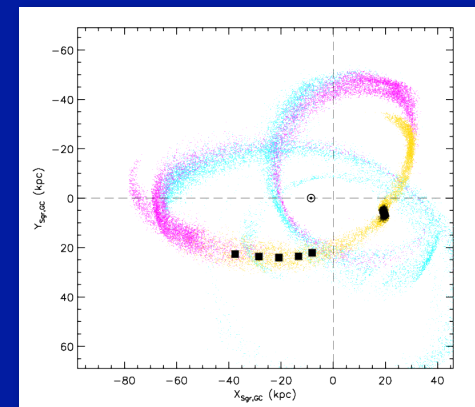
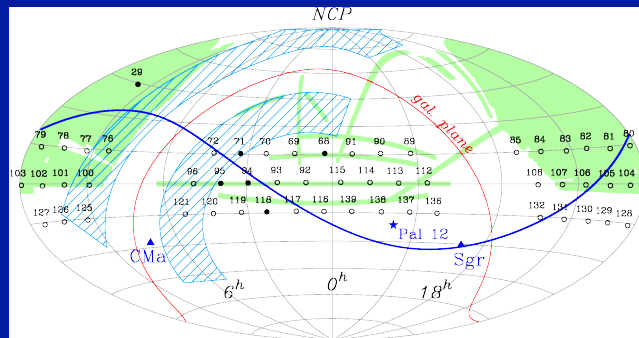


Characterizing Halo Substructure with Kapteyn Proper-motion Survey and SEGUE/SDSS Data

Dana Casetti - Yale University



The Team and Collaborations

- *Yale*: Dana Casetti, Terry Girard
- *Rensselaer P I*: Jeff Carlin
- *University of Virginia*: Steve Majweski

- SDSS
- Mike Siegel, David Law, Kathy Vivas, Ron Wilhelm, Tim Beers, Bill van Altena, Ricky Patterson, Carl Grillmair, Jorge Penarrubia

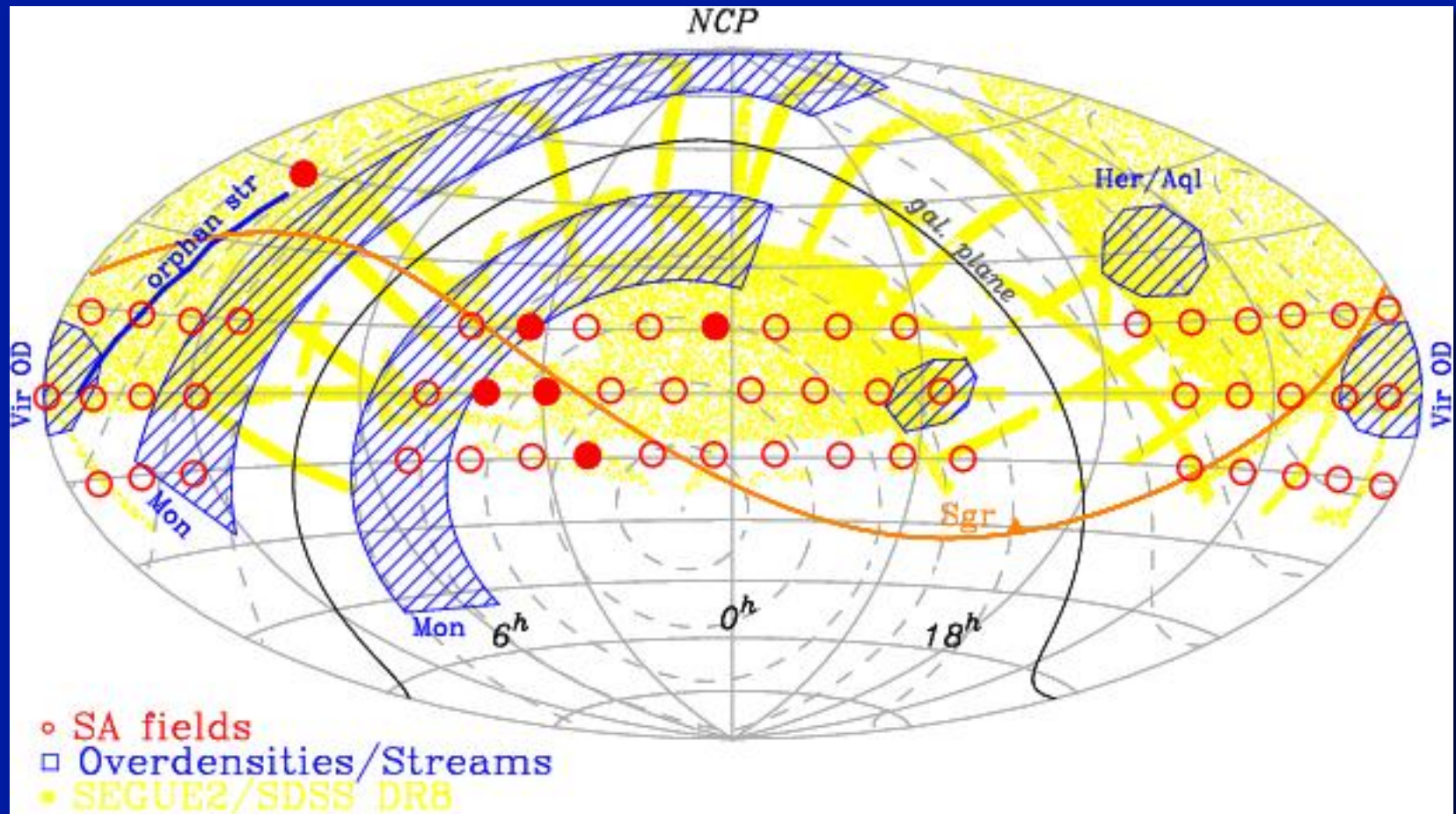
Outline

- Survey Description
- Comparison with current proper-motion surveys (Munn et al. 2008-DR7, Bramich et al. 2008)
- Science highlights

