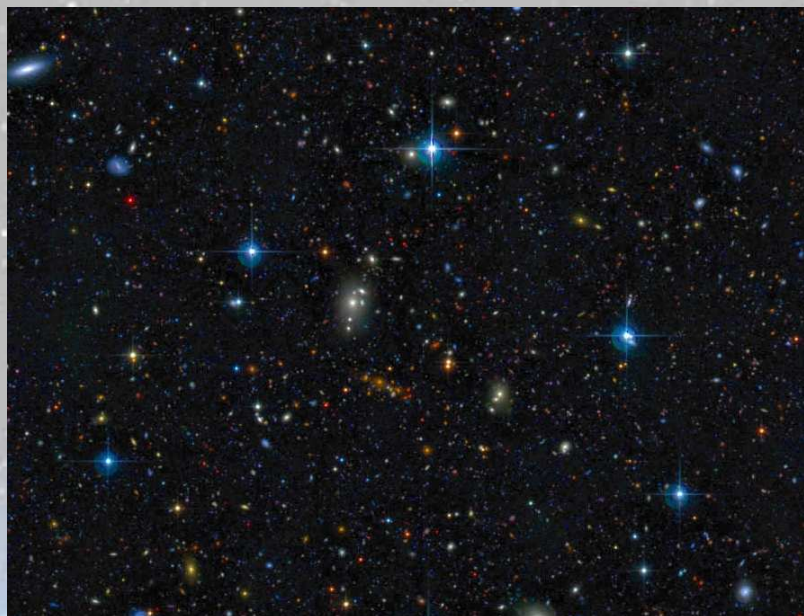


# Canada–France–Hawaii Telescope MegaCam Surveys

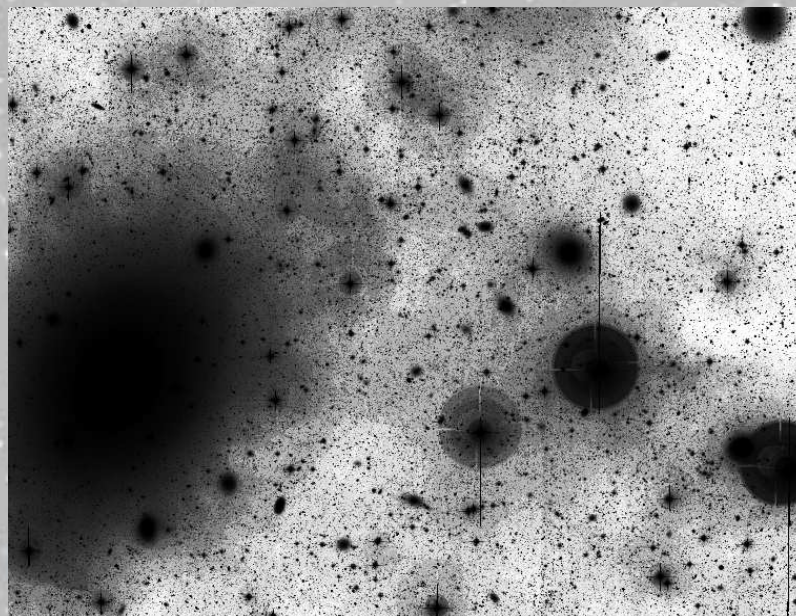
*History, strategy, and lessons learned*

*Jean–Charles Cuillandre (CFHT)*

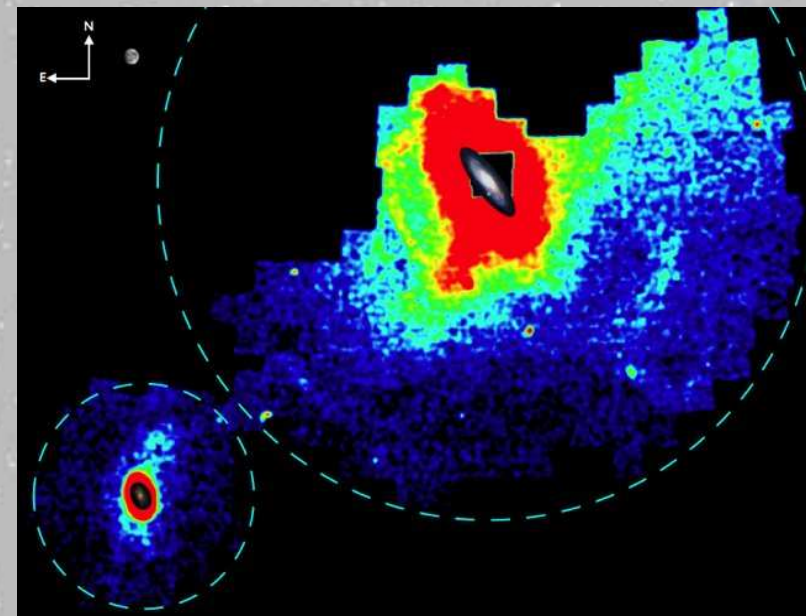
*One Degree Imager Survey Workshop – Yale – October 2009*



**Legacy Survey**  
**2003–2009**



**Virgo Survey**  
**2009–2012**



**Pan-Andromeda Survey**  
**2009–2012**



NATIONAL RESEARCH  
COUNCIL CANADA  
CONSEIL NATIONAL  
DE RECHERCHES CANADA



CENTRE NATIONAL  
DE LA RECHERCHE  
SCIENTIFIQUE



cea



CFH



# CFHT and Scientific Impact: citation metric (2006)

naturenews

Published online 6 February 2009 | Nature | doi:10.1038/news.2009.81

News

## The world's top ten telescopes revealed

The best observatories ranked by their scientific impact.

**Eric Hand**

It doesn't take a big mirror to have a big impact. The Sloan Digital Sky Survey, a project conducted with a modest 2.5-metre-wide telescope in New Mexico, performed the most highly cited science in 2006, according to a new analysis of the top ten 'high impact' astronomical observatories<sup>1</sup>.

"It measures how hot the science of the telescope is," says Juan Madrid of McMaster University in Hamilton, Canada, of the top-ten table he has released for most years since 1998. "In a way it measures how good the time-allocation committee is and how good the telescope is. I will also say it measures how good the scientists are."

Also in the top five is another modest telescope — Swift, a satellite that looks for  $\gamma$ -ray bursts — followed by three technological giants of the astronomy world: the Hubble Space Telescope, the four 8-metre telescopes of the European Southern Observatory in Paranal, Chile, and the twin 10-metre Keck telescopes in Hawaii.

The table shows that a telescope's technological advantages can push it to the top of the list, but also operating institution are important. However, some astronomers caution that citations are just one of ways to assess an observatory's value.

HIGH-IMPACT OBSERVATORIES

Rank	Facility	Citations	Participation
1	SDSS	1892	14.3%
2	Swift	1523	11.5%
3	HST	1078	8.2%
4	ESO	813	6.1%
5	Keck	572	4.3%
6	CFHT	521	3.9%
7	Spitzer	469	3.5%
8	Chandra	381	2.9%
9	Boomerang	376	2.8%
10	HESS	297	2.2%

Key  
SDSS - Sloan Digital Sky Survey  
HST - Hubble Space Telescope  
ESO - European Southern Observatory  
CFHT - Canada France Hawaii Telescope  
HESS - High Energy Stereoscopic System

Madrid, J. P. & Macchetto, D.



