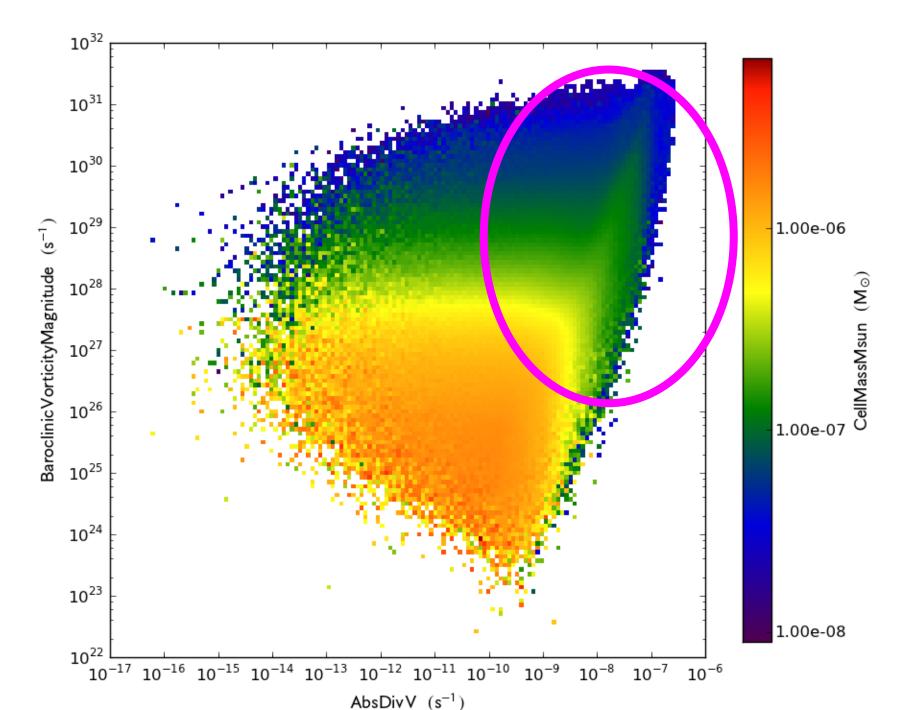
Increased core temperature

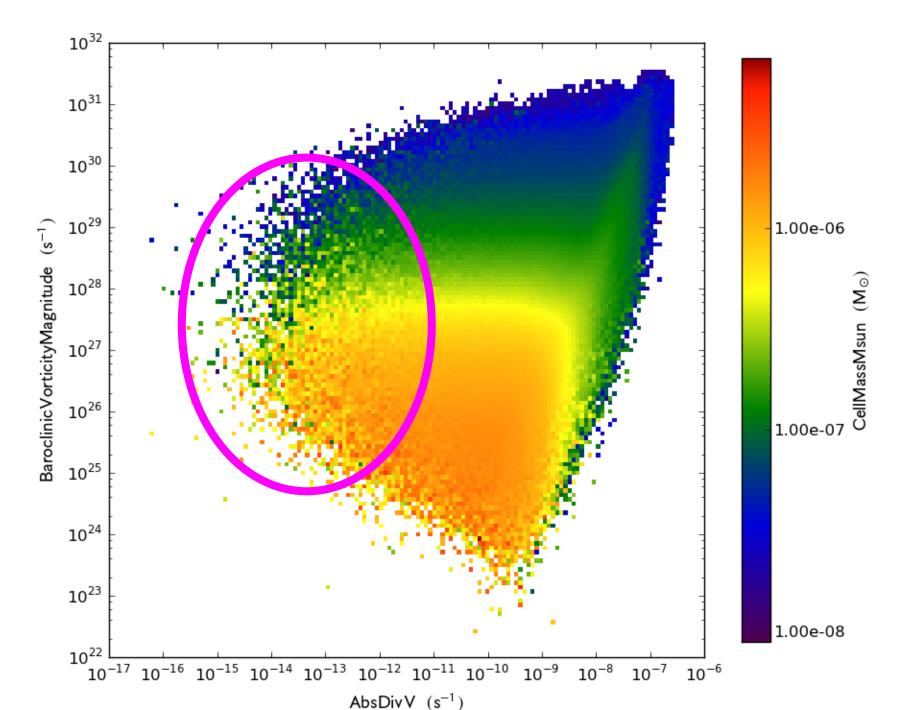
Decreased H_2 fraction



Shocks, chemothermal gradients, and convection.







Disk structure is changed, but perhaps not by the magnetic field

Dynamo action operates in cosmological initial conditions



Collaborators:

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