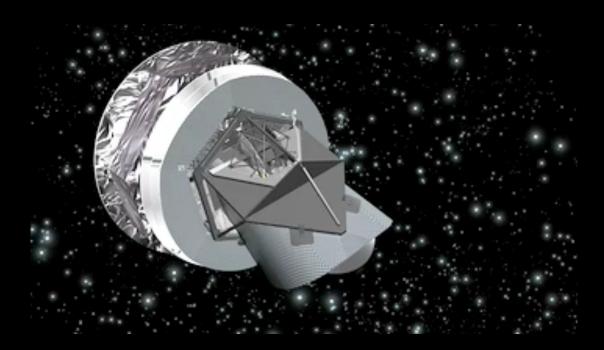
## Astro 220: Galaxies and cosmology Nov 10 2010



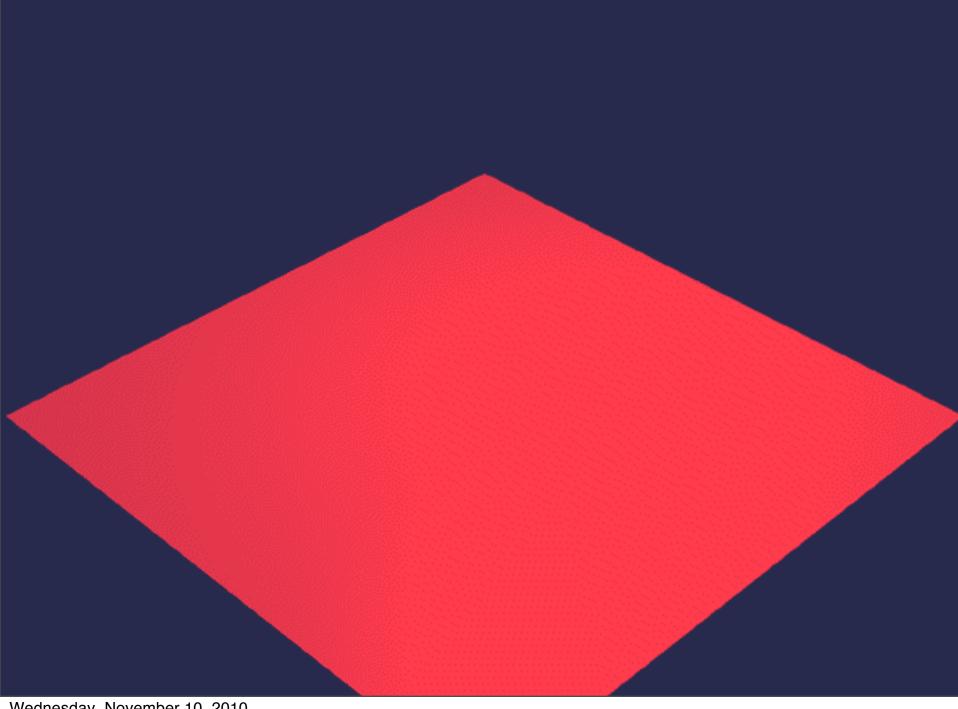
Description	Symbol	Value	+ uncertainty	<ul> <li>uncertainty</li> </ul>
Total density	$\Omega_{tot}$	1.02	0.02	0.02
Equation of state of quintessence	w	< -0.78	95% CL	—
Dark energy density	$\Omega_{\Lambda}$	0.73	0.04	0.04
Baryon density	$\Omega_b h^2$	0.0224	0.0009	0.0009
Baryon density	$\Omega_b$	0.044	0.004	0.004
Baryon density (cm <sup>-3</sup> )	$n_b$	$2.5 \times 10^{-7}$	$0.1 \times 10^{-7}$	$0.1 \times 10^{-7}$
Matter density	$\Omega_m h^2$	0.135	0.008	0.009
Matter density	$\Omega_m$	0.27	0.04	0.04
Light neutrino density	$\Omega_{\nu}h^2$	< 0.0076	95% CL	_
CMB temperature (K) <sup>a</sup>	$T_{\rm emb}$	2.725	0.002	0.002
CMB photon density (cm <sup>-3</sup> ) <sup>b</sup>	ny	410.4	0.9	0.9
Baryon-to-photon ratio	η	$6.1\times10^{-10}$	$0.3\times10^{-10}$	$0.2 \times 10^{-10}$
Baryon-to-matter ratio	$\Omega_b \Omega_m^{-1}$	0.17	0.01	0.01
Fluctuation amplitude in $8h^{-1}$ Mpc spheres	$\sigma_8$	0.84	0.04	0.04
Low- $z$ cluster abundance scaling	$\sigma_8 \Omega_m^{0.5}$	0.44	0.04	0.05
Power spectrum normalization (at $k_0 = 0.05 \text{ Mpc}^{-1})^c$	A	0.833	0.086	0.083
Scalar spectral index (at $k_0 = 0.05 \text{ Mpc}^{-1})^c$	$n_s$	0.93	0.03	0.03
Running index slope (at $k_0 = 0.05 \text{ Mpc}^{-1})^c$	$dn_s/d\ln k$	-0.031	0.016	0.018
Tensor-to-scalar ratio (at $k_0 = 0.002 \text{ Mpc}^{-1}$ )	r	< 0.90	95% CL	_
Redshift of decoupling	$z_{dec}$	1089	1	1
Thickness of decoupling (FWHM)	$\Delta z_{dec}$	195	2	2
Hubble constant	h	0.71	0.04	0.03
Age of universe (Gyr)	$t_0$	13.7	0.2	0.2
Age at decoupling (kyr)	$t_{dec}$	379	8	7
Age at reionization (Myr, 95% CL))	$t_r$	180	220	80
Decoupling time interval (kyr)	$\Delta t_{dec}$	118	3	2
Redshift of matter-energy equality	$z_{eq}$	3233	194	210
Reionization optical depth	τ	0.17	0.04	0.04
Redshift of reionization (95% CL)	$z_r$	20	10	9
Sound horizon at decoupling (°)	$\theta_A$	0.598	0.002	0.002
Angular size distance to decoupling (Gpc)	dA	14.0	0.2	0.3
Acoustic scale <sup>d</sup>	$\ell_A$	301	1	1
Sound horizon at decoupling (Mpc) <sup>d</sup>	Ts.	147	2	2