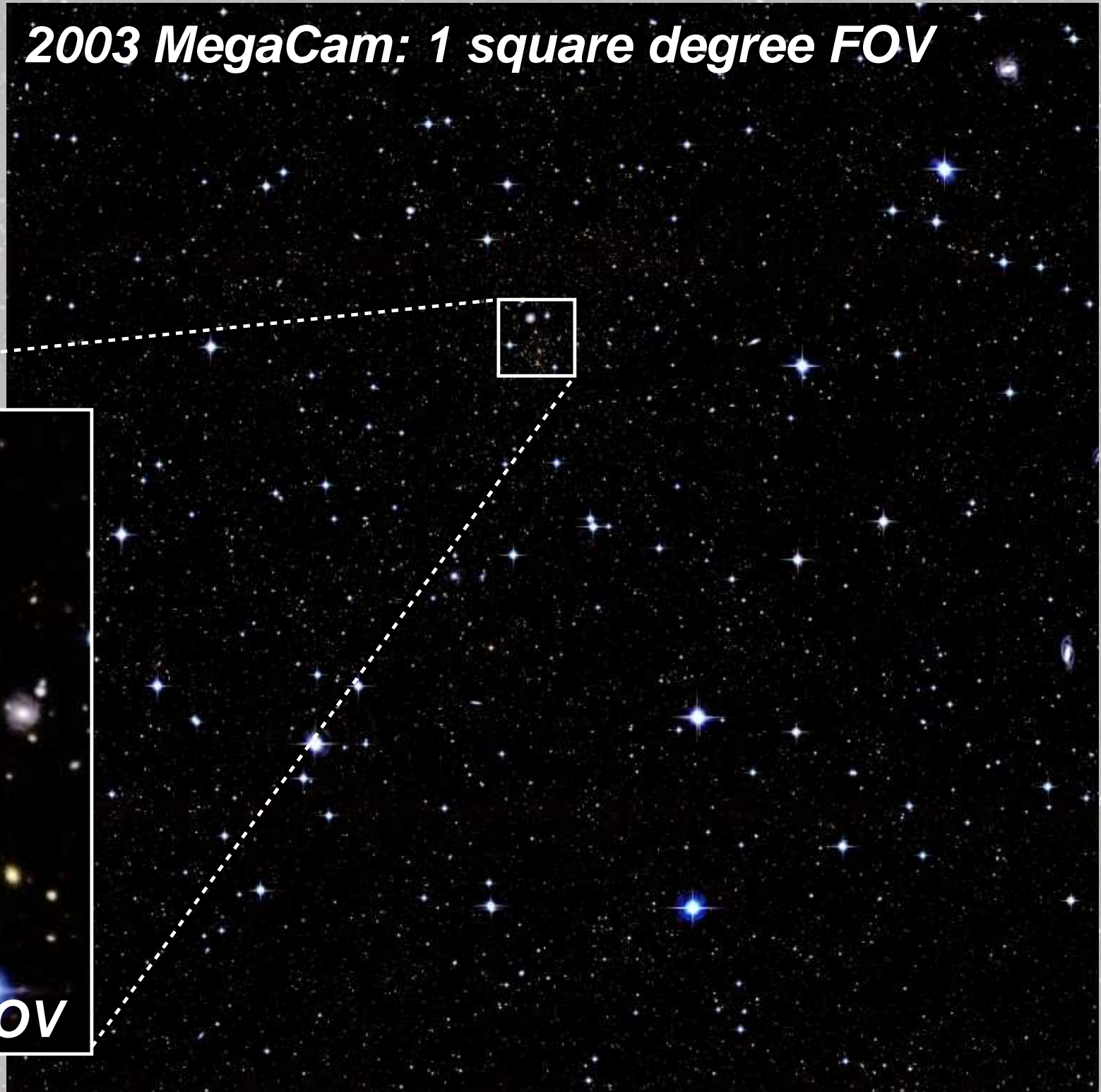
# Gravitational lensing: a key science driver for CFHT

*Discovery of the first giant arc by CFHT in 1985 (Soucail et al.)*

**2003 MegaCam: 1 square degree FOV**

**Cluster Abell 370**

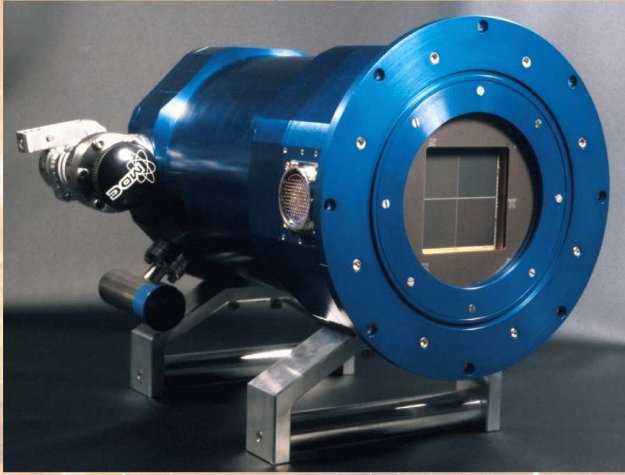
**1985 FOCAM: ~4'x4' FOV**



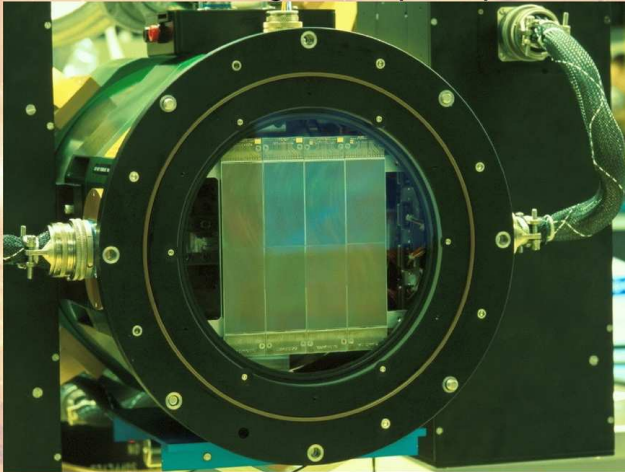


# 1983–2003: 2 decades to get CFHT full frame digital

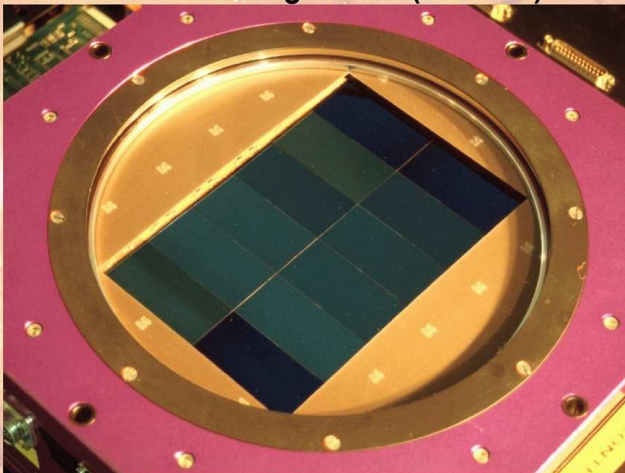
1994: 16 MegaPixels (MOCAM)



1995: 64 MegaPixels (UH8K)



1999: 100 MegaPixels (CFH12K)



2003: 350 MegaPixels (MegaCam)



New Era in Wide Field Imaging: MegaPrime at CFHT





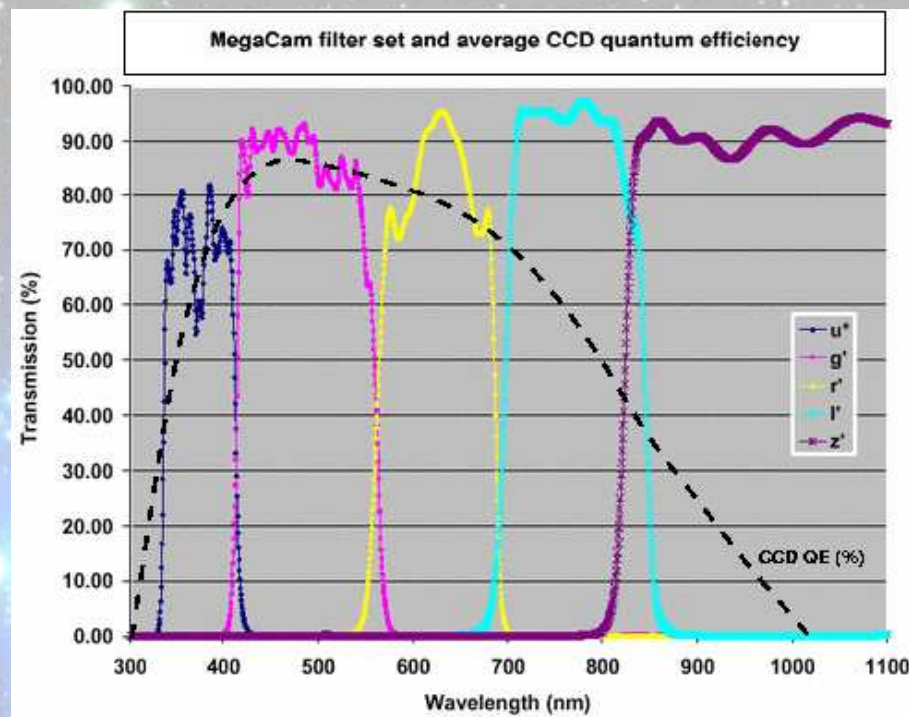
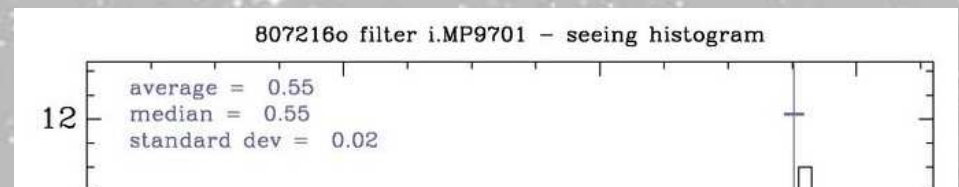
# The instrument: MegaCam on CFHT atop Mauna Kea