Survey Layout and Status

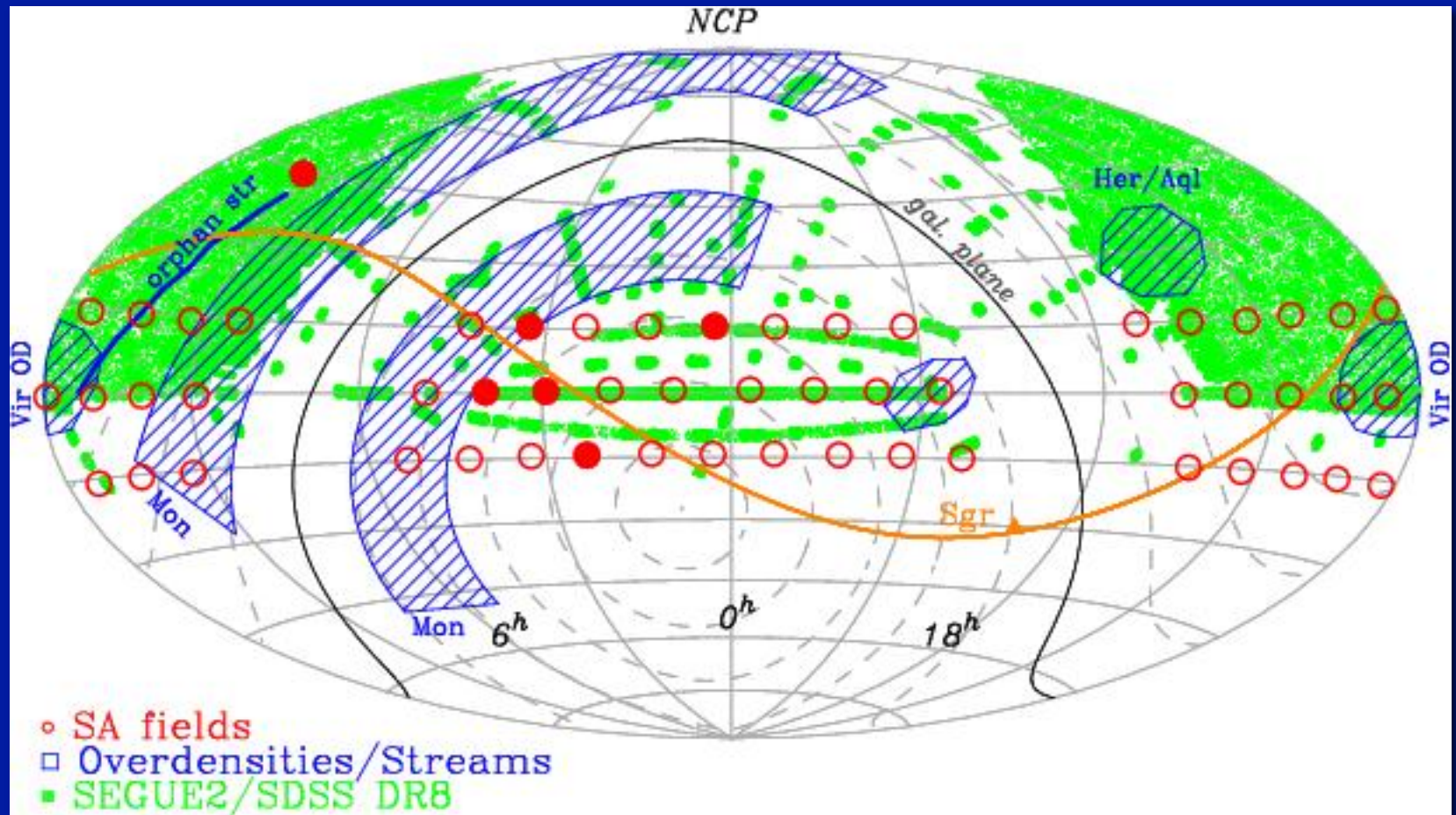
- 51 Kapteyn Selected Area fields of 40' x40' (originally 55, but plates are not available for 4 fields)
- 3 Dec. zones: 0° , $+15^\circ$, -15° ; every $\sim 15^\circ$ in RA
- 51 fields measured, 28 reduced
- 35 (of 51) overlap with DR8; 22 (of 28 reduced) overlap with DR8
- 7 are in stripe 82, and overlap with Bramich et al. 2008