## Weak Anthropic Principle:

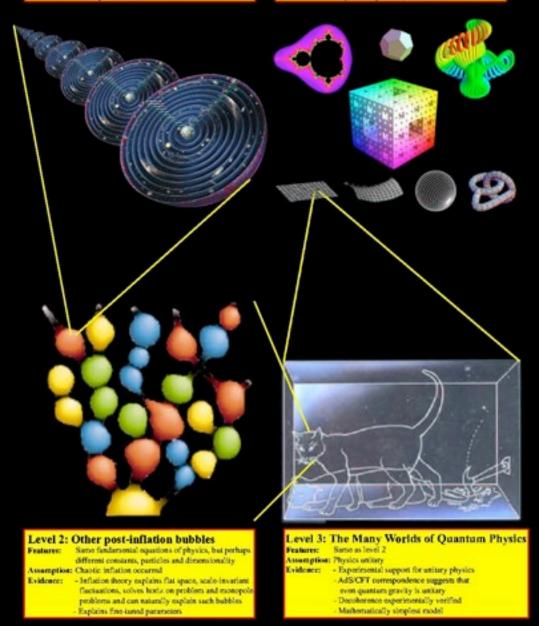
"The observed values of all physical and cosmological quantities are not equally probable but they take on values restricted by the requirement that there exist sites where carbon-based life can evolve and by the requirement that the Universe be old enough for it to have already done so."

## Strong Anthropic Principle:

"The universe must have those properties which allow life to develop within it at some stage in its history."