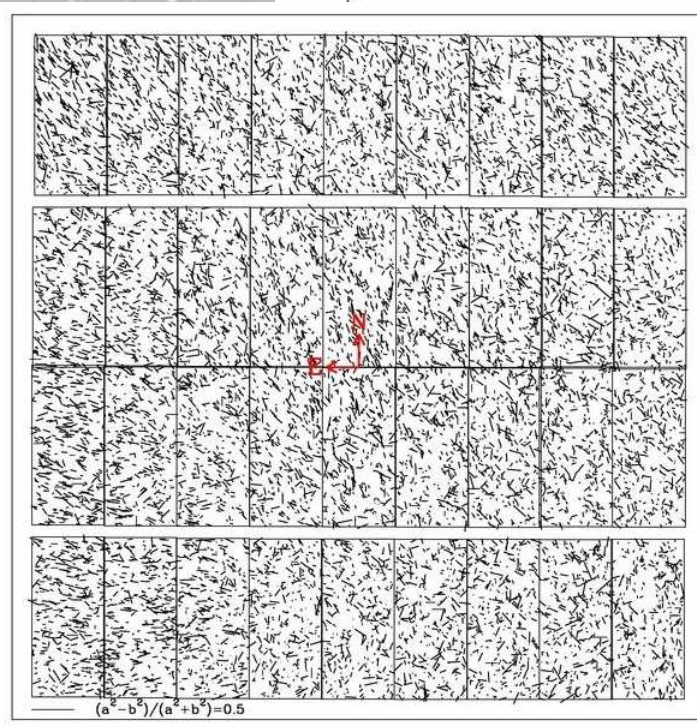
MegaPrime on CFHT (3.6m mirror)



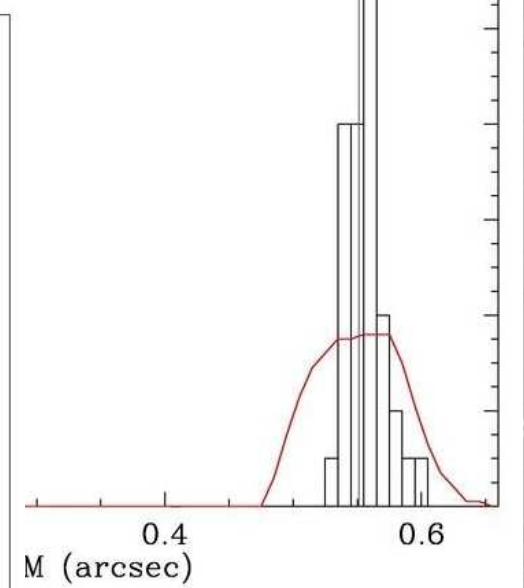
MegaCam by CEA : 340 megapixels CCD mosaic



Wavelength coverage: from u-band to z-band



Excellent image quality (uniformity) over 1 sq. deg.

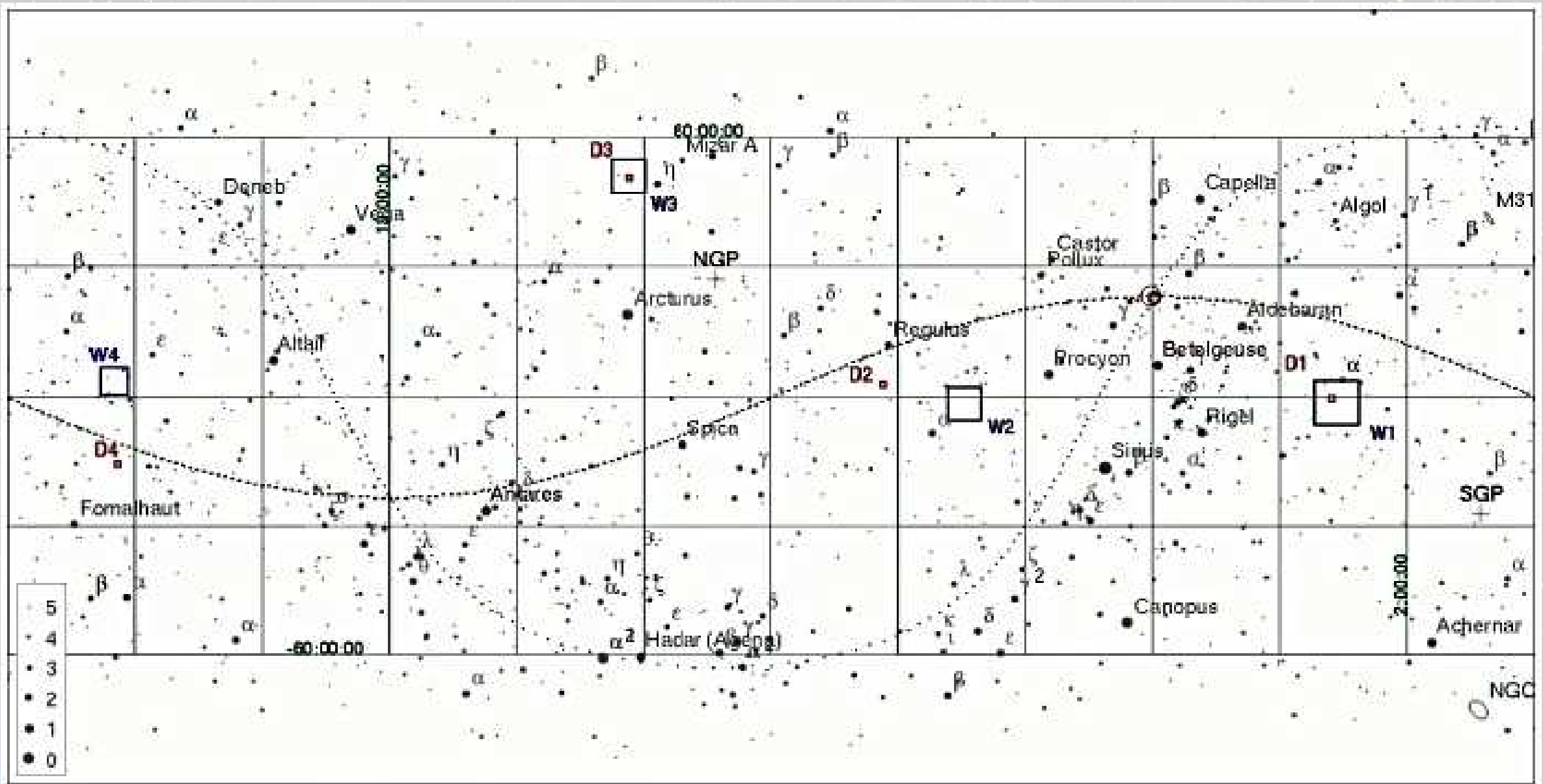




# Origins and evolution of the CFHT Legacy Survey

- **1998: CEA offered to build MegaCam and offer it to the entire CFHT scientific community on the condition a major general interest survey is undertaken with it.**
- **1999: CFH12K operational, and with MegaCam and the CFHTLS on the horizon (4 years), an evolution of the CFHT observing process is needed: start of the CFHT's New Observing Process, including Queued Service Observing (QSO).**
- **2000: with the brisk success of the CFH12K across many communities (solar system, stellar physics, nearby Universe, cosmology), resulting with a consistent telescope allocation of 55% over its 4 years lifetime, the call for ideas for the CFHTLS with MegaCam is a success. The French and Canadian agencies allocate 30% of all the telescope time to the survey (they own 85% of the time together).**
- **2001: a science definition group is created, seeking the support of external (of C+F) reviewers. The three core topics addressing the most burning astrophysical questions tailored to make an optimal use of the u to z band sensitivity, the 0.7" median seeing, the 1 sq. degree field, the 3.6m aperture, and the dark skies of Mauna Kea were:**
  - **Dark energy study using a large sample (500) of high redshift SNe type Ia (Deep survey)**
  - **Cosmic shear and cosmological constraints (Wide survey)**
  - **Origins of the Solar system: Kuiper Belt Objects (Very Wide survey)**
  - **...with MANY other science programs enabled: High-z QSOs, brown dwarfs, clustering, etc.**
- **2002: CFHTLS green-lighted (500 nights over 5 years) by the CFHT board and C+F agencies after review by the Scientific Advisory Committee (SAC). Steering Group created (9 members, PI-less), under a continuous review process by the SAC+Board.**