Survey Layout and DR8



DR8 Imaging

Survey Layout and DR8



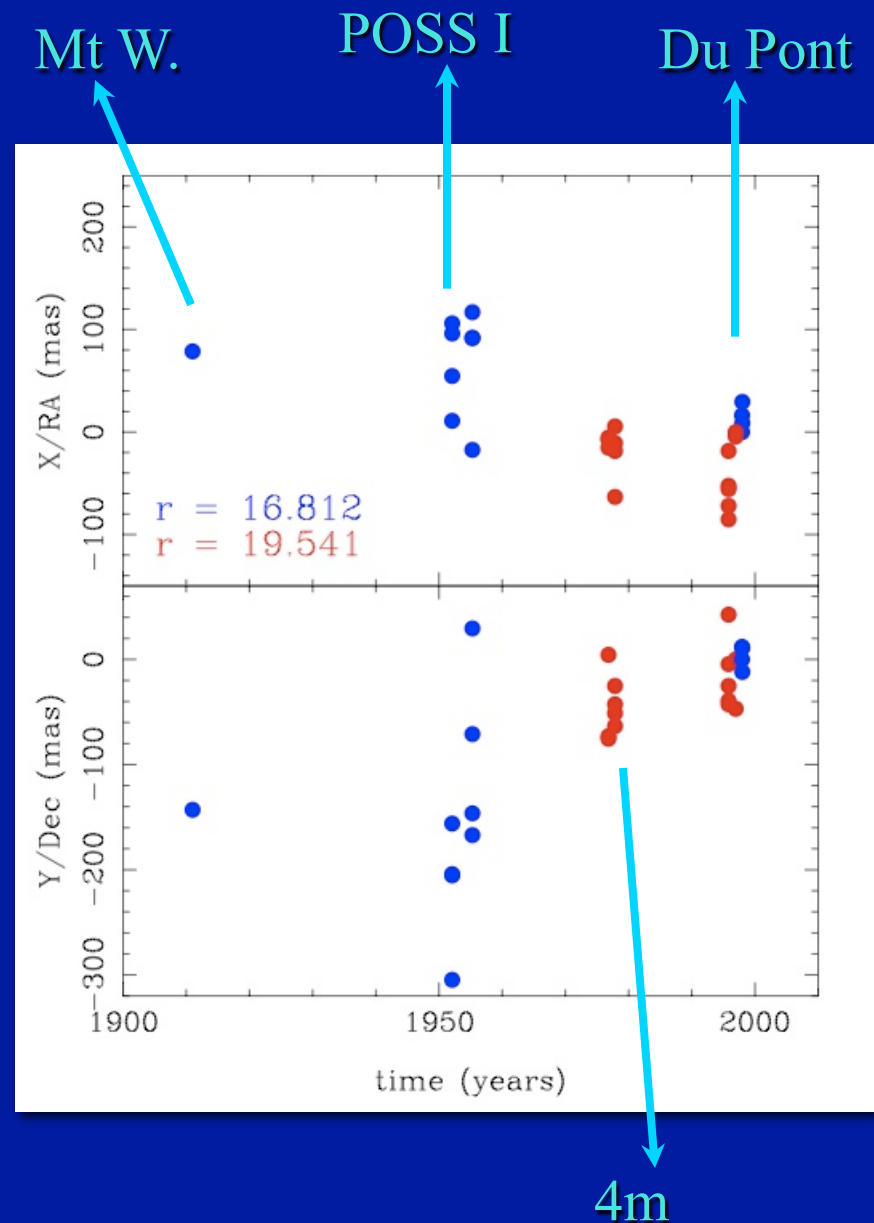
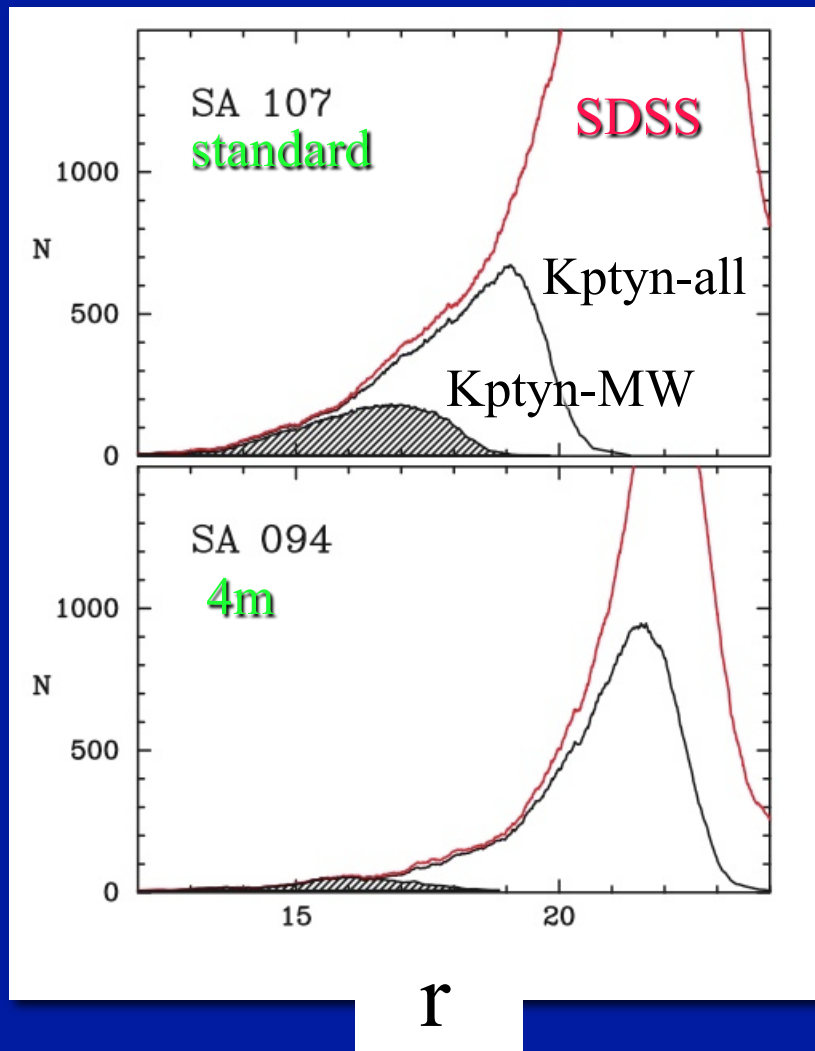
Proper Motions

- LCO du Pont 2.5m, 1996-2000, (Majewski), IIIaJ-F; $V_{\text{lim}} = 21$ to 22; 10.9"/mm
- KPNO Mayall 4m PF: mid 1970; (Sandage) and mid 1990s (Majewski); IIaO,D and IIIaJ,F; $V_{\text{lim}} \sim 23$; 18.6"/mm
- Mt. Wilson 60-inch, 1909 - 1912, to $V_{\text{lim}} = \sim 18$; 27.1"/mm
- POSS-I Oschin Schmidt 1.2m, early-mid 1950s; scanned at USNO and STScI (scans processed at Yale); $V_{\text{lim}} \sim 21.5$; 67.2"/mm

-Errors: 1-2 mas/yr per well-measured star ($V \sim 17$ and $V \sim 19$ for 4m); zero point ~ 0.3 to 0.6 mas/yr

-Status: 28 fields reduced with $\sim 57,000$ objects

Proper Motions



Radial Velocities

-WIYN 3.5m (Hydra): ~1600 spectra; errors ~ 5 -10 km/s

$\lambda = 5200\text{\AA}$ (4400-7200 $\text{\AA}$, $R \sim 1500$) and $\lambda = 5170\text{\AA}$ ($R \sim 5400$)

-CTIO 4m (Hydra): ~500 spectra; errors 10-20 km/s

$\lambda = 5500\text{\AA}$ (4600-7000 $\text{\AA}$, $R \sim 1600$)

-MMT 6.5m (Hectospec): ~1000 spectra; errors ~10-15 km/s

$\lambda = 6400\text{\AA}$ (3700-9150 $\text{\AA}$, $R \sim 1300$)

-SDSS / SEGUE

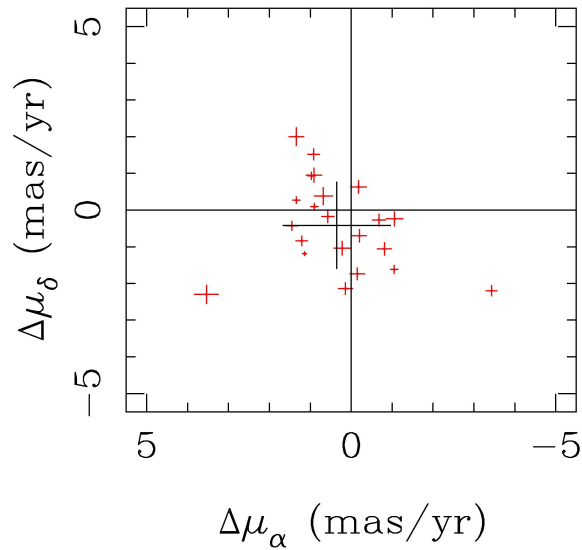
Photometry/Object classification

-SDSS

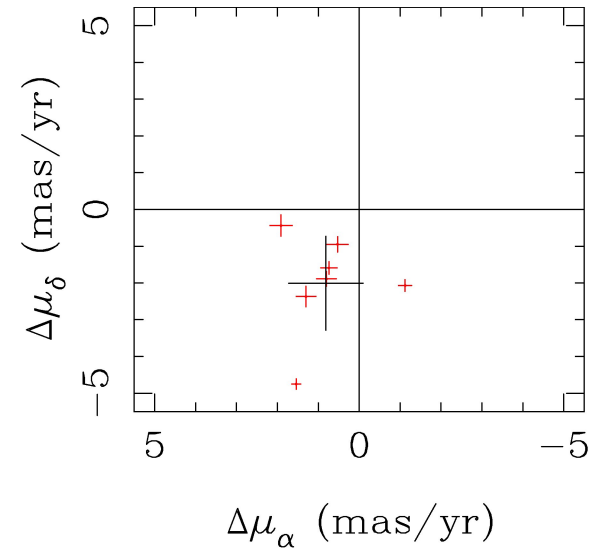
-Other: LCO Swope 1m, KPNO 0.9m, literature