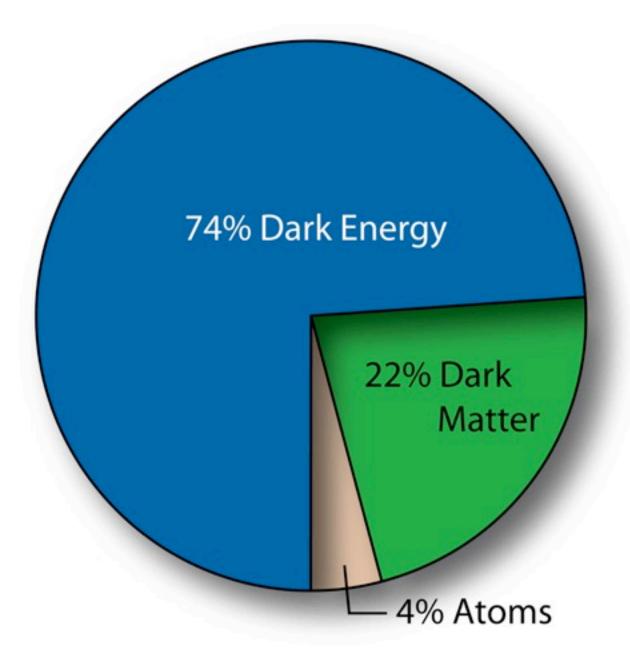
Level 1: Regions beyond our cosmic horizon Features: Same laws of physics, different initial conditions Assumptions: Infinite space, ergodic matter distribution Evidence: • Microwave background measurements point to flat, infinite space, large-scale smoothness • Simplest model Level 4: Other mathematical structures Feature: Different fundamental equations of physics Assumption: Mathematical existence – physical existence Evidence: - Unreasonable effectiveness of math in physics - Answers Wheeler/Flawking question: "why these equations, not others"

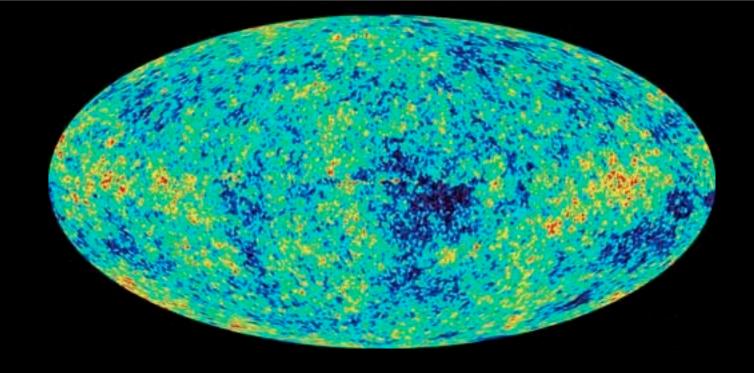


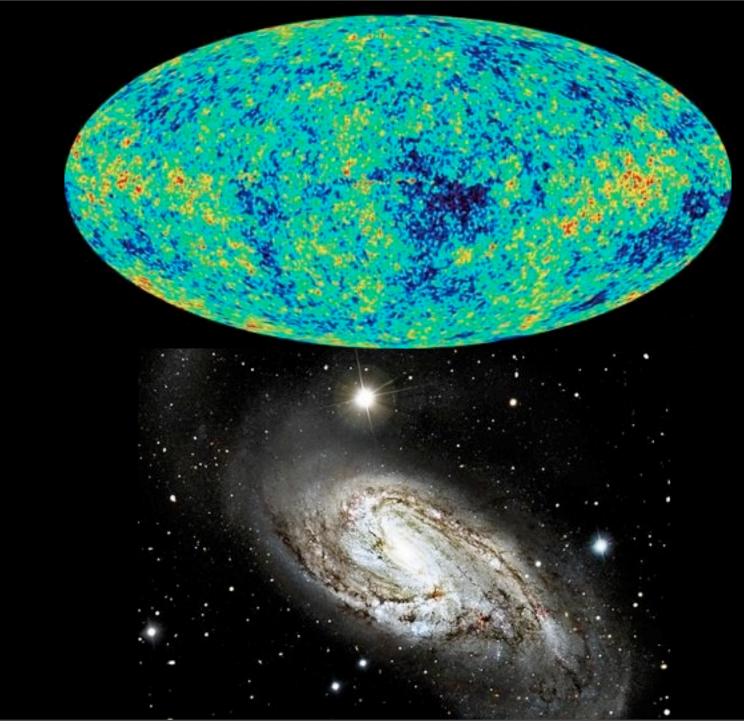


"Our aesthetic judgment therefore comes down to what we find more wasteful: many worlds or many words. Perhaps we will gradually get more used to the weird ways of our cosmos, and even find its strangeness to be part of its charm."

- Max Tegmark







Wednesday, November 10, 2010





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