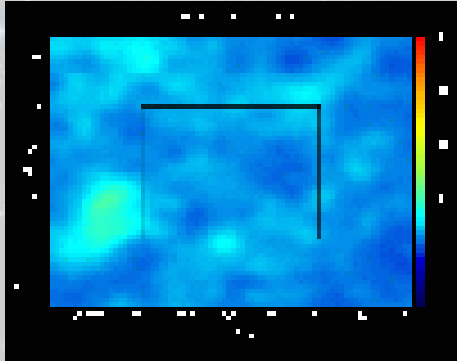
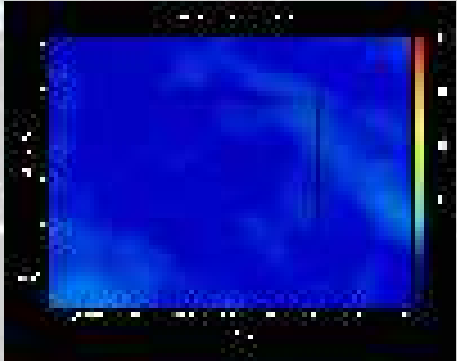
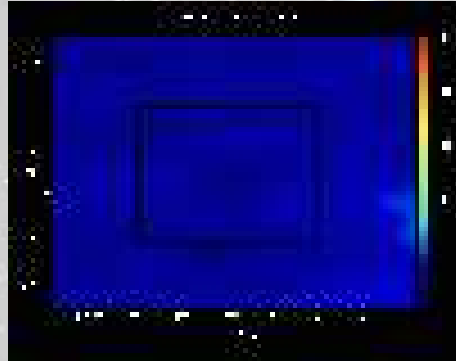
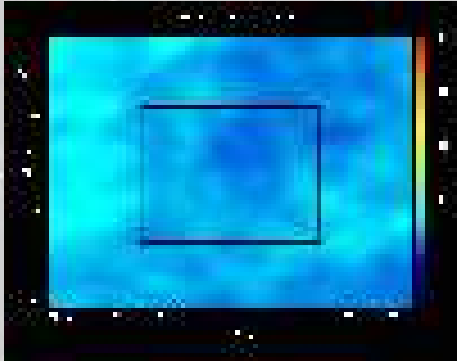
# CFHTLS completed in Feb.09, 2400 hr of valid data



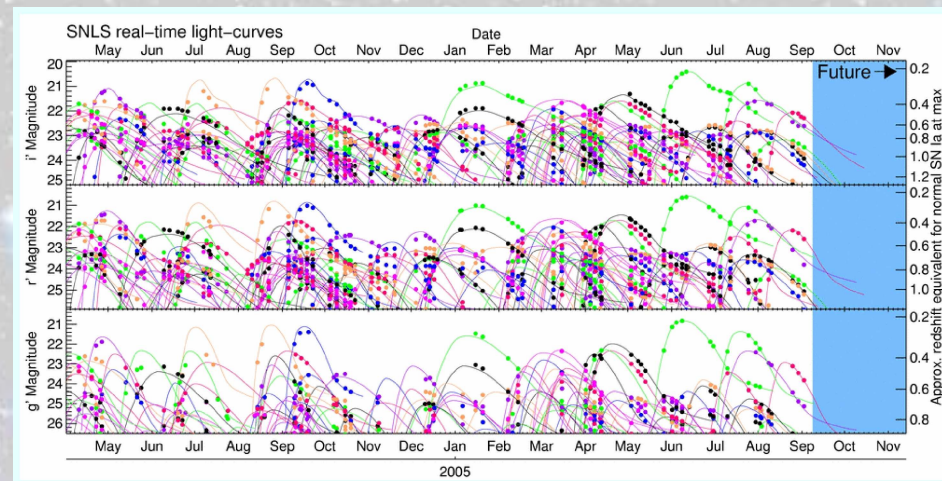
Survey	Area (sq. deg.)	Location	u*	g'	r'	i'	z'
Deep	4	D1/2/3/4	28.7	28.9	28.5	28.4	27.0
Wide	170	W1/2/3/4	26.4	26.6	25.9	25.5	24.8
Very Wide	410	On ecliptic		25.5	25.0	24.4	



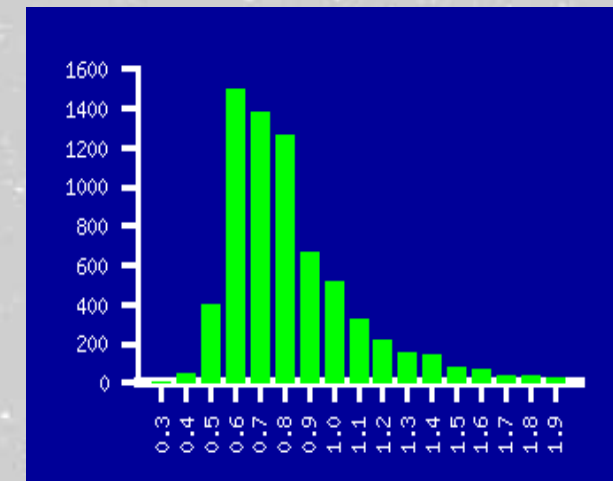
# The Deep Survey – Completed integration

Field	D1	D2	D3	D4
RA/DEC	02:26/−04:30	10:00/02:12	14:17/+52:30	22:15/−17:44
Overlap	in W1	On COSMOS field	in W3	LBQS2212−17
Dust				
Int.	289 hr.	252 hr.	285 hr.	270 hr.

Filter	Fraction	Mean IQ
u*	9%	0.91"
g'	10%	0.88"
r'	20%	0.79"
i'	40%	0.76"
z'	21%	0.75"



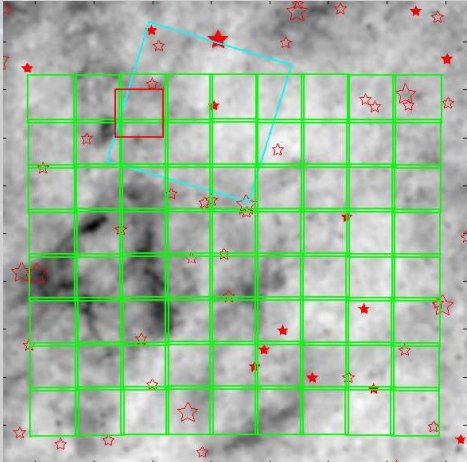
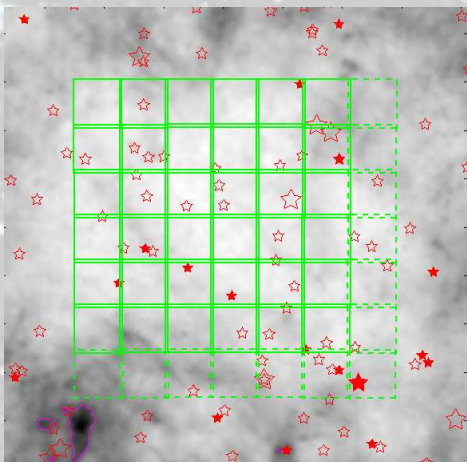
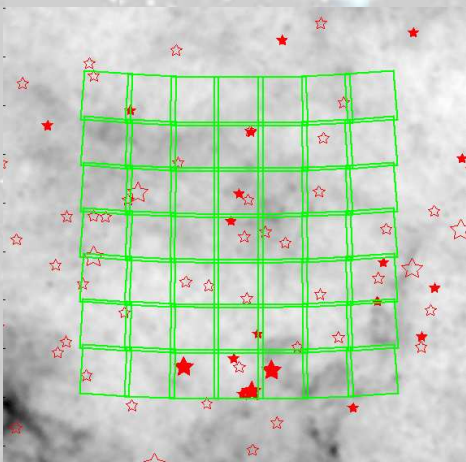
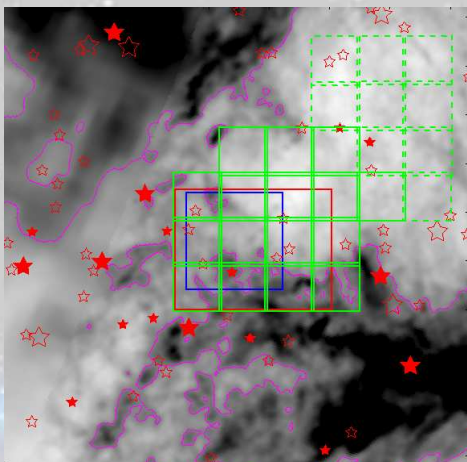
SNLS' SNe Ia light curve tracking