Proper-motion Comparisons: Differences

dr7-K : 22 fields

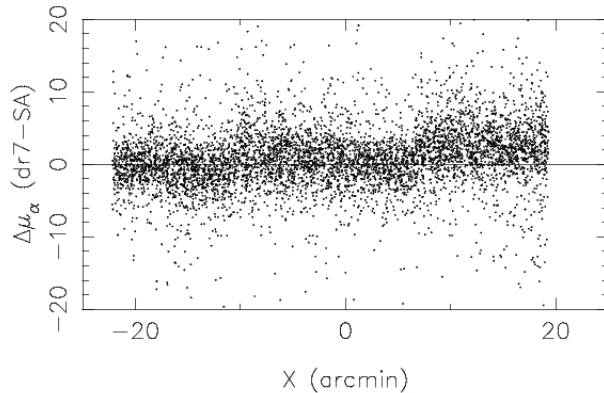


Bramich-K : 7 fields



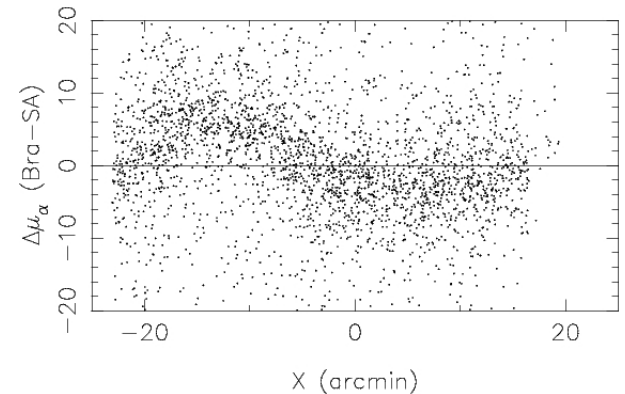
dr7-K : per field

sa112

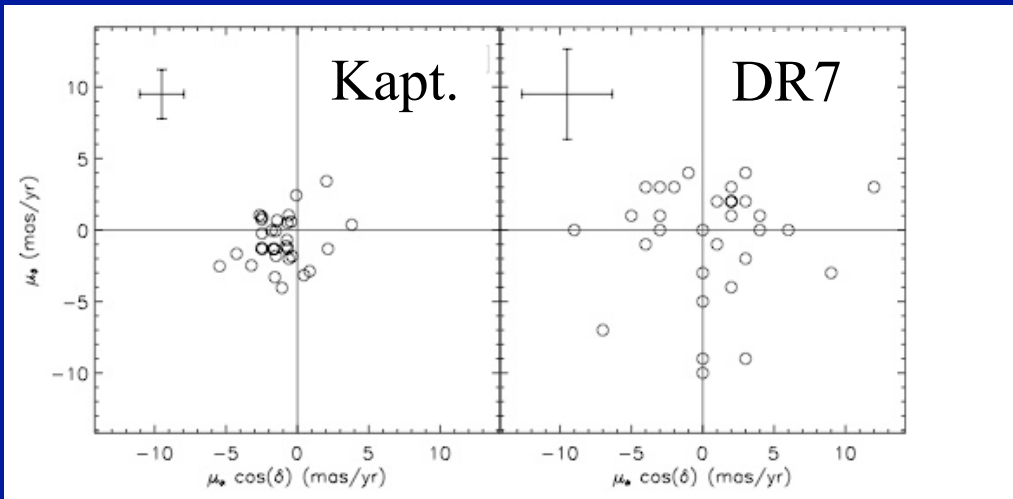


Bramich-K : per field

sq094



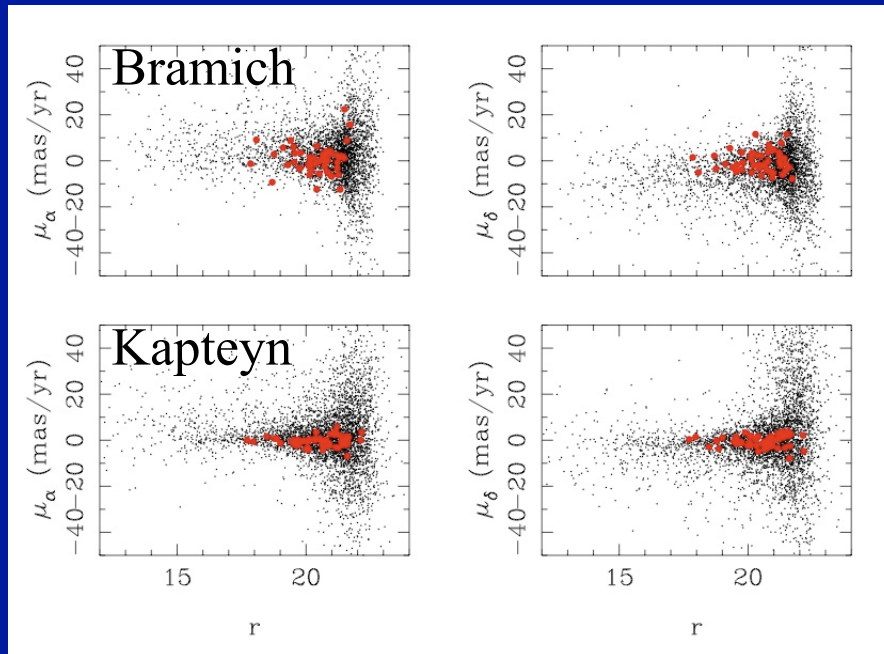
Proper-motion Comparisons: Actual values



SA 76 - comparison with dr7

Stars in the Mon/Anticenter stream (RV-CMD selected)

Carlin et al. 2010



SA 94 – comparison with
Bramich2008, stripe 82

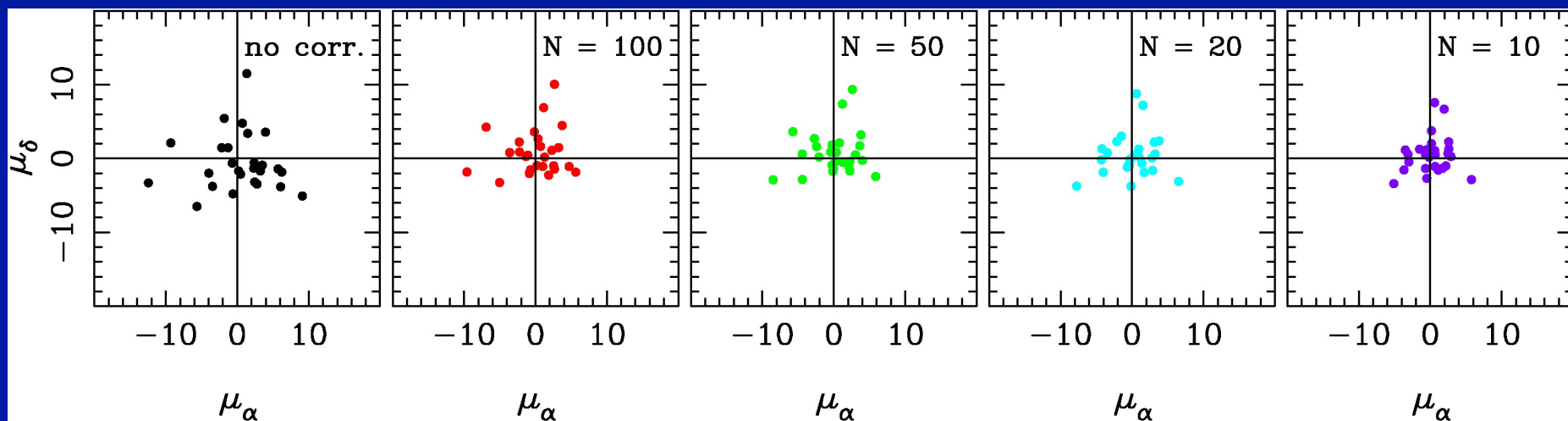
Stars

QSOs

Proper-motion Corrections to Bramich 2008

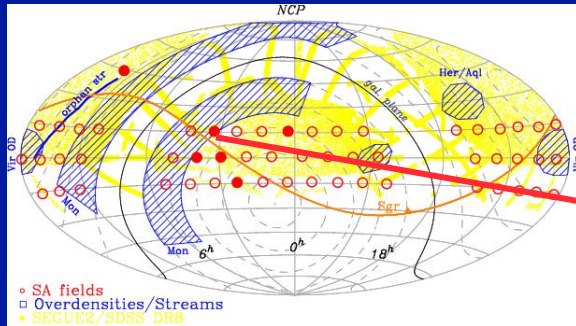
SA 94

Proper-motion distribution of QSOs (mas/yr)

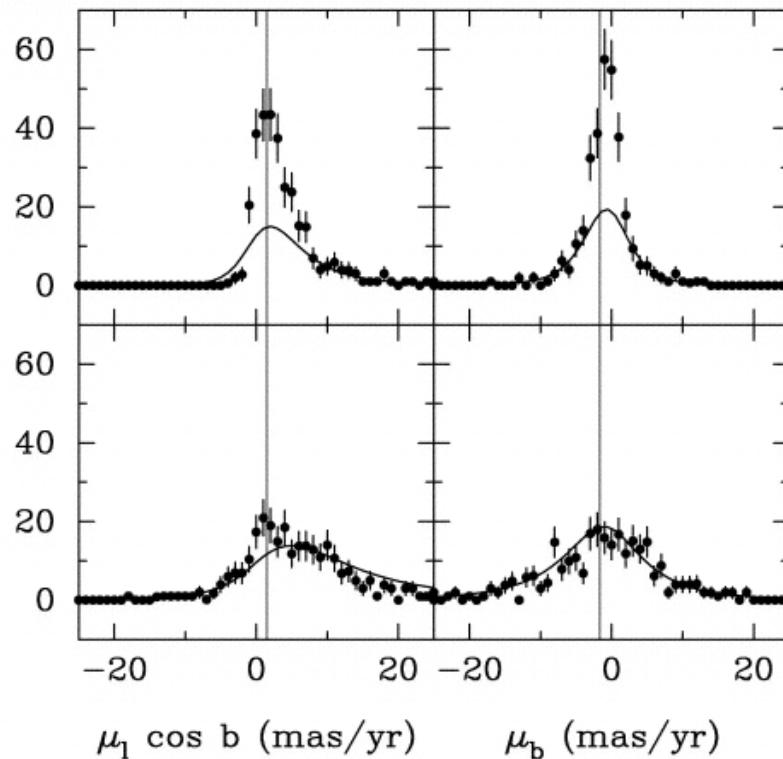
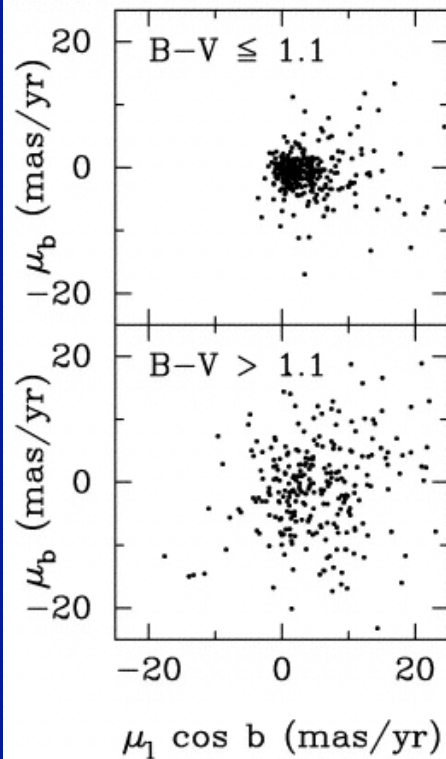