i' FWHM



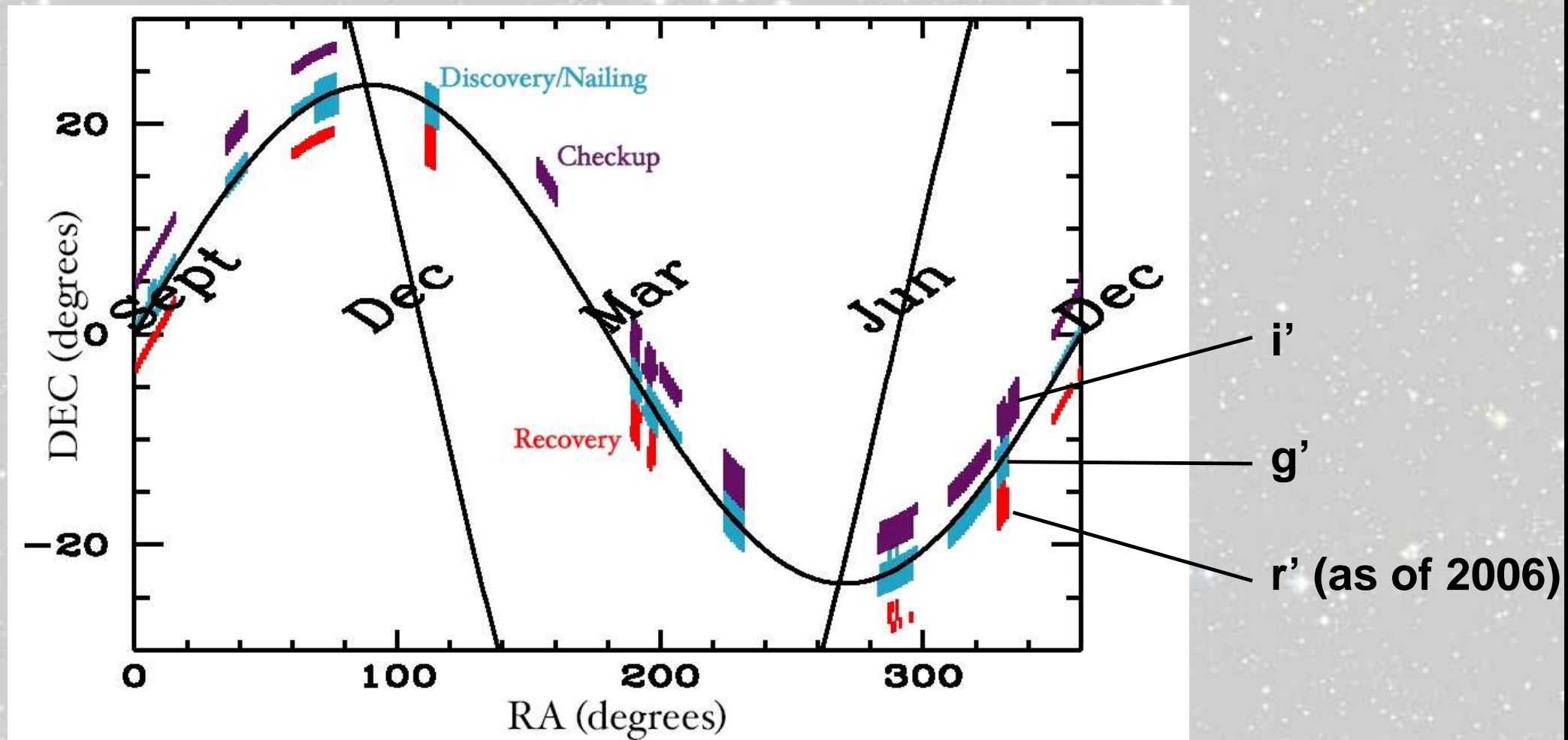
# The Wide Survey – Completed integration

<b>Field</b>	<b>W1 – 8x9 sq.deg.</b>	<b>W2 – 5x5 sq.deg.</b>	<b>W3 – 7x7 sq.deg.</b>	<b>W4 – 5x5 sq.deg.</b>
<b>RA/DEC</b>	<b>02:18/–07:00</b>	<b>08:54/–04:15</b>	<b>14:17/+54:30</b>	<b>22:13/+01:19</b>
<b>Overlap</b>	<b>XMM LSS</b>		<b>Groth Strip</b>	<b>VVDS&amp;UKIDSS</b>
<b>Dust</b>				
<b>Int.</b>	<b>361 hr.</b>	<b>139 hr.</b>	<b>242 hr.</b>	<b>126 hr.</b>

<b>Filter</b>	<b>u*</b>	<b>g'</b>	<b>r'</b>	<b>i'</b>	<b>z'</b>
<b>Fraction</b>	<b>17%</b>	<b>16%</b>	<b>17%</b>	<b>28%</b>	<b>22%</b>
<b>Mean IQ</b>	<b>0.90"</b>	<b>0.88"</b>	<b>0.76"</b>	<b>0.67"</b>	<b>0.72"</b>



# The Very Wide Survey – Completed integration



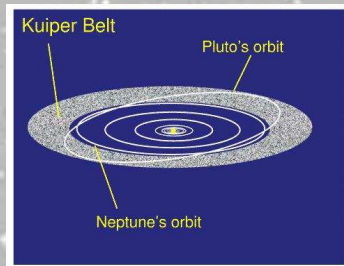
Filter	g'	r'	i'
Integration	47 hr.	75 hr.	70 hr.
Fraction	24%	40%	36%
Mean IQ	0.90"	0.80"	0.80"

**Sky Coverage: 410 sq. deg. ...**



# Science with the CFHT Legacy Survey

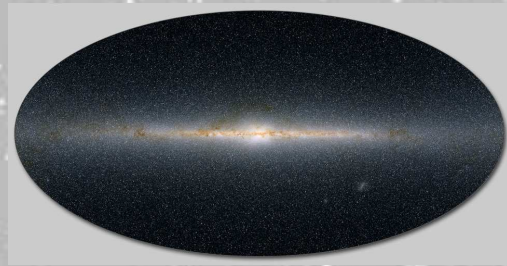
## Solar System



*Very Wide*

- The Kuiper Belt
- Asteroids

## The Galaxy



*Very Wide / Wide / Deep*

- Stellar Populations
- Brown Dwarfs
- Low Mass Stars
- White Dwarfs
- Dynamics
- Variability

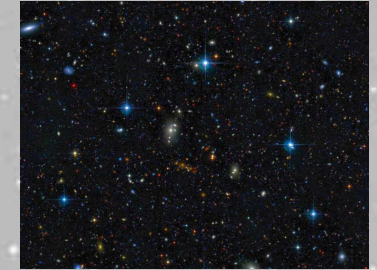
## Galaxies & Clusters



*Wide / Deep*

- Redshift Distribution
- Evolution
- Clusters
- Morphology
- Clustering
- Weak Lensing
- Star Formation
- Luminosity Function
- Environment
- AGNs

## Cosmology



*Very Wide / Wide / Deep*

- Dark Energy
- Cosmic Shear
- Strong Lensing
- Large Scale Struct.
- Supernovae
- GRBs
- QSOs