sdev μ_α	4.8	3.5	3.3	3.1	2.4
sdev μ_δ	3.9	3.1	2.9	2.9	2.6

Finding Streams in Proper-motion Space

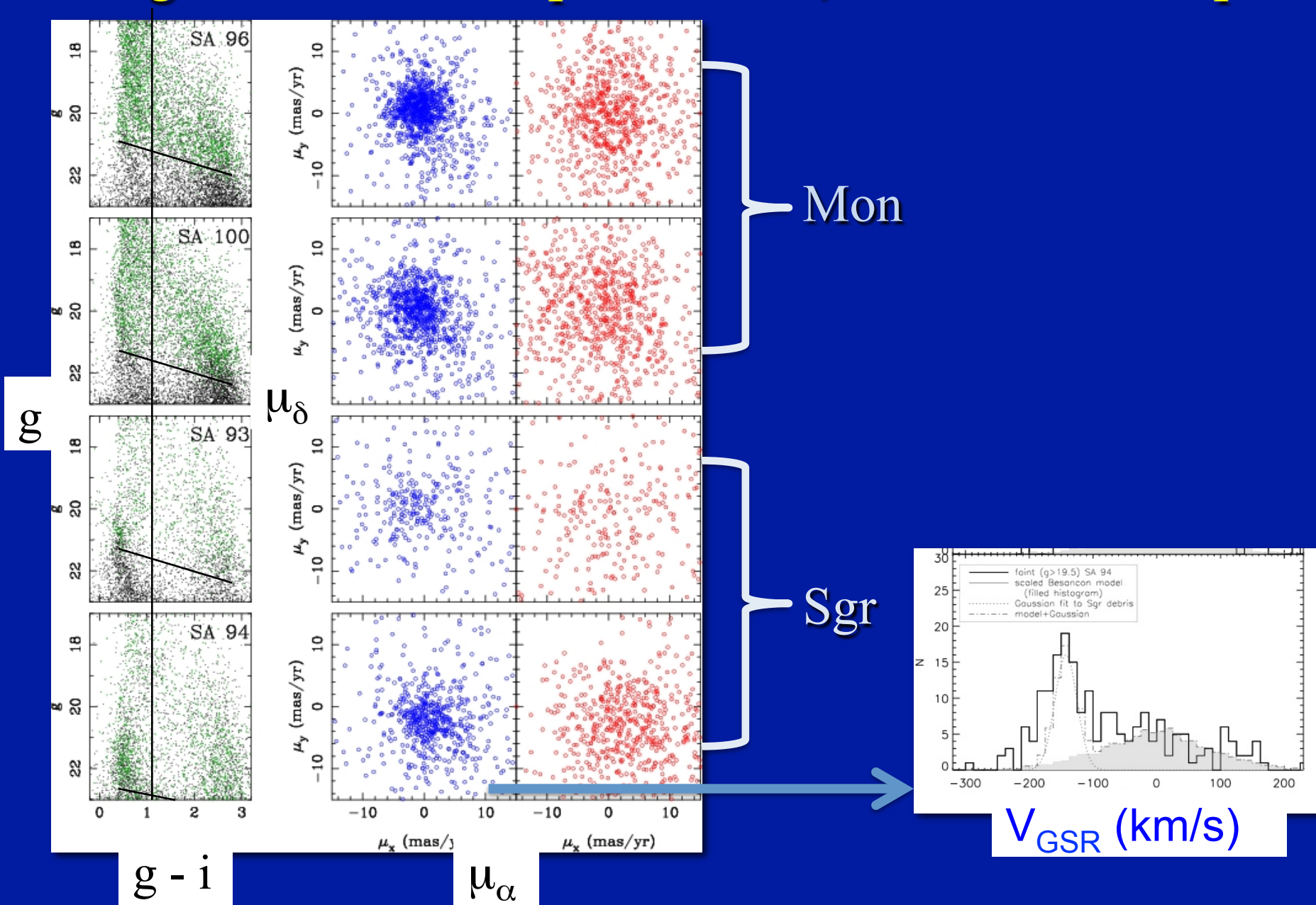


SA 71



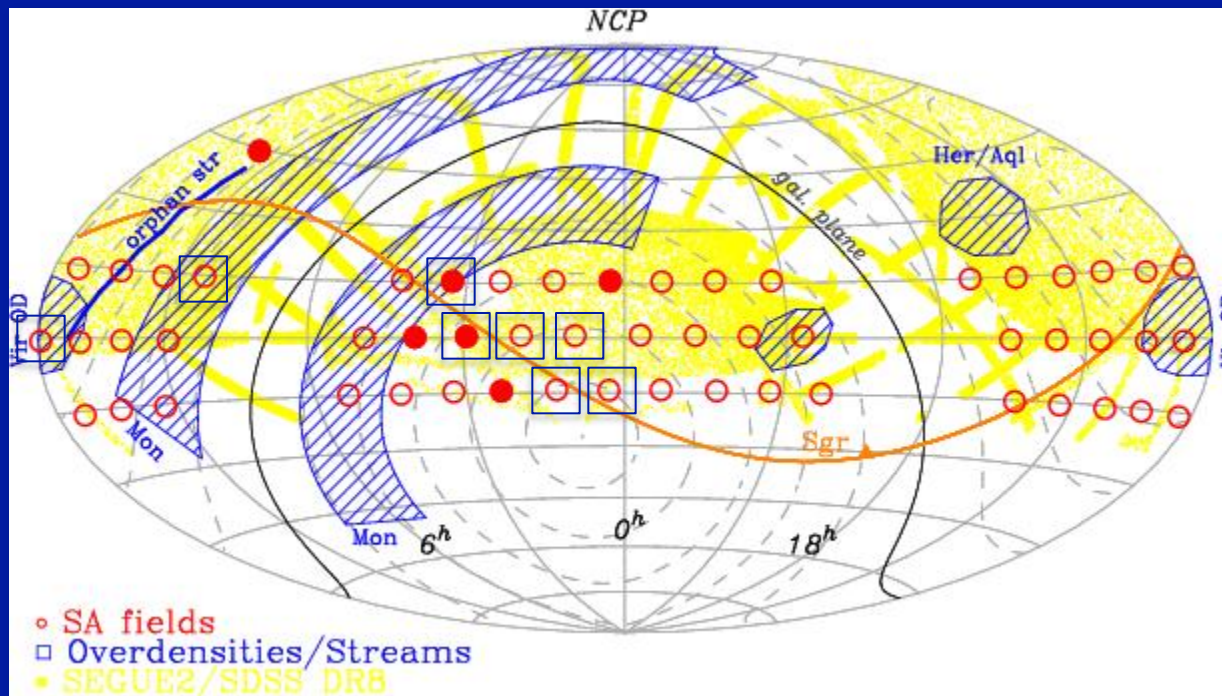
-clumpiness
-stellar excess

Finding Streams in Proper-motion, CMD and RV Space



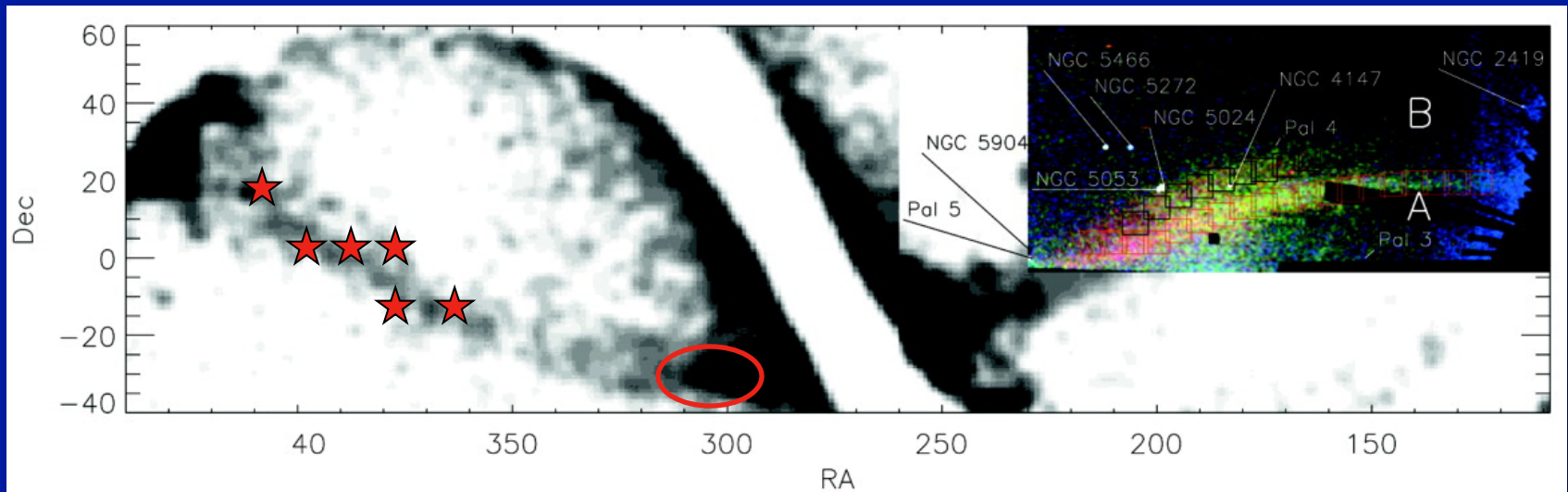
Science Highlights: Results

- The Monoceros/Anticenter structure (SA 71, 76)
- The Virgo Stellar Stream (SA 103) - poster by Yam et al.
- The Sagittarius trailing tail (SA 71, 93, 94, 117, 92, 116)



Science Highlights: Characterizing Streams

Sagittarius trailing tidal tail and constraining Θ_{LSR}

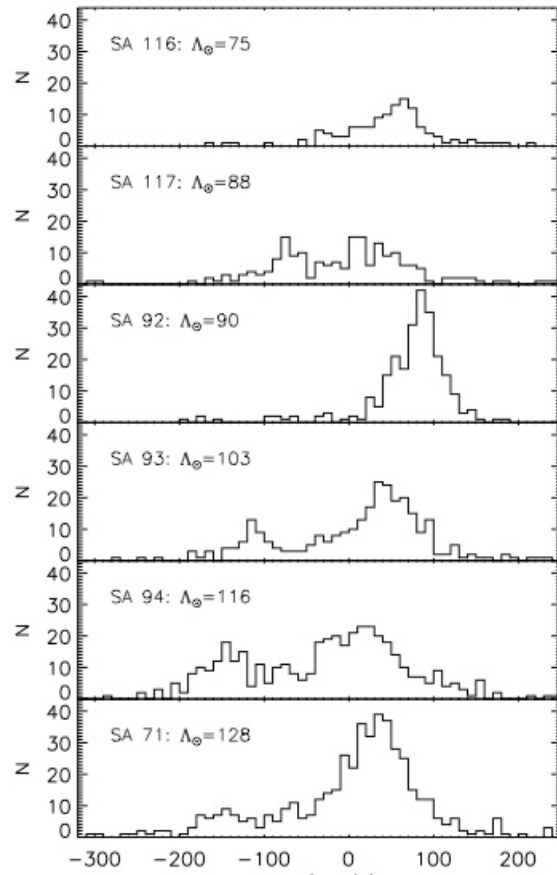


SDSS and 2MASS:
Belokurov et al. 2006,
2007, 2008, Majewski et
al. 2003

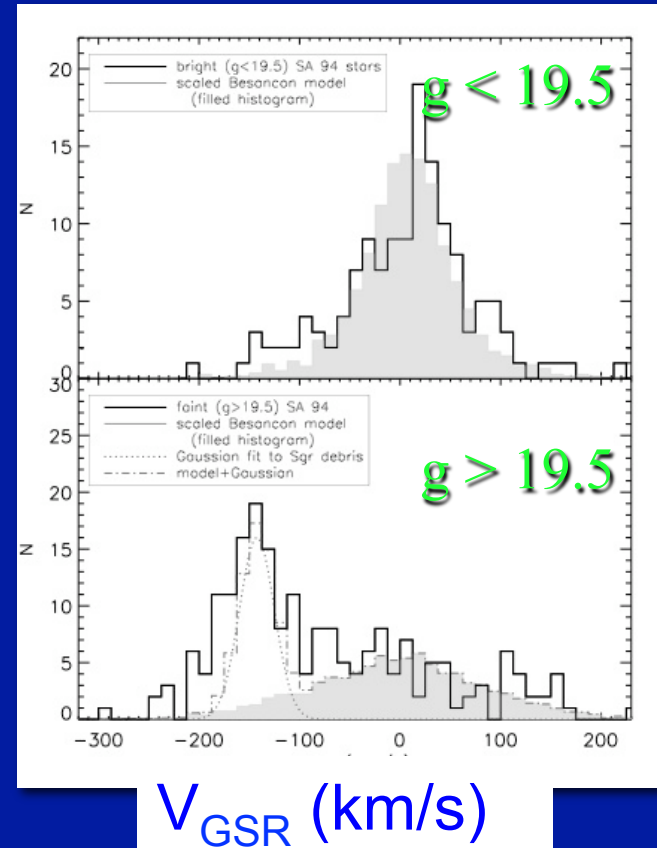
Carlin et al. 2012

Sagittarius trailing tidal tail and constraining Θ_{LSR}

Λ increasing

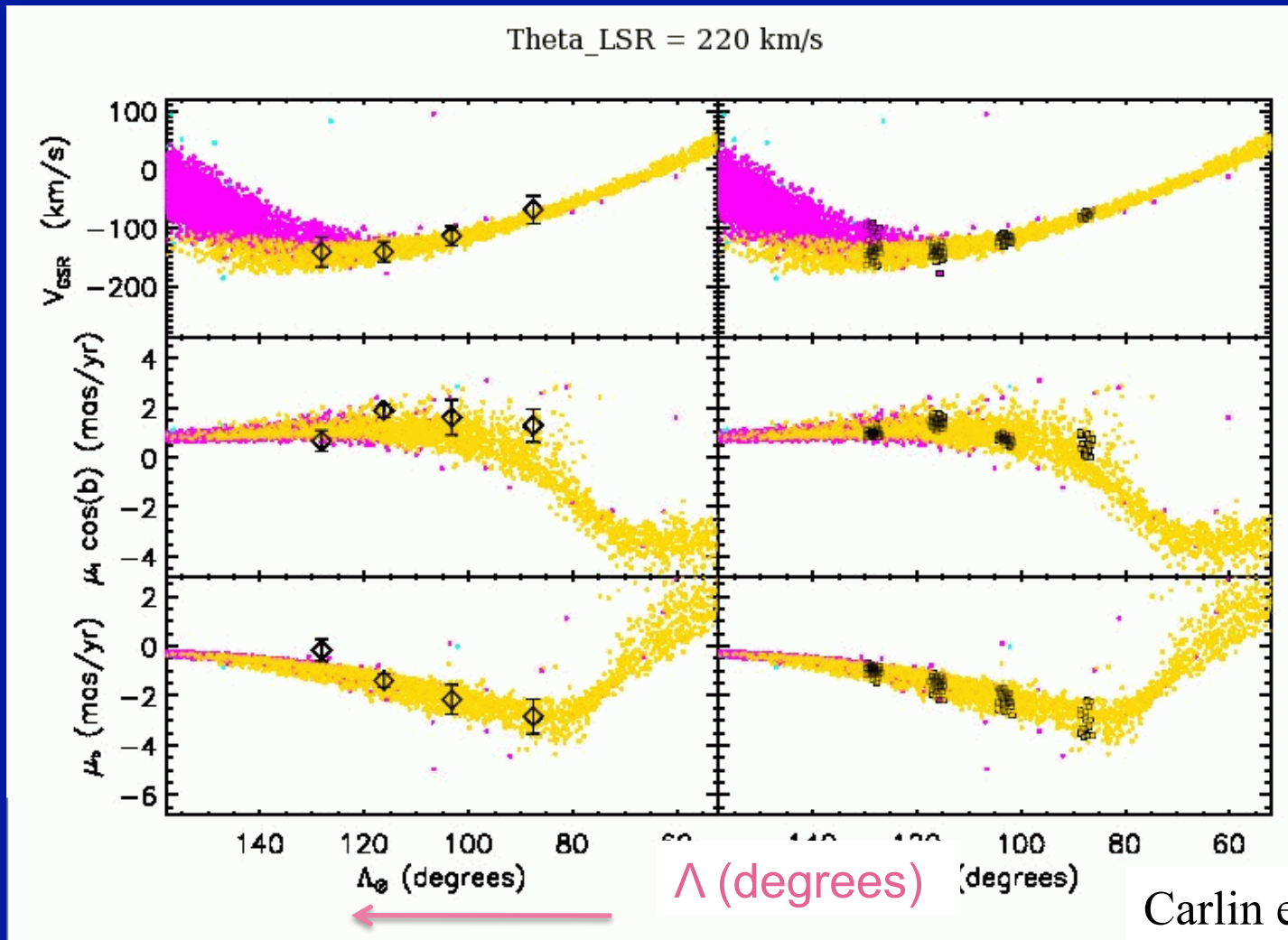


V_{GSR} (km/s)



V_{GSR} (km/s)

Sagittarius trailing tidal tail and constraining Θ_{LSR}

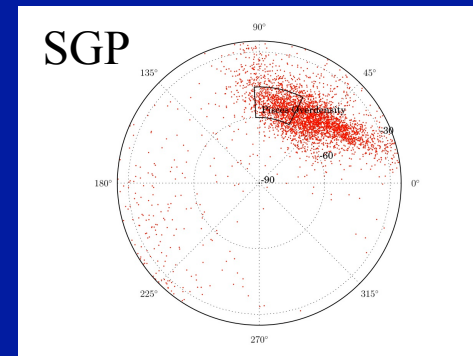
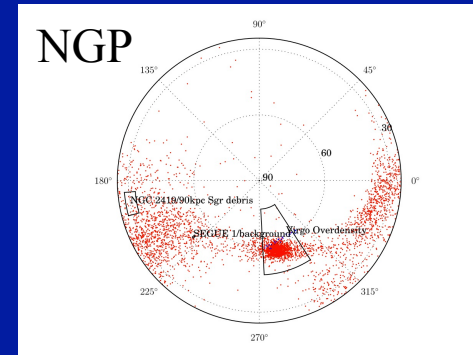
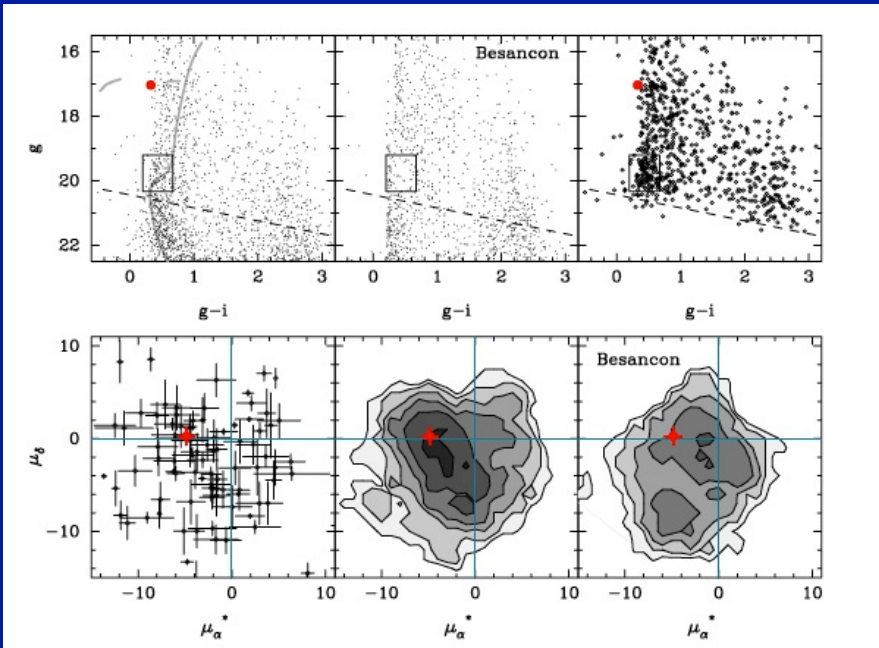


Carlin et al. 2012

3D vel.fit $\Rightarrow \Theta_{\text{LSR}} = 232 \pm 14 \text{ km/s}$

Science Highlights: Characterizing Streams

SA 103: Virgo Stellar Stream



Casetti-Dinescu et al. 2009

1 RR Lyr $\Rightarrow$ $\text{ecc} \sim 0.8$;

$R_p \sim 11 \text{ kpc}$; $R_a \sim 90 \text{ kpc}$

Pericentric debris of a former satell.

Assoc with NGC 2419

Yam et al. 2012 – poster 252.16

16 stars $\Rightarrow$ similar orbit as CD09

Assoc with NGC 2419, Segue1b/

Segue, Pisces

Summary

- To date, this survey is best suited for the study of 3D kinematics of streams
- More fields to be reduced/analyzed
- Potential for calibrating other surveys
- Du Pont plates to be used in combination with modern ccd data for deeper/more precise catalogs
- Sgr trailing tail=> constraining MW potential
- VSS/VOD – its nature/formation