# The 5 most cited CFHTLS publications

Article, citations, CFHTLS component, and title (as of June 2009):

Astier et al. 2006	987	SNLS	The Supernova Legacy Survey: Measurement of $\Omega_M$ , $\Omega_\Lambda$ and $w$ from the First Year Data Set
Hoekstra et al. 2006	128	Wide	First cosmic shear results from the Canada–France–Hawaii Telescope Wide Synoptic Legacy Survey
Ilbert et al. 2006	121	Deep	Accurate photometric redshifts for the CFHT Legacy Survey calibrated using the VIMOS VLT Deep Survey
Sullivan et al. 2006	109	SNLS	Rates and Properties of Type Ia Supernovae as a Function of Mass and Star Formation in Their Host Galaxies
Semboloni et al. 2006	74	Deep	Cosmic Shear Analysis with CFHTLS Deep data

**TOP**  
**5**

Yearly publication rate:

2005: 4

2006: 19

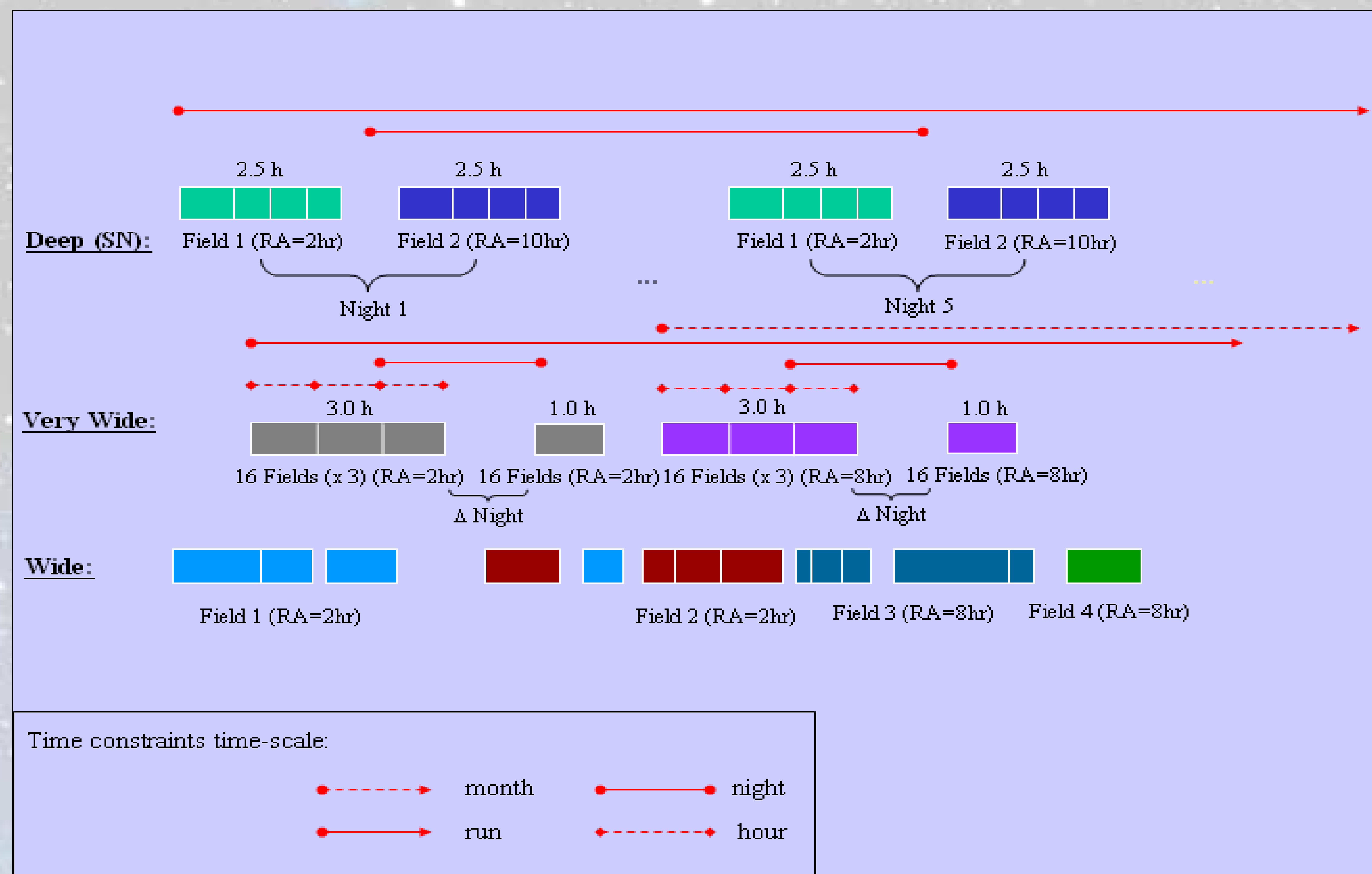
2007: 25

2008: 19



# The CFHTLS seen by Queued Service Observing

- ✓ *Time critical observations on a large-scale*  
*SNe fields every 4 nights, 2 fields per run*  
*Each Very Wide patch has 5 time constraints*
- ✓ *Balancing Agency Time*  
*PI programs (50%) versus CFHTLS (50% or "49" nights at best per semester)*
- ✓ *Balancing time share between CFHTLS programs*  
*Deep (44%), Wide (34%), Very Wide (22%)*
- ✓ *Example of conflicts just with the CFHTLS:*



- ✓ *Time constraints are a real challenge, require planning, understanding of programs and priorities, and constant attention to the general balance.*

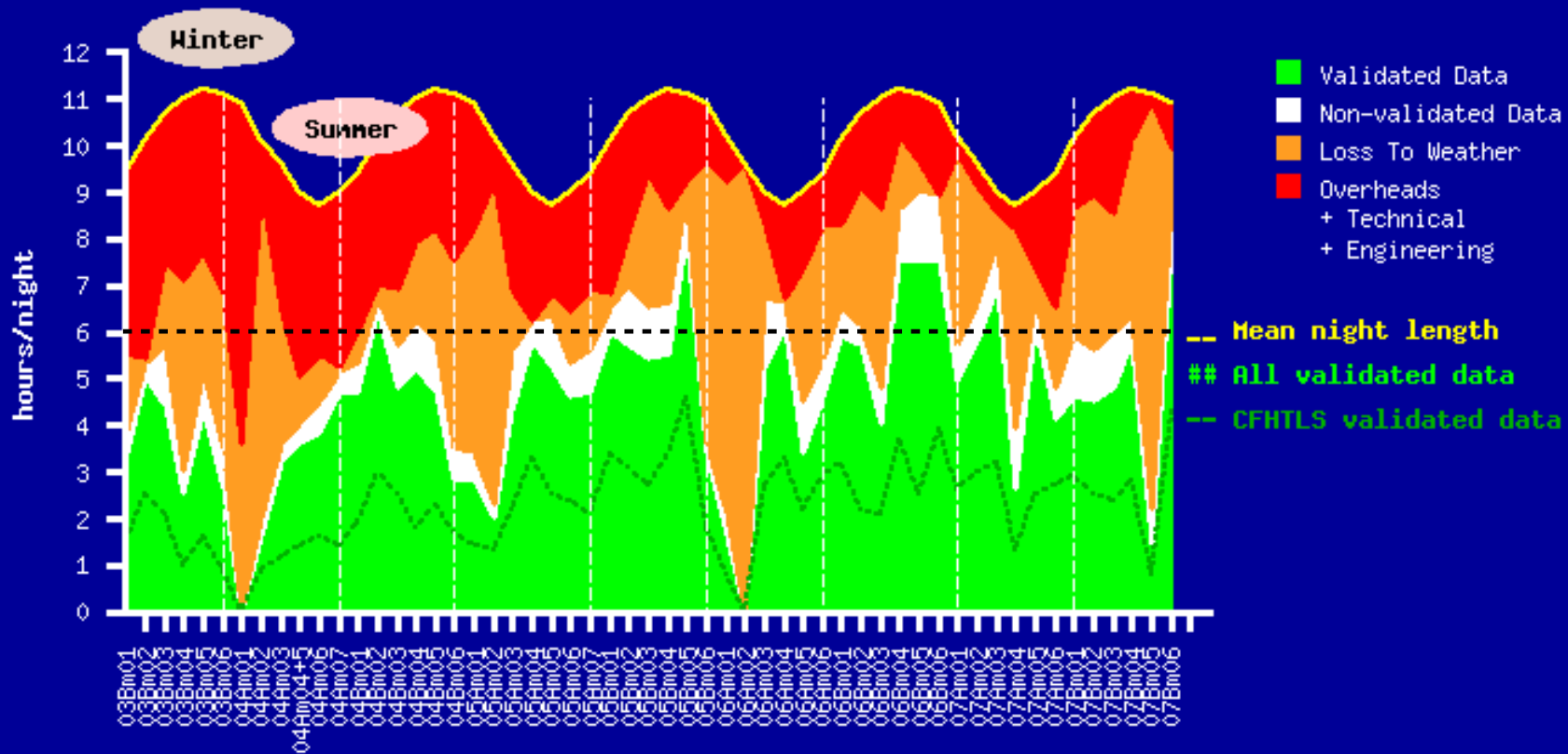


# The real world: weather and technical downtime

Period	Runs	Nights	Validation	CFHTLS	Unvalidated	Weather	Overheads	Night Length
All	55	16.5	4.5	50%	0.8	2.5	28%	10.1
03B	6	16.8	3.7	43%	0.7	2.3	48%	10.6
04A	6	18.7	2.8	39%	0.4	2.6	36%	9.5
04B	6	19.0	4.8	46%	0.8	1.7	38%	10.6
05A	7	16.4	4.2	52%	0.8	2.3	32%	9.5
05B	6	16.3	5.7	56%	0.8	2.1	26%	10.8
06A	6	15.8	3.5	57%	0.7	4.0	18%	9.3
06B	6	16.8	6.4	46%	1.0	1.8	20%	10.8
07A	6	14.3	5.1	50%	0.8	2.3	15%	9.3
07B	6	14.2	4.9	53%	1.0	3.7	18%	10.8

## CFHT's MegaPrime Observing Statistics

Starting June 2003, with 6 to 7 runs per semester since (m01->m06/7)  
 "A" semesters = February to July / "B" semesters = August to January  
 The night length varies from 8.7 hr (June) to 11.2 hr (December)



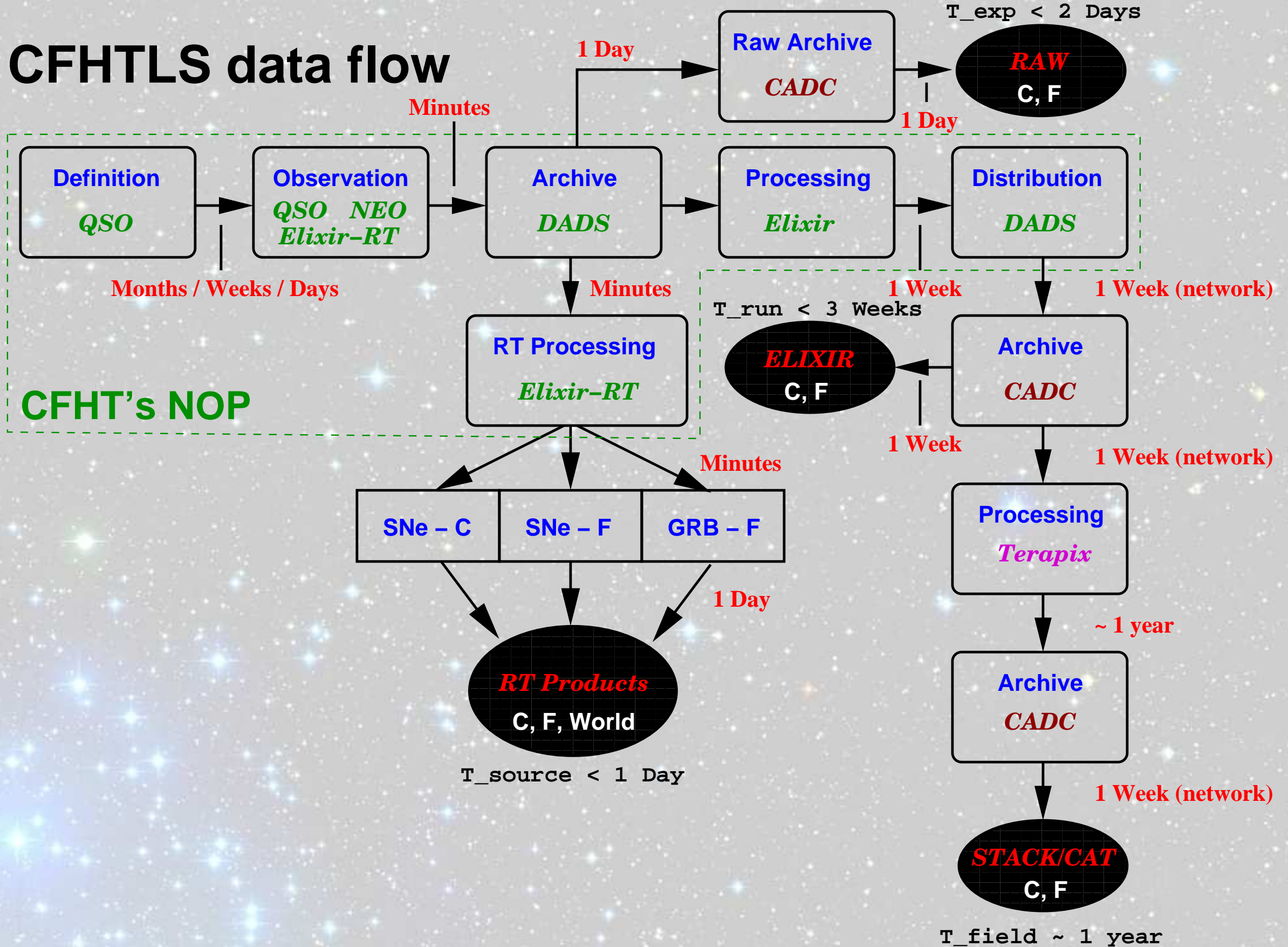
## Mid-term review advanced to May 2005

- Priority to core science (SNLS, Cosmic Shear)
  - Wide cut to u/2 & z/2
  - Very Wide ramped down
- ... and addition of the W4 field  
 ... plus a 6 months extension with top priority to get the core CFHTLS completed.

**Flexibility = Salvation**



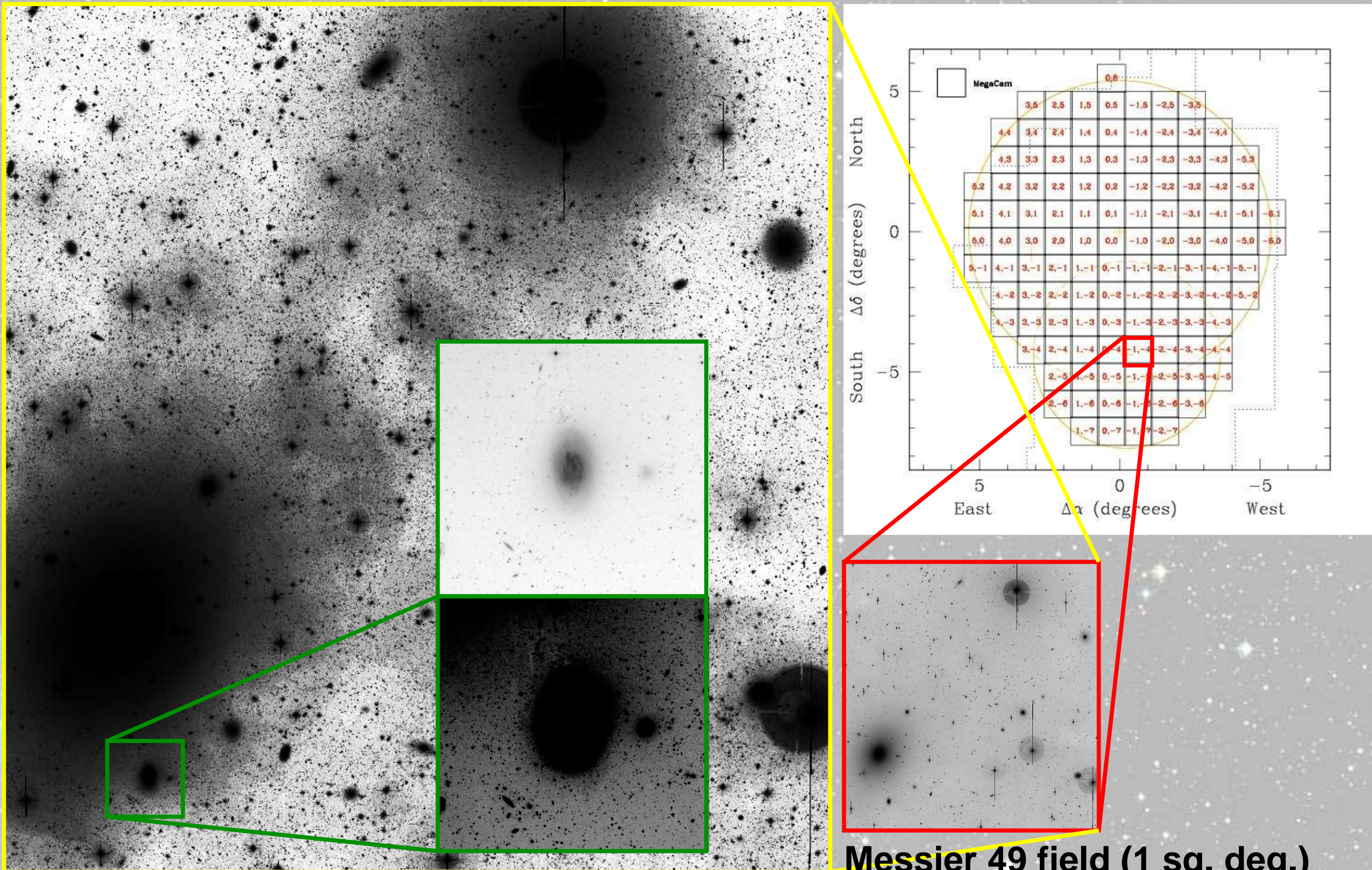
# CFHTLS data flow





# The Next Generation Virgo Survey

700 hours over 3 years (2009–12) to cover 104 sq. deg. in u\*g'r'i'z'

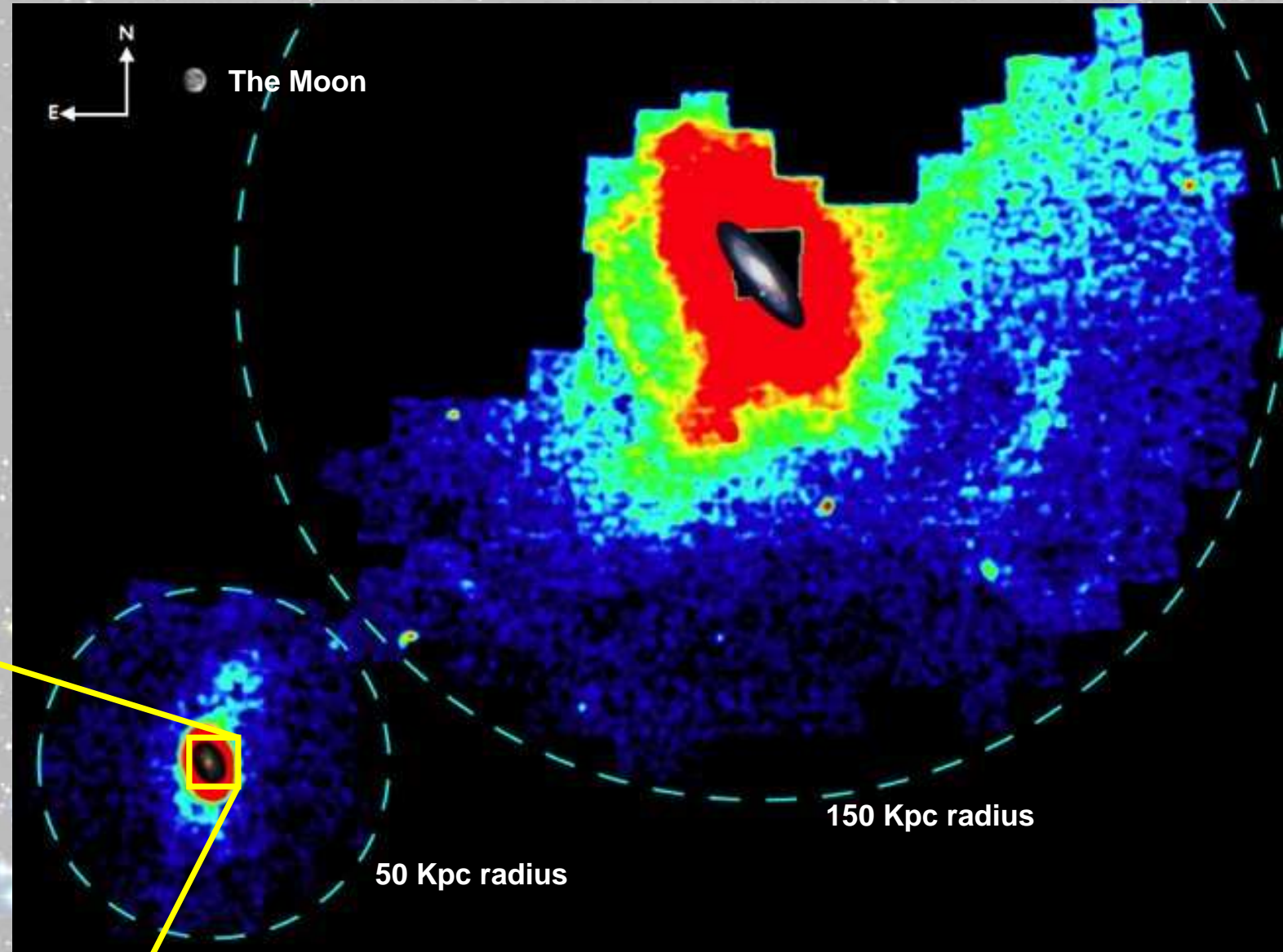
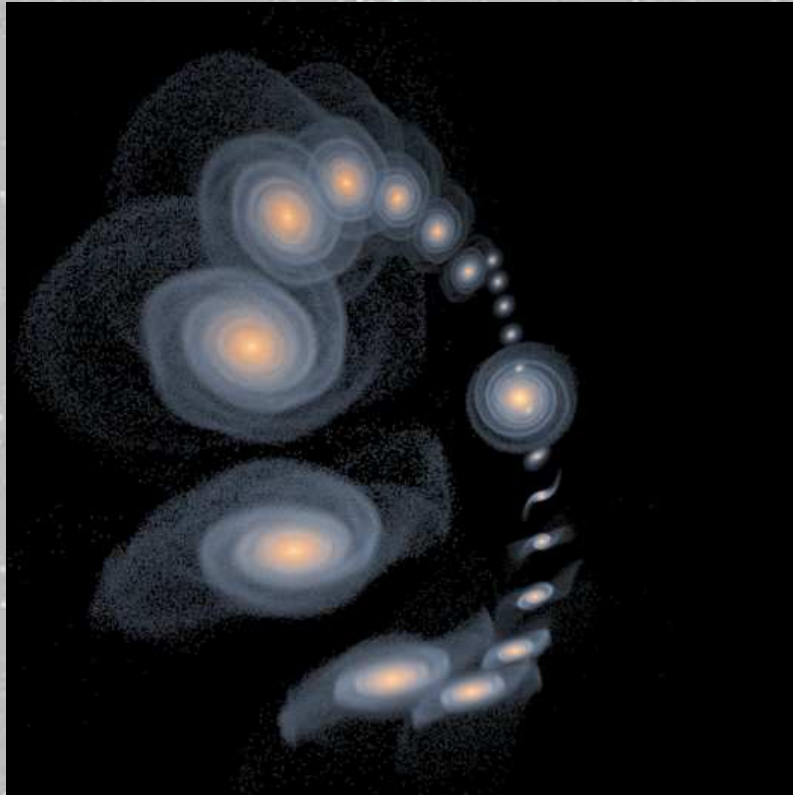


Messier 49 field (1 sq. deg.)



# The Pan-Andromeda Survey

450 hours over 5 years (2007–12) to cover 350 sq. deg. in g' and i'



Depth:  $g=25.5$   $i=24.5$

Detects surface brightness  $\sim 32$  mag/sq.arcsec

Messier 33 field (1 sq. deg